

# Class Progress

Basics of Linux, gnuplot, C

Visualization of numerical data

Roots of nonlinear equations

(Midterm 1)

Solutions of systems of linear equations

Solutions of systems of nonlinear equations

Monte Carlo simulation

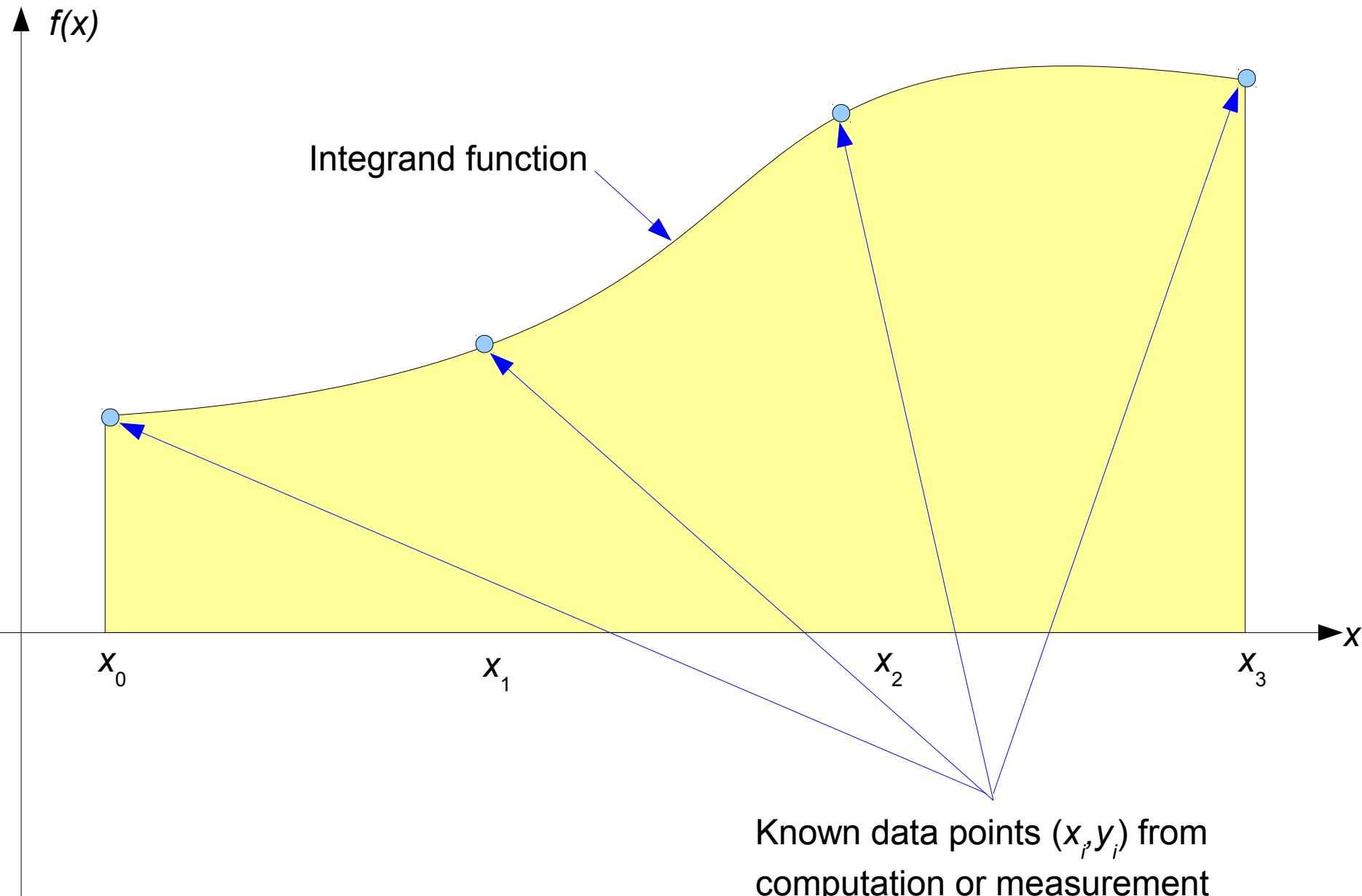
Interpolation of sparse data points

**Numerical integration**

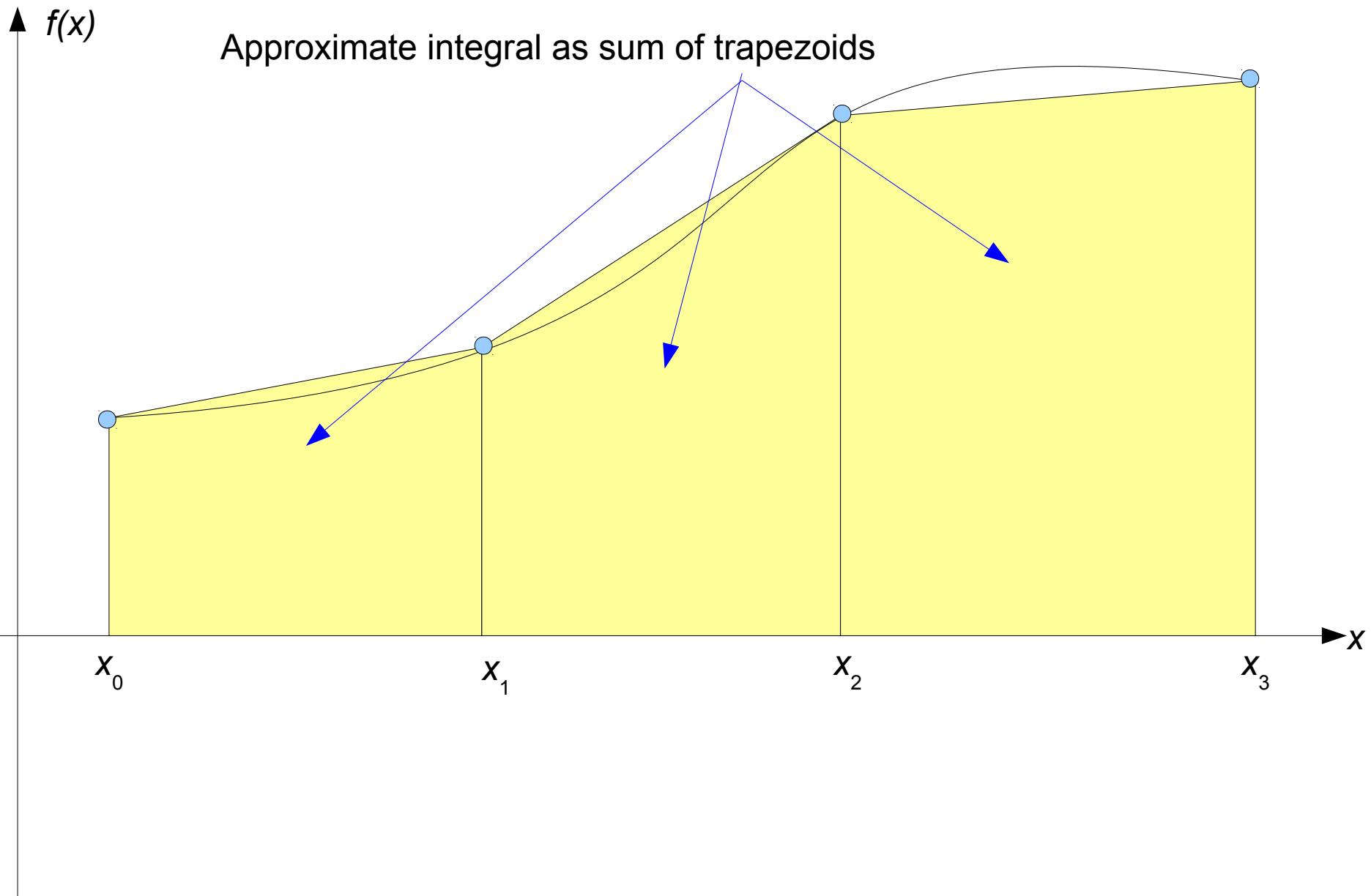
(Midterm 2)

Solutions of ordinary differential equations

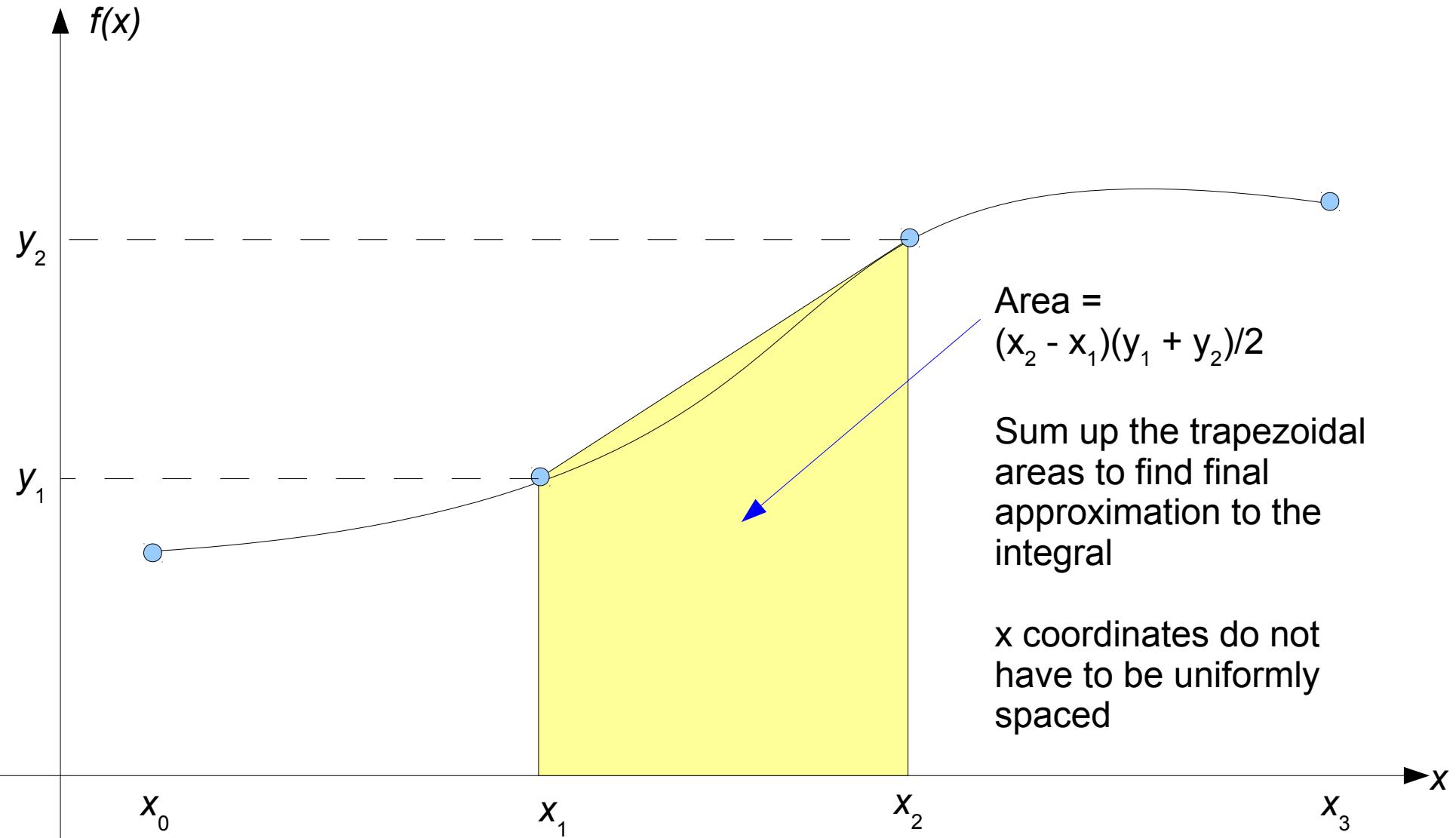
# Integration of a Function from Known Data Points



# Integration of a Function with Trapezoidal Rule



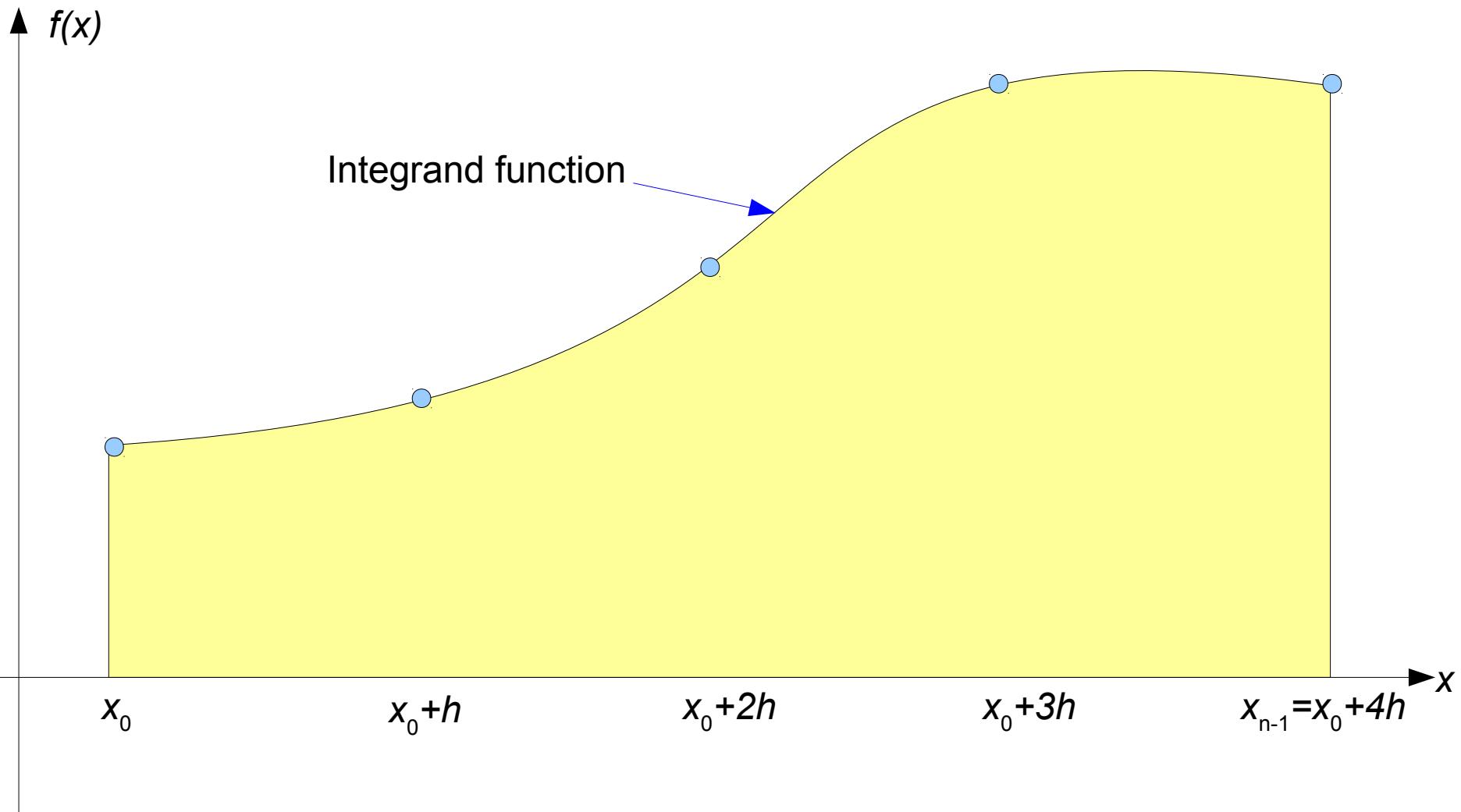
# Sum the Area of Each Trapezoid



# Integration C Code for Tabulated Data Points

```
int trapezoidal_data(int n, struct point2d *sample) {  
    double integral;  
    int i;  
  
    integral = 0.0;  
    i = 1;  
    while (i < n) {  
        integral += 0.5 * (sample[i].y + sample[i-1].y) * (sample[i].x -  
sample[i-1].x);  
        i++;  
    }  
    printf("%.12g\n", integral);  
    return(0);  
}
```

# Trapezoidal Rule with Uniformly Spaced $x$ Values



# Integration of a Function with Trapezoidal Rule

Each trapezoidal area now =  $\frac{h}{2}(y_i + y_{i+1})$

So sum up the areas of all the trapezoids:

$$\frac{h}{2}y_0 + \frac{h}{2}y_1$$

$$\frac{h}{2}y_1 + \frac{h}{2}y_2$$

$$\frac{h}{2}y_2 + \frac{h}{2}y_3$$

.....

+

$$\frac{h}{2}y_{n-2} + \frac{h}{2}y_{n-1}$$

---

$$\frac{h}{2}y_0 + h y_1 + h y_2 + \cdots + h y_{n-2} + \frac{h}{2}y_{n-1}$$

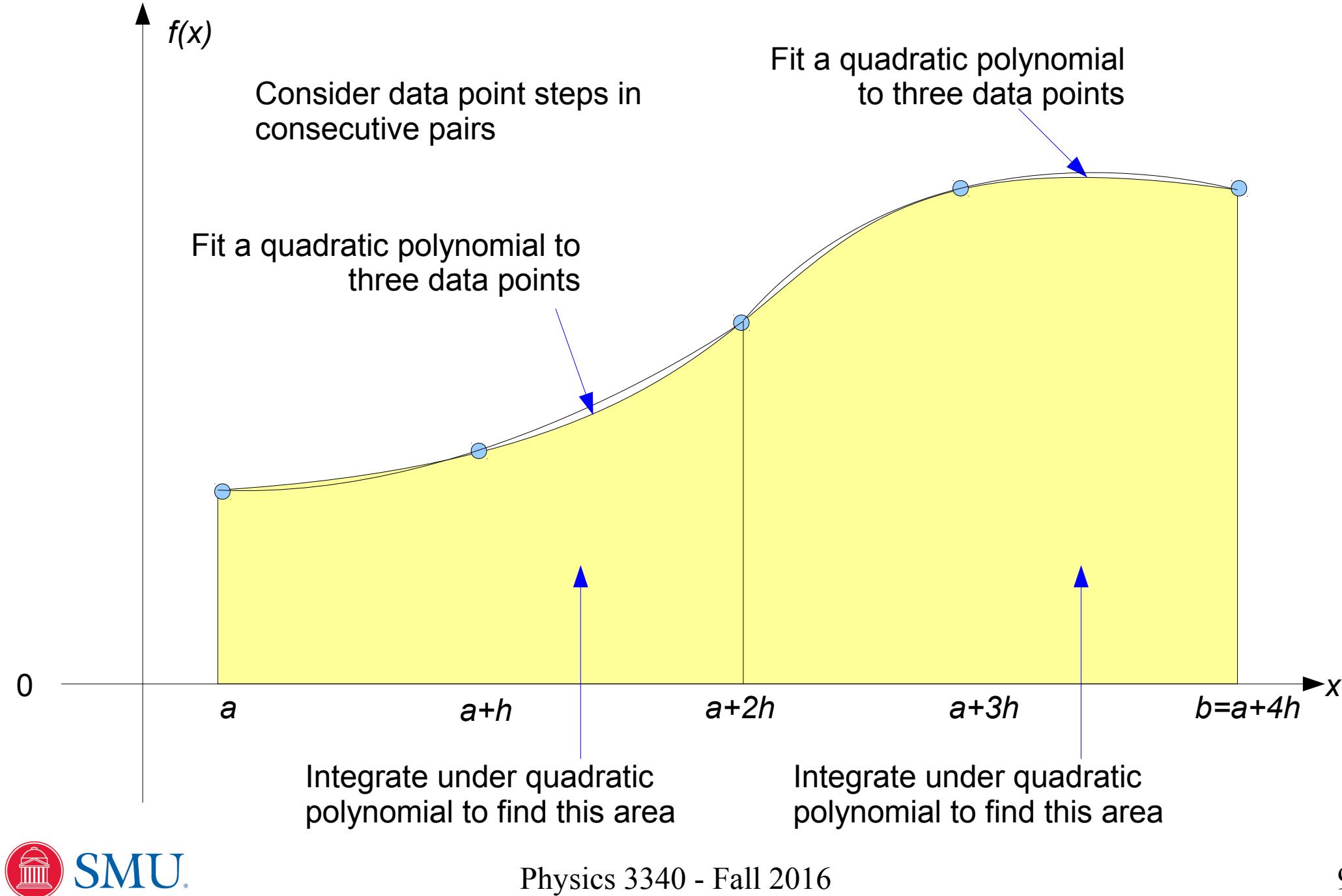
# Integration of a Function with Trapezoidal Rule

For  $n$  uniformly spaced data points:

$$\int_{x_0}^{x_{n-1}} f(x) dx \approx h \cdot \left[ \frac{(y_0 + y_{n-1})}{2} + \sum_{i=1}^{n-2} y_i \right]$$

$$\text{Error term} = -\left(\frac{1}{12}\right)(x_{n-1} - x_0)h^2 f''(\xi)$$

# Simpson's Rule with Uniformly Spaced $x$ Values



# Integration of a Function with Simpson's 1/3 Rule

Fit second order Lagrange interpolation polynomial to first three data points:

$$f(x) = \frac{(x-x_1)(x-x_2)}{(x_0-x_1)(x_0-x_2)} \cdot y_0 + \frac{(x-x_0)(x-x_2)}{(x_1-x_0)(x_1-x_2)} \cdot y_1 + \frac{(x-x_0)(x-x_1)}{(x_2-x_0)(x_2-x_1)} \cdot y_2$$

Change  $x$  axis so that  $x_0=0$   $x_1=h$   $x_2=2h$  and rewrite Lagrange interpolation polynomial for uniformly spaced data points:

$$f(x) = \frac{(x-h)(x-2h)}{2h^2} \cdot y_0 - \frac{x(x-2h)}{h^2} \cdot y_1 + \frac{x(x-h)}{2h^2} \cdot y_2$$

Integrate under quadratic polynomial from 0 to  $2h$ :

$$\int_0^{2h} f(x) dx \approx \left(\frac{h}{3}\right) \cdot [y_0 + 4y_1 + y_2]$$

# Integration of a Function with Simpson's 1/3 Rule

So sum up the areas under all the quadratic polynomials:

$$\frac{h}{3} y_0 + \frac{4h}{3} y_1 + \frac{h}{3} y_2$$

$$\frac{h}{3} y_2 + \frac{4h}{3} y_3 + \frac{h}{3} y_4$$

$$\frac{h}{3} y_4 + \frac{4h}{3} y_5 + \frac{h}{3} y_6$$

.....

+

$$\frac{h}{3} y_{n-3} + \frac{4h}{3} y_{n-2} + \frac{h}{3} y_{n-1}$$

$$\frac{h}{3} y_0 + \frac{4h}{3} y_1 + \frac{2h}{3} y_2 + \frac{4h}{3} y_3 + \frac{2h}{3} y_4 + \cdots + \frac{2h}{3} y_{n-3} + \frac{4h}{3} y_{n-2} + \frac{h}{3} y_{n-1}$$

# Integration of a Function with Simpson's 1/3 Rule

So for  $n$  uniformly spaced data points:

$$\int_{x_0}^{x_{n-1}} f(x) dx \approx \left(\frac{h}{3}\right) \cdot \left[ (y_0 + y_{n-1}) + 4 \cdot \sum_{i=1, i \text{ odd}}^{n-2} y_i + 2 \cdot \sum_{i=2, i \text{ even}}^{n-3} y_i \right]$$

Applicable for only an even number of steps, so  $n-1$  must be an even number, or the number of function points must be odd

$$\text{Error term} = -\left(\frac{1}{180}\right)(x_{n-1} - x_0) h^4 f^{(4)}(\xi)$$

# Integration C Code for C Function

```
#include <stdio.h>
#include <stdlib.h>
#include "integrate.h"

int trapezoidal(double (*func) (double x), double xstart, double
xstop, double xinc) {
    double x, integral;

    x = xstart;
    integral = 0.5 * xinc * (*func) (x);
    x = x + xinc;
    xstop = xstop - (xinc * 0.5);
    while (((xinc > 0.0) && (x < xstop)) || ((xinc < 0.0) && (x >
xstop))) {
        integral += xinc * (*func) (x);
        x = x + xinc;
    }
    integral += 0.5 * xinc * (*func) (x);
    printf("%.8g\n", integral);
    return(0);
}
```

Note '-' sign to stop loop  
before the last desired point!

# Integration C Code for C Function

```
int simpson(double (*func) (double x), double xstart, double xstop, double xinc) {  
    double x, integral;  
    int state;  
  
    x = xstart;  
    integral = xinc * (*func) (x) / 3.0;  
    x = x + xinc;  
    xstop = xstop - (xinc * 0.5);  
    state = 0;  
    while (((xinc > 0.0) && (x < xstop)) || ((xinc < 0.0) && (x > xstop))) {  
        if (state == 0) {  
            integral += 4.0 * xinc * (*func) (x) / 3.0;  
            state = 1;  
        }  
        else {  
            integral += 2.0 * xinc * (*func) (x) / 3.0;  
            state = 0;  
        }  
        x = x + xinc;  
    }  
    if (state != 1) {  
        fprintf(stderr, "Warning: Odd number of integration segments\n");  
    }  
    integral += xinc * (*func) (x) / 3.0;  
    printf("%.8g\n", integral);  
    return(0);  
}
```



# Integration C Code

```
#include <stdio.h>
#include <stdlib.h>
#include <math.h>
#include "integrate.h"
#include "constants.h"

double mu,sigma;

double gaussian(double x) {
    return(exp(-(x - mu) * (x - mu) / (2.0 * sigma * sigma)) / (sqrt(2.0 * PI) * sigma));
}

int main(int argc,char *argv[]) {
    double xstart,xstop,xinc;

    if (argc != 6) {
        fprintf(stderr,"%s <mu> <sigma> <xstart> <xstop> <xinc>\n",argv[0]);
        exit(1);
    }
    mu = atof(argv[1]);
    sigma = atof(argv[2]);
    xstart = atof(argv[3]);
    xstop = atof(argv[4]);
    xinc = atof(argv[5]);
    printf("Trapezoidal rule:\n");
    trapezoidal(gaussian,xstart,xstop,xinc);
    printf("Simpson's rule:\n");
    simpson(gaussian,xstart,xstop,xinc);
    exit(0);
}
```

# Rotational Inertia

For a given object, define

$$I = \int r^2 dm$$

As the rotational inertia of an object about a particular axis of rotation

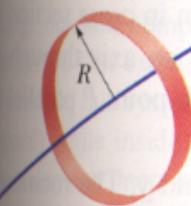
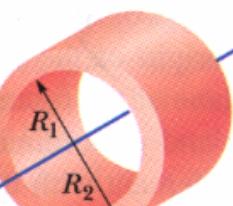
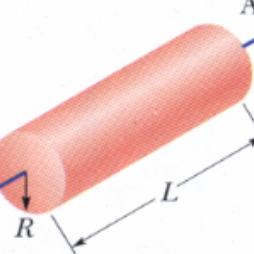
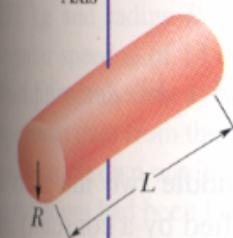
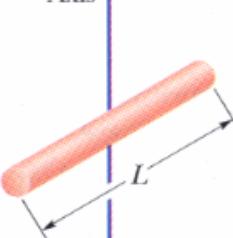
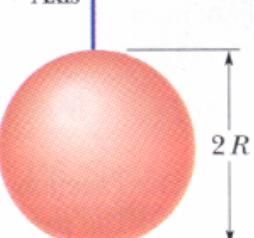
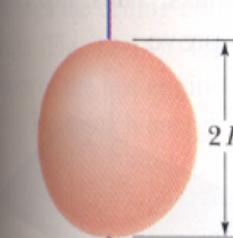
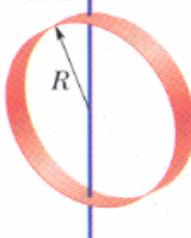
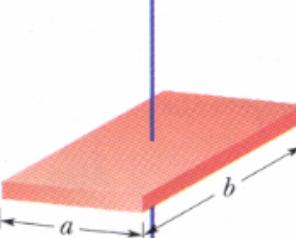
Then simple rotational equivalent of the Newtonian translational relationships can be derived, such as

$$\text{Torque } \tau = I \alpha$$

$$\text{Kinetic energy } K = \frac{1}{2} I \omega^2$$

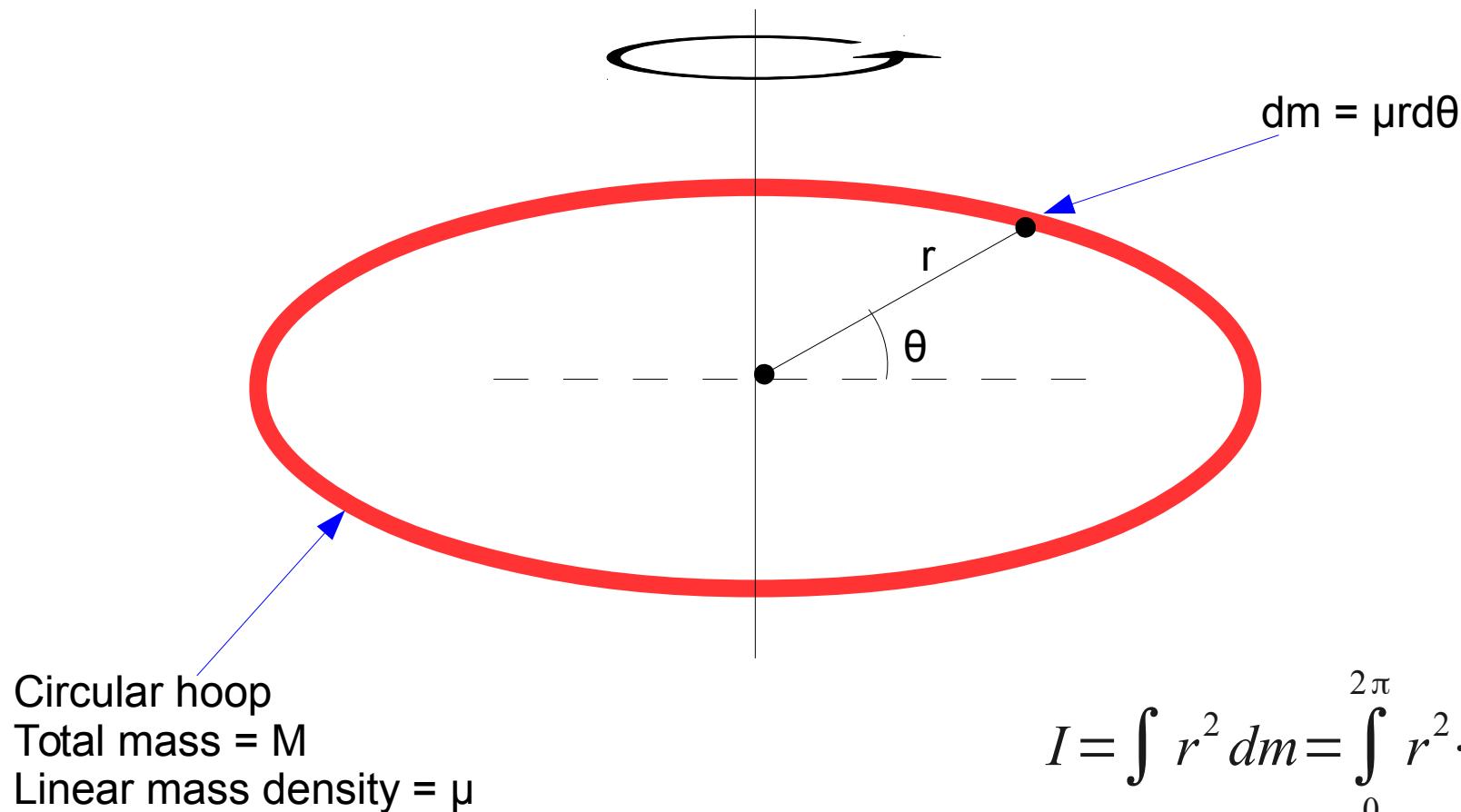
# Sample Rotational Inertia Integrals

TABLE 11-2 Some Rotational Inertias

 <p>Hoop about central axis</p> $I = MR^2$	 <p>Annular cylinder (or ring) about central axis</p> $I = \frac{1}{2}M(R_1^2 + R_2^2)$	 <p>Solid cylinder (or disk) about central axis</p> $I = \frac{1}{2}MR^2$
 <p>Solid cylinder (or disk) about central diameter</p> $I = \frac{1}{4}MR^2 + \frac{1}{12}ML^2$	 <p>Thin rod about axis through center perpendicular to length</p> $I = \frac{1}{12}ML^2$	 <p>Solid sphere about any diameter</p> $I = \frac{2}{5}MR^2$
 <p>Thin spherical shell about any diameter</p> $I = \frac{2}{3}MR^2$	 <p>Hoop about any diameter</p> $I = \frac{1}{2}MR^2$	 <p>Slab about perpendicular axis through center</p> $I = \frac{1}{12}M(a^2 + b^2)$



# Simple Rotational Inertial Problem



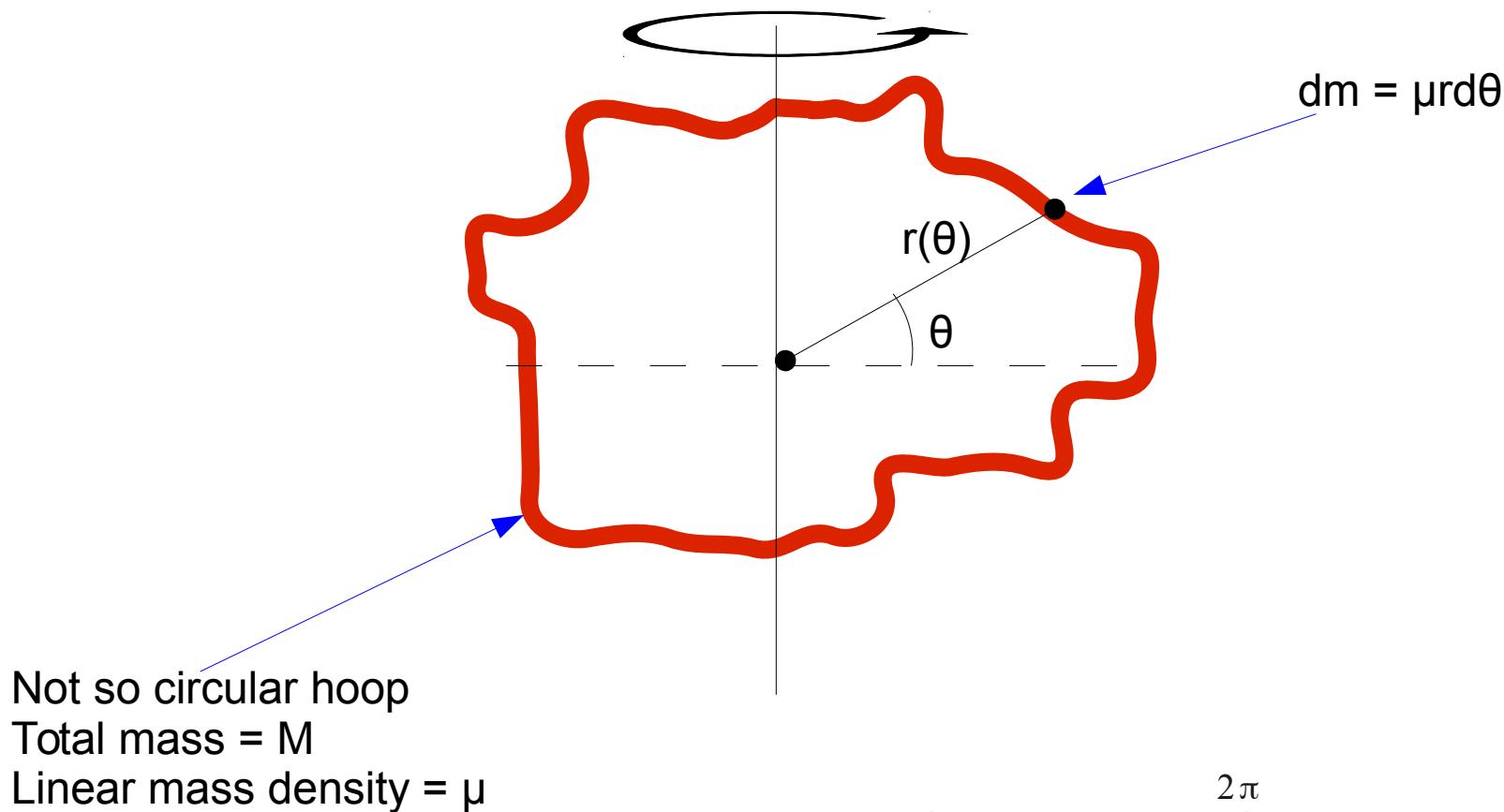
$$I = \int r^2 dm = \int_0^{2\pi} r^2 \cdot (\mu r d\theta)$$

Let  $r=R$ , a constant

$$I = R^3 \mu \int_0^{2\pi} d\theta = 2\pi R^3 \mu$$

$$= (2\pi R \mu) \cdot R^2 = M R^2$$

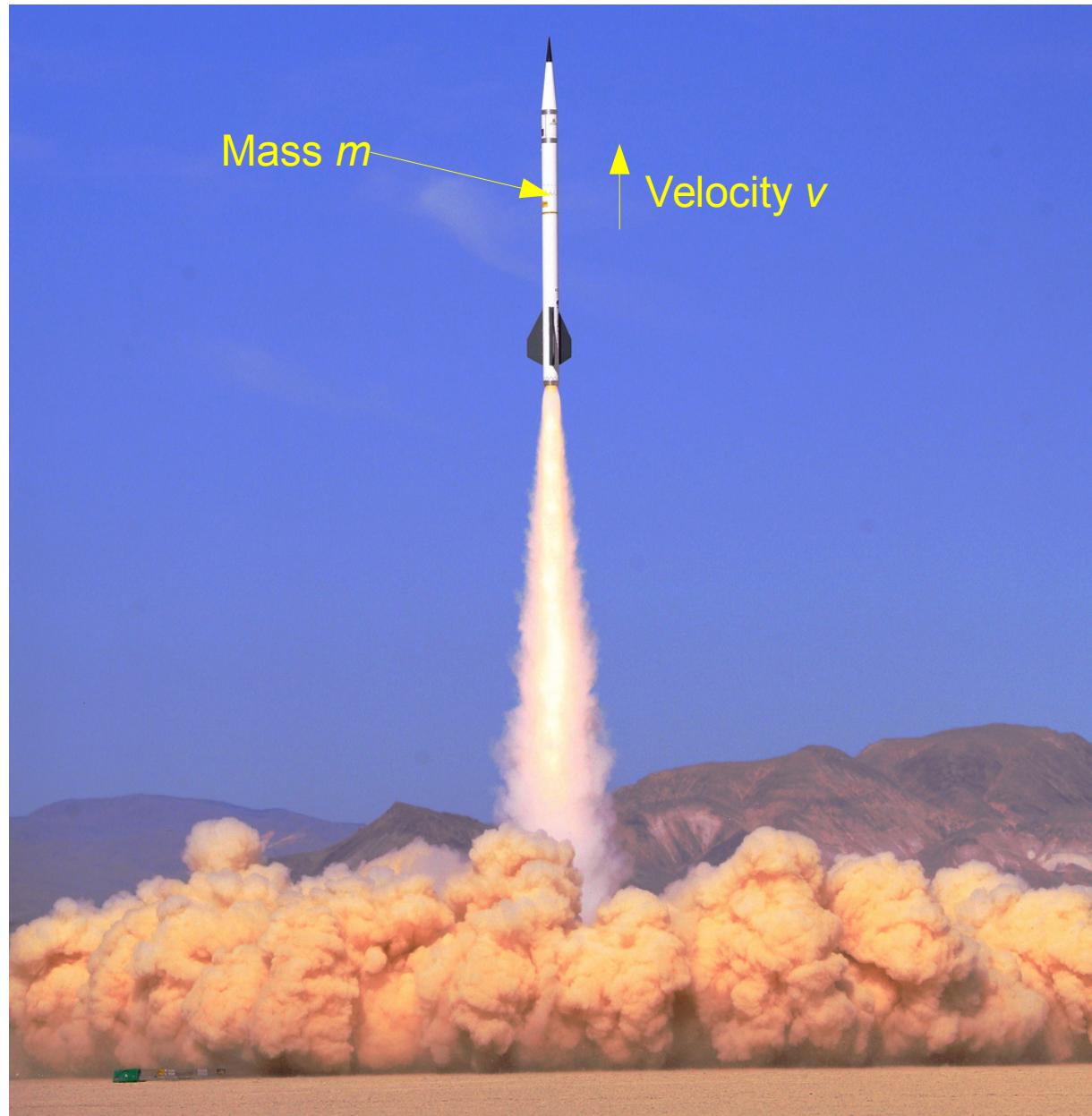
# Not So Simple Rotational Inertial Problem



$$\begin{aligned} I &= \int r^2(\theta) dm = \int_0^{2\pi} r^2(\theta) \cdot (\mu r(\theta) d\theta) \\ &= \mu \int_0^{2\pi} r^3(\theta) d\theta \end{aligned}$$

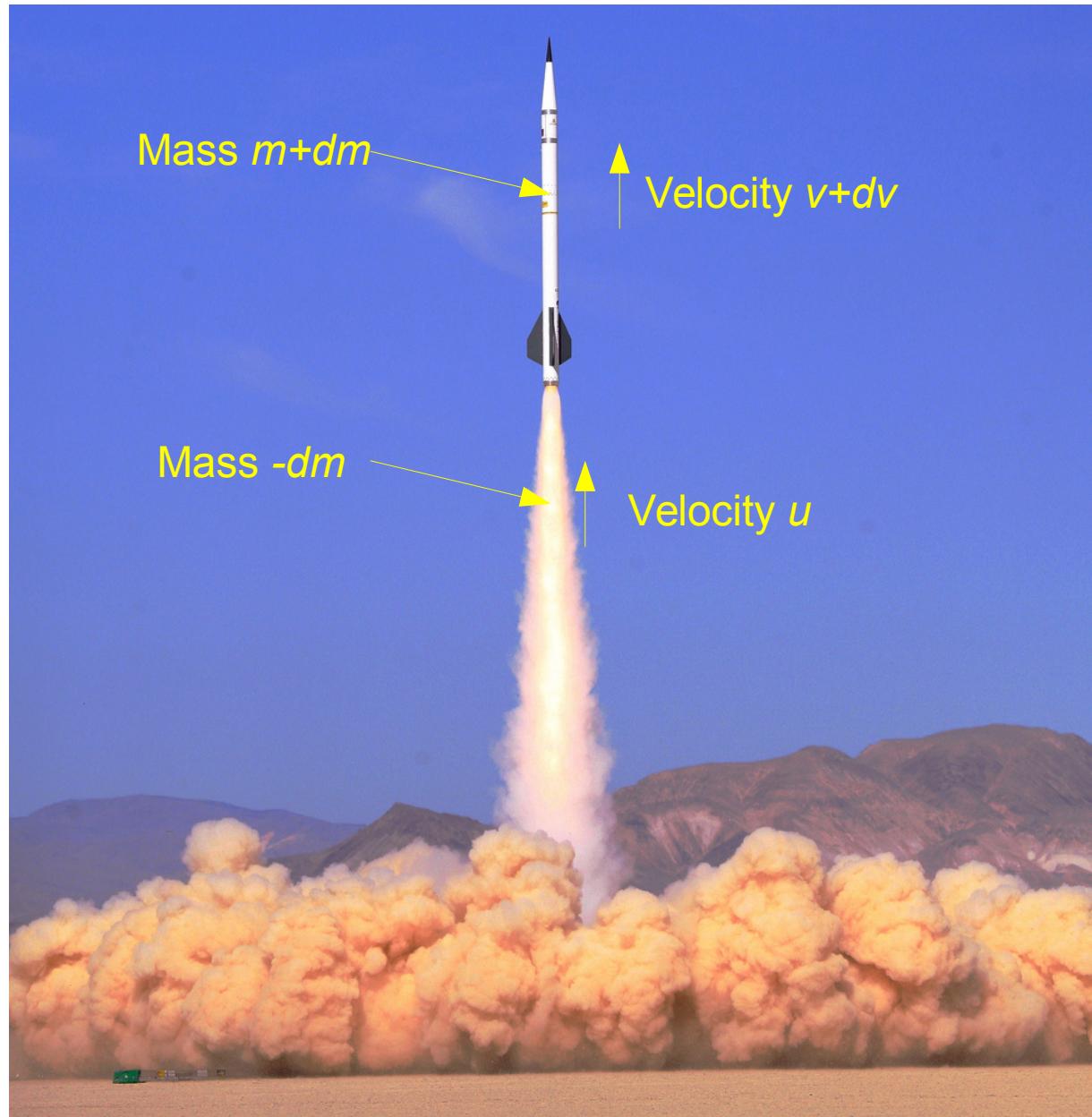
# Rocket Propulsion

“Initial” state:



# Rocket Propulsion

“Final” state,  
after small mass of  
fuel consumed and  
propelled out the  
nozzle is  $-dm$ :



# Rocket Propulsion

Conservation of linear momentum:

$$P_i = P_f$$
$$mv = -dm \cdot u + (m + dm)(v + dv)$$

Look at velocity of rocket relative to the exhaust in final state  $v_{rel}$

$$v + dv = v_{rel} + u$$

or

$$u = v + dv - v_{rel}$$

Substituting for  $u$ :

$$-dm \cdot v_{rel} = m dv$$

Divide by  $dt$  to see thrust and acceleration of rocket:

$$-\frac{dm}{dt} v_{rel} = T = m \frac{dv}{dt}$$

# Rocket Propulsion

Or integrate  $dv$  to find increase in rocket velocity over time:

$$dv = -v_{rel} \frac{dm}{m}$$
$$\int_{v_i}^{v_f} dv = - \int_{m_i}^{m_f} v_{rel} \cdot \frac{dm}{m}$$

If  $v_{rel}$  is constant

$$v_f - v_i = v_{rel} \ln\left(\frac{m_i}{m_f}\right)$$

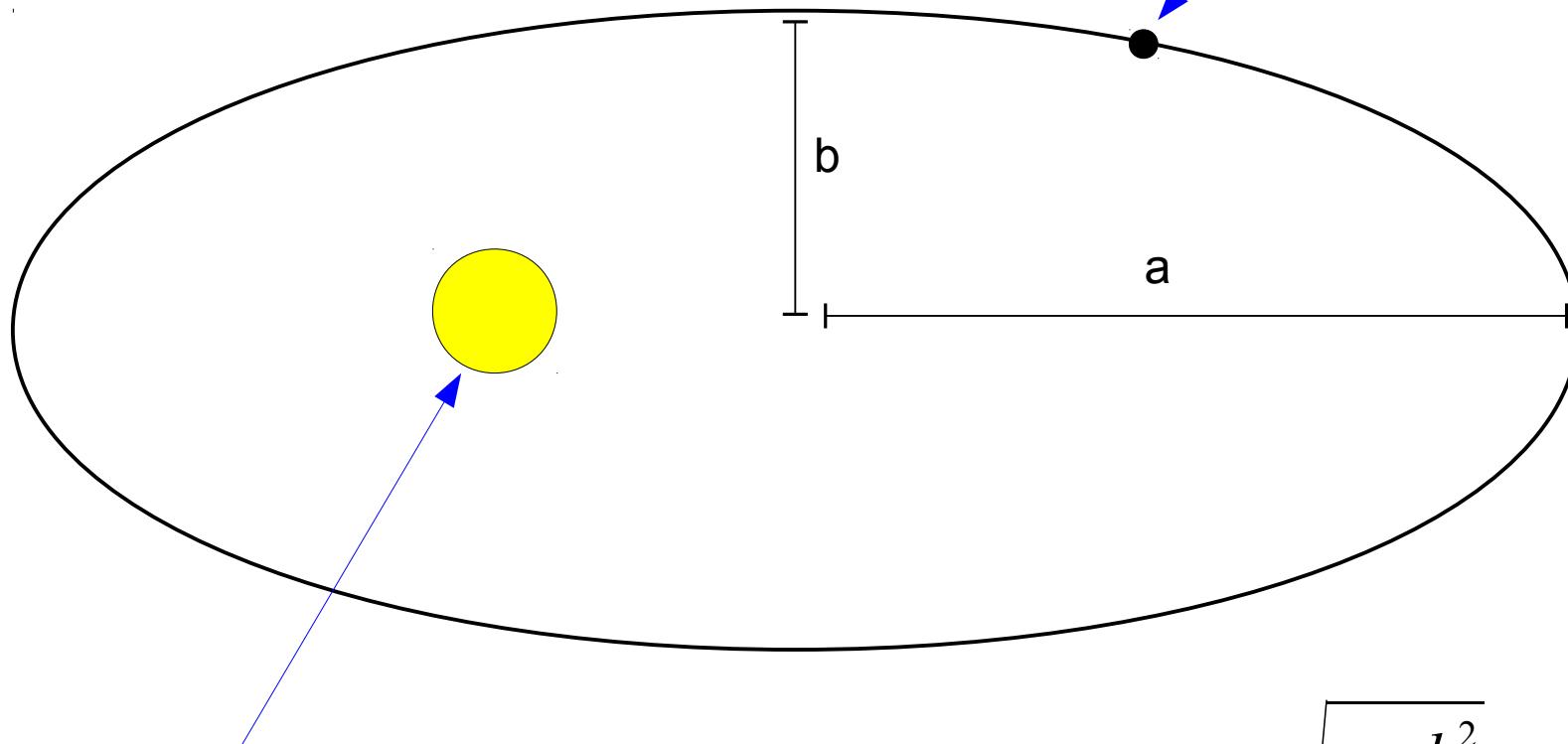
But in general

$$v_f - v_i = - \int_{m_i}^{m_f} \frac{v_{rel}(m)}{m} dm$$

# Circumference of Elliptical Orbits

$a$  = "Semi-major axis"  
 $b$  = "Semi-minor axis"

Comet in highly eccentric elliptical orbit



Sun at one focus of ellipse

$$\text{Eccentricity } \epsilon = \sqrt{1 - \frac{b^2}{a^2}}$$

# Circumference of Elliptical Orbits

The circumference of an elliptical orbit  $C = 4aE(\epsilon)$

where  $a$  is the length of the semi-major axis,  $\epsilon$  is the eccentricity and where the function  $E$  is the complete elliptic integral of the second kind:

$$E(\epsilon) = \int_0^{\frac{\pi}{2}} \sqrt{1 - \epsilon^2 \sin^2(\theta)} d\theta$$

This integral cannot be evaluated analytically.

Note: for a circular orbit,  $a=b=\text{radius}$ ,  $\epsilon=0$  and

$$E(\epsilon) = \int_0^{\frac{\pi}{2}} d\theta = \frac{\pi}{2}$$

so  $C = 4a \frac{\pi}{2} = 2\pi a$  as expected

# Plotting Vectors with gnuplot

We have used:

```
plot 'problem1.dat' using 1:2 with lines
```

```
plot 'problem2.dat' using 1:2 with points
```

Column with  
X coordinate

Column with  
Y coordinate

Another plot mode:

```
plot 'helmholtz.dat' using 1:2:3:4 with vectors
```

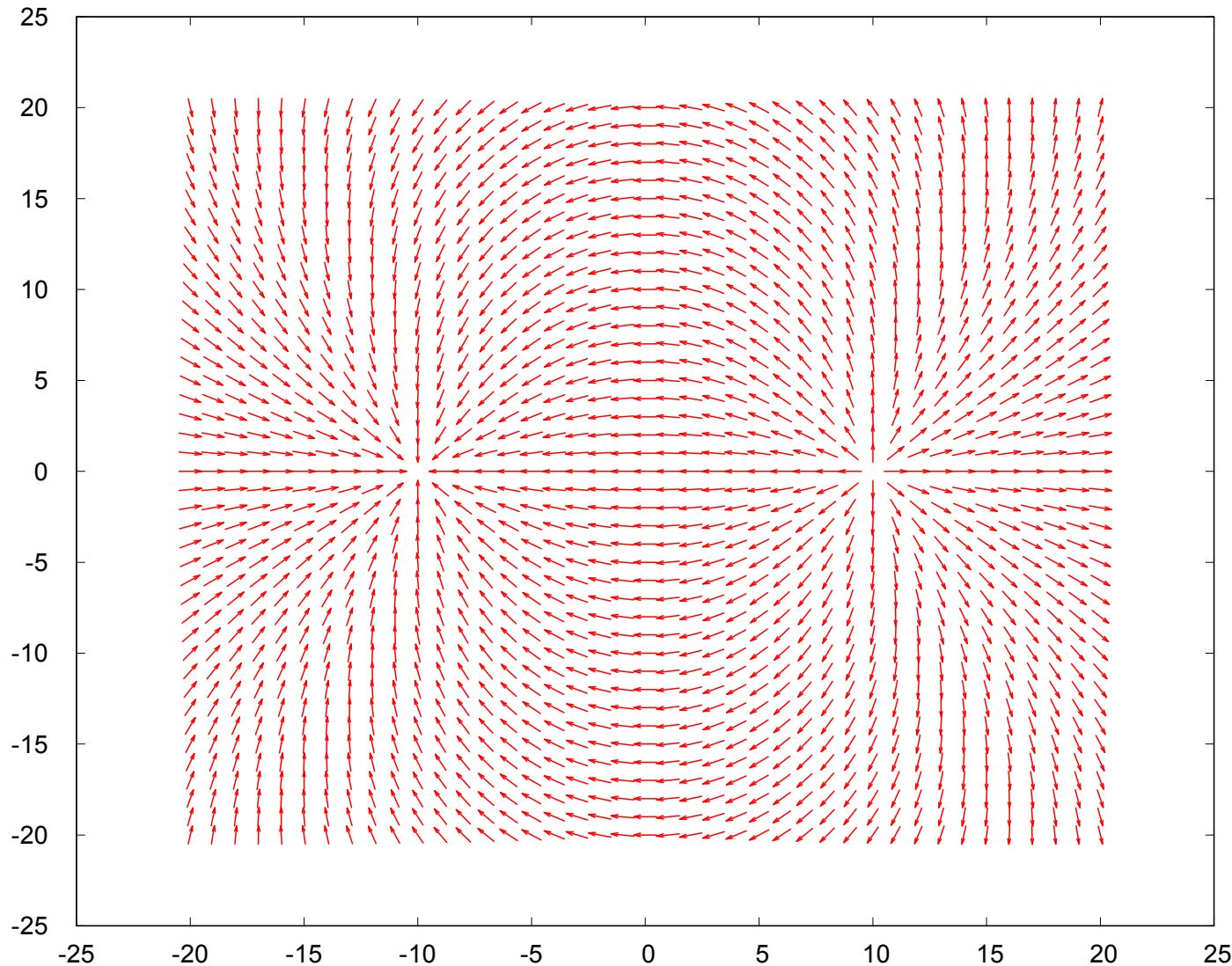
Column with  
starting X  
coordinate

Column with  
starting Y  
coordinate

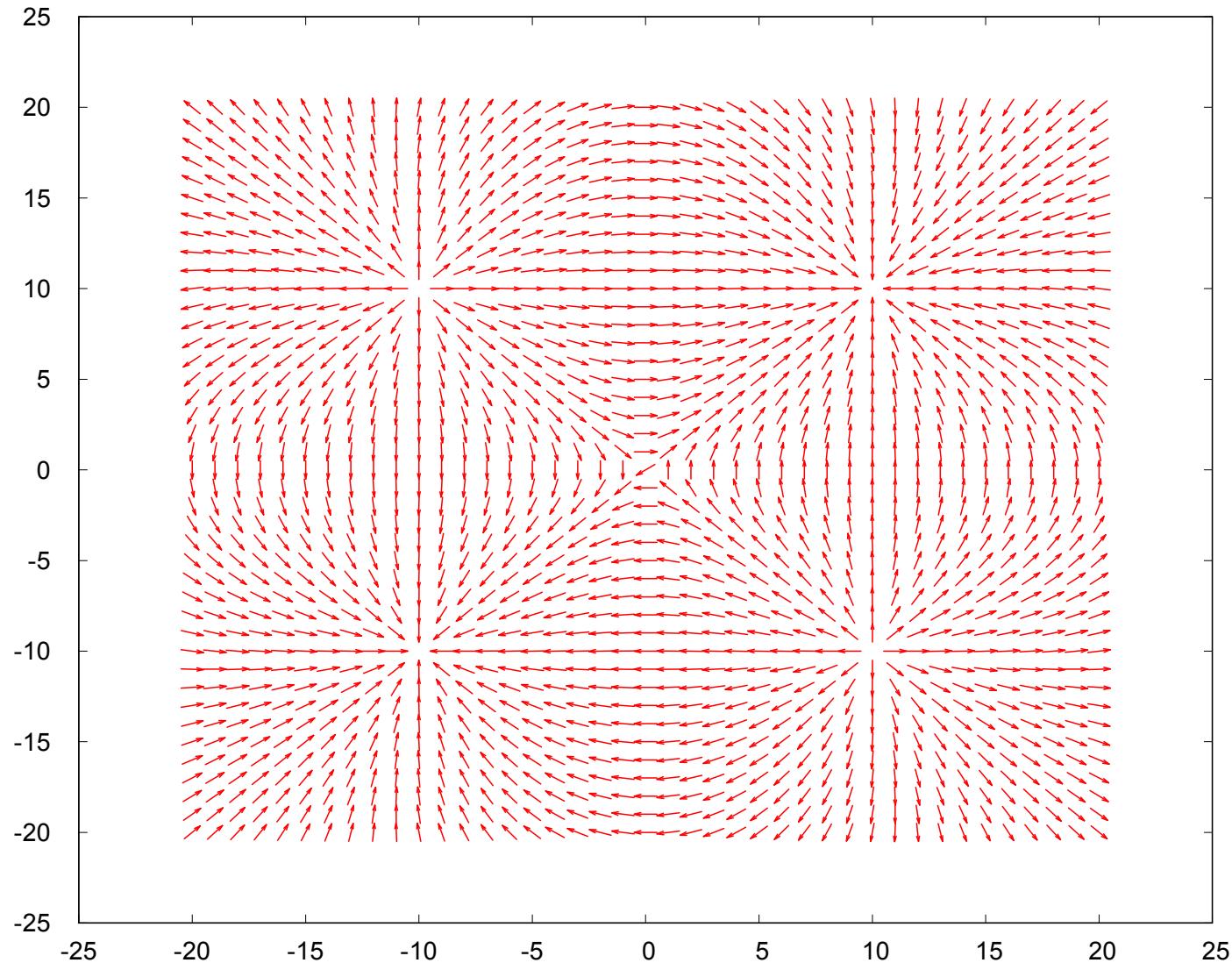
Column with  
X distance

Column with  
Y distance

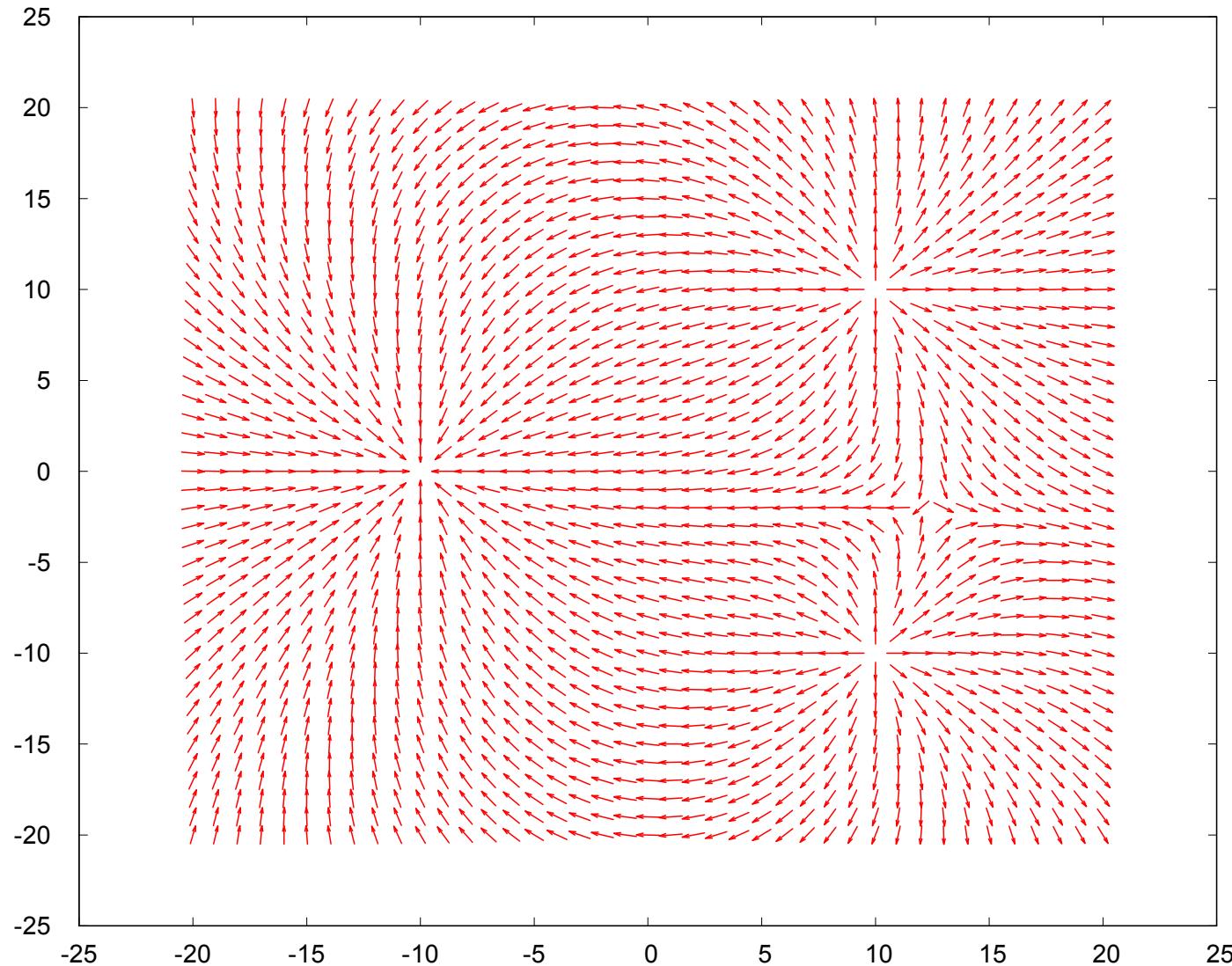
# Electric Field Vectors for Dipole



# Electric Field Vectors for Quadrapole



# Electric Field Vectors for Example Arbitrary Charges



# C Code for Mapping Electric Field Vectors

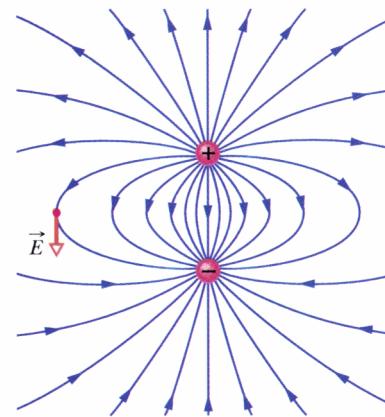
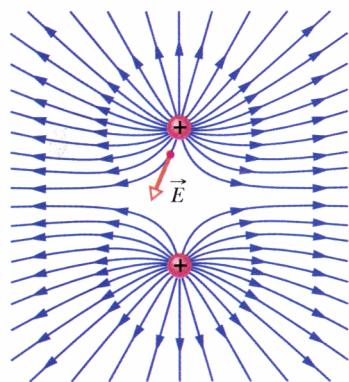
```
xstop = xstop + xinc * 0.5;
ystop = ystop + yinc * 0.5;
rtest = 1.0e-6 * (xinc*xinc + yinc*yinc);
x = xstart;
while (((xinc > 0.0) && (x < xstop)) || ((xinc < 0.0) && (x > xstop))) {
    y = ystart;
    while (((yinc > 0.0) && (y < ystop)) || ((yinc < 0.0) && (y > ystop))) {
        ex = 0.0;
        ey = 0.0;
        i = 0;
        while (i < n) { ←
            rsq = (x-chg[i].x)*(x-chg[i].x) + (y-chg[i].y)*(y-chg[i].y);
            if (rsq < rtest) break;
            r = sqrt(rsq);
            ex += chg[i].q * (x-chg[i].x) / (r * rsq);
            ey += chg[i].q * (y-chg[i].y) / (r * rsq);
            i++;
        } ←
        if (i >= n) {
            emag = sqrt(ex*ex + ey*ey);
            vx = xinc * ex / emag;
            vy = yinc * ey / emag;
            printf("%.8g %.8g %.8g %.8g\n", x-0.5*vx, y-0.5*vy, vx, vy);
        }
        y = y + yinc;
    }
    x = x + xinc;
}
```

Loop over  $n$  charges.  
Structure array  $\text{chg}[]$  holds  
each charge  $q_i$  and  
 $x_i$  and  $y_i$  coordinates

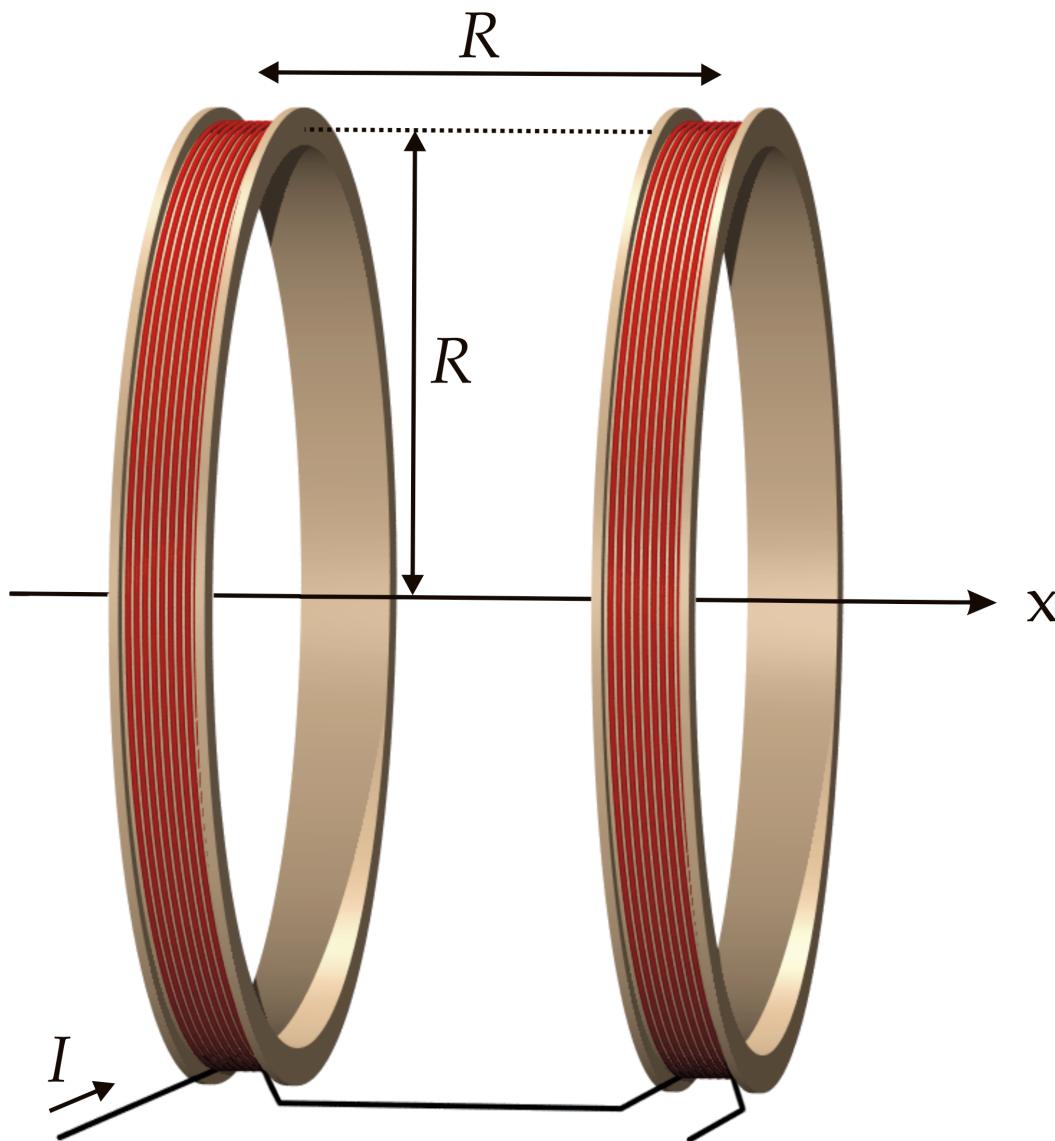


# Classical Electric Field Lines

Note that the vectors plotted by this method only record the direction of the E field vector at an array of sampled points. These are not quite equivalent to the classical field lines, which convey both the direction of the field and its magnitude from the density of drawn lines.



# Helmholtz Coil for Uniform Magnetic Field



# Magnetic Field of a Helmholtz Coil

Coordinates of two conductors follow the parameterized equations

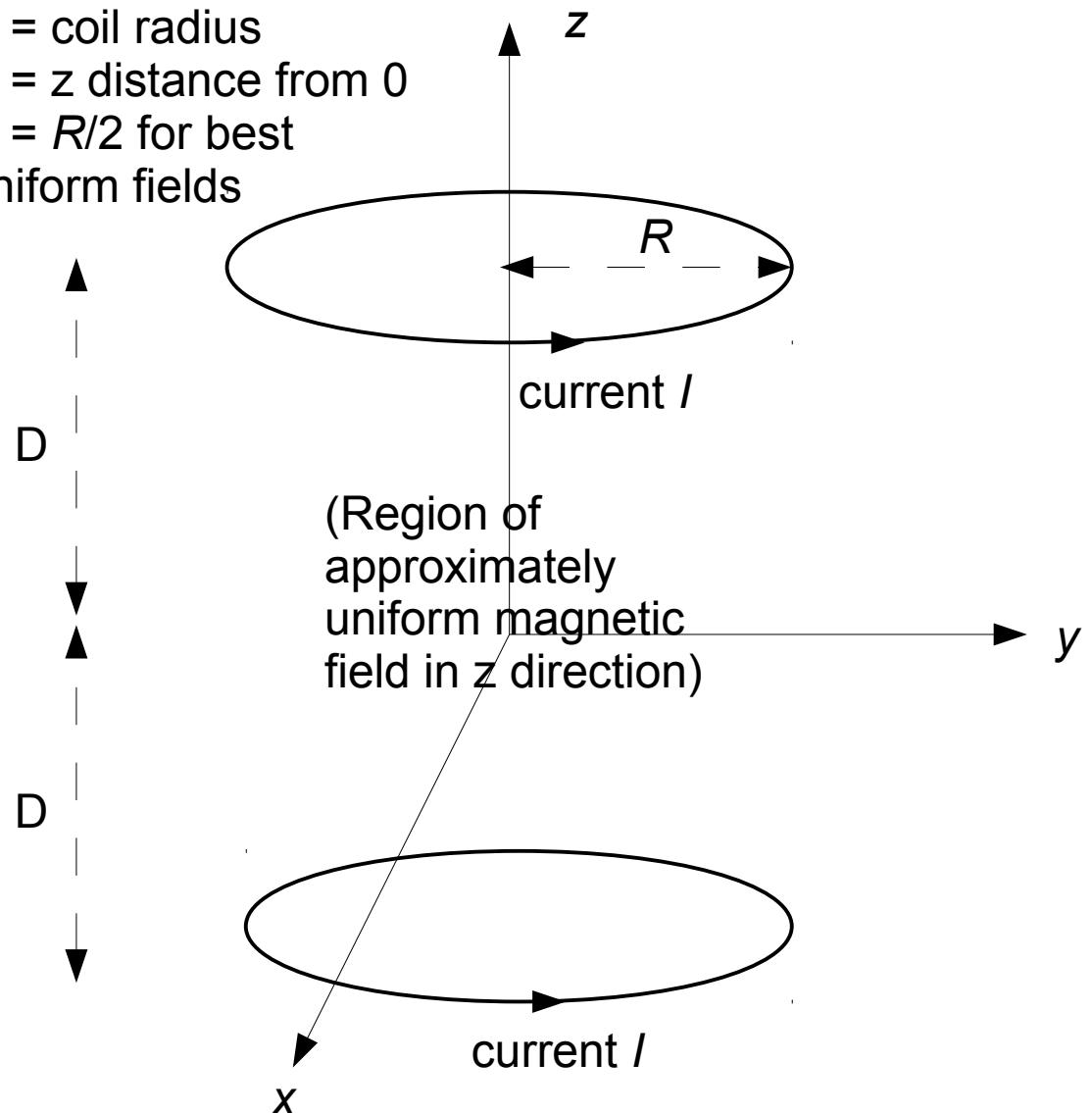
$$x = R \cos(\theta)$$

$$y = R \sin(\theta)$$

$$z = \pm D = \pm \frac{R}{2}$$

for  $-\pi \leq \theta \leq \pi$

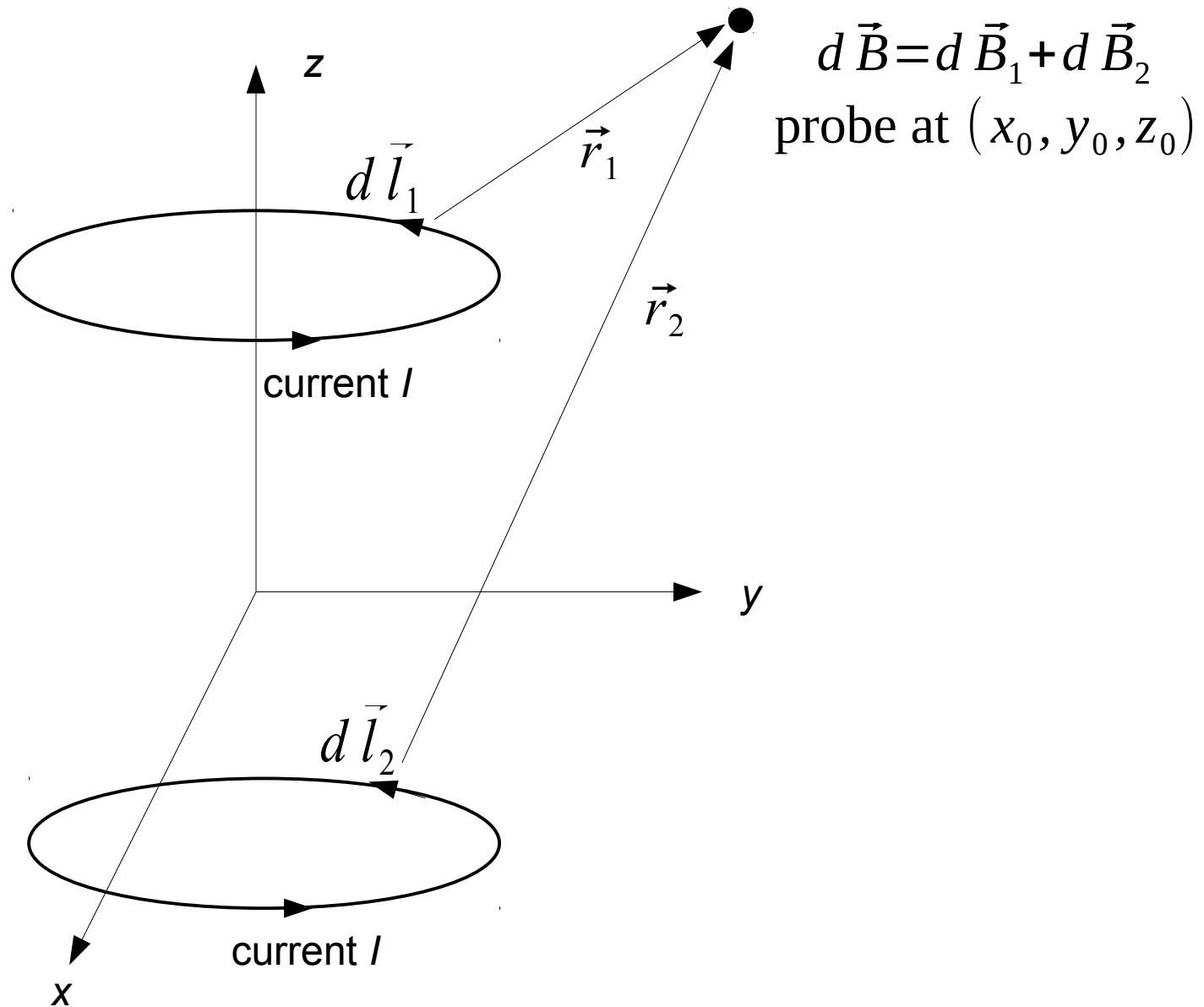
$R$  = coil radius  
 $D$  = z distance from 0  
 $D = R/2$  for best uniform fields



# Finding Magnetic Field with Biot-Savart Law

$$d\vec{B}_i = \frac{\mu_0}{4\pi} \frac{I d\vec{l}_i \times \vec{r}_i}{r_i^3}$$

$$\vec{B}(\vec{r}) = \int d\vec{B}$$



# Review of Vector Cross Product

if  $\vec{C} = \vec{A} \times \vec{B}$

then

$$C_x = A_y B_z - A_z B_y$$

$$C_y = A_z B_x - A_x B_z$$

$$C_z = A_x B_y - A_y B_x$$

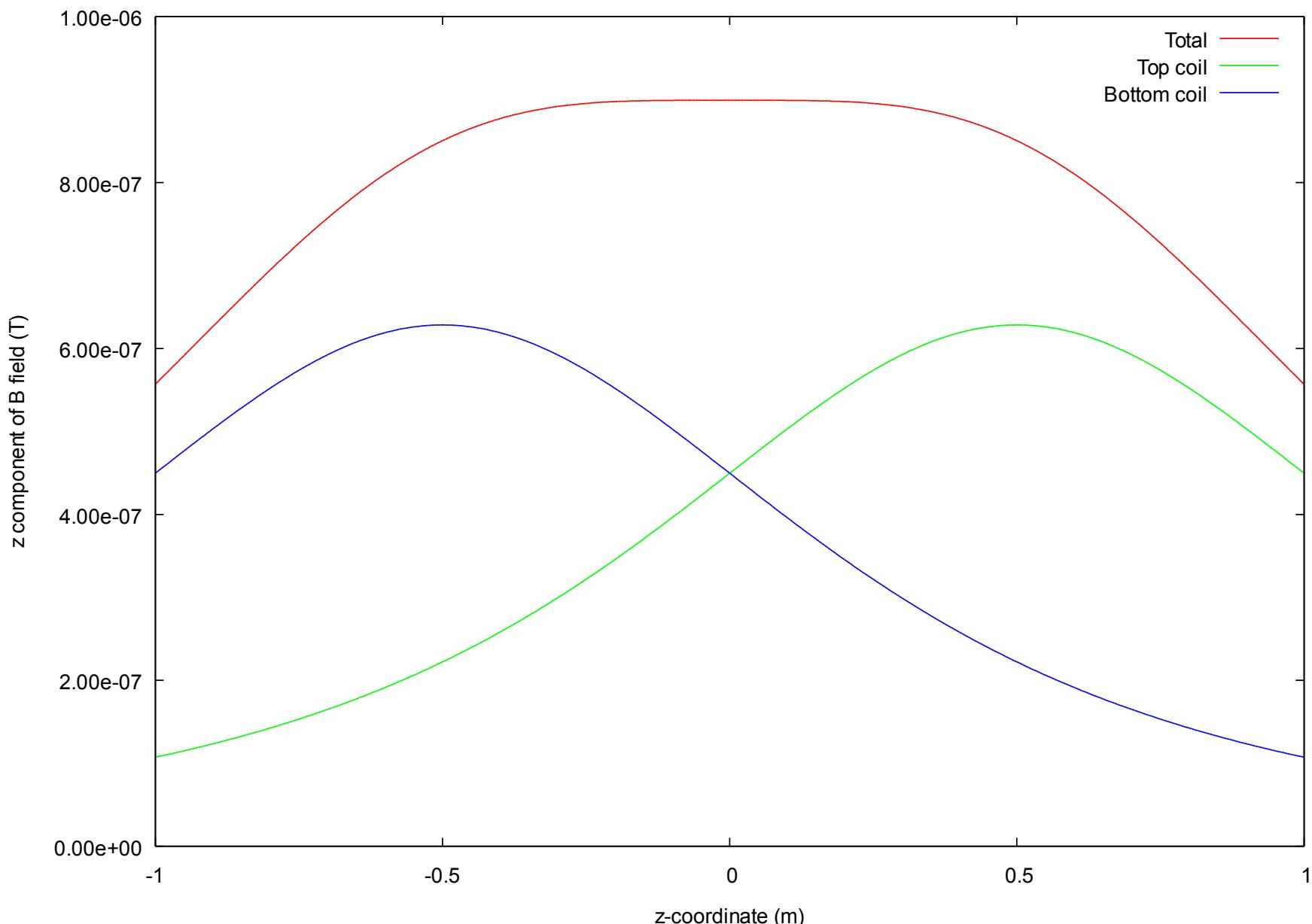
When computing

$$d\vec{l} \times \vec{r}$$

for this geometry, look for several components and terms that are zero!

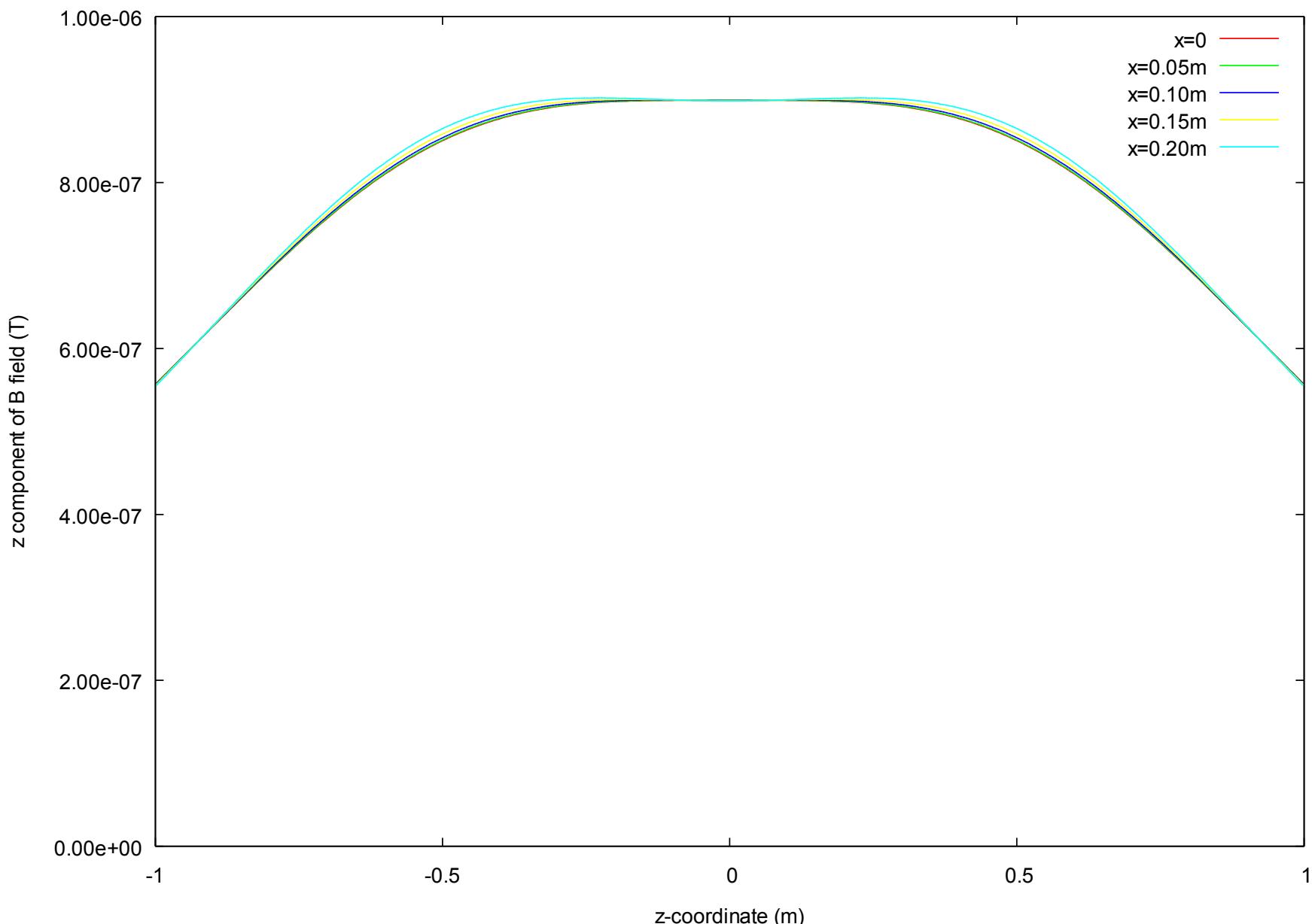
# Magnetic Field Along z-Axis

R=1m I=1A x=0 y=0



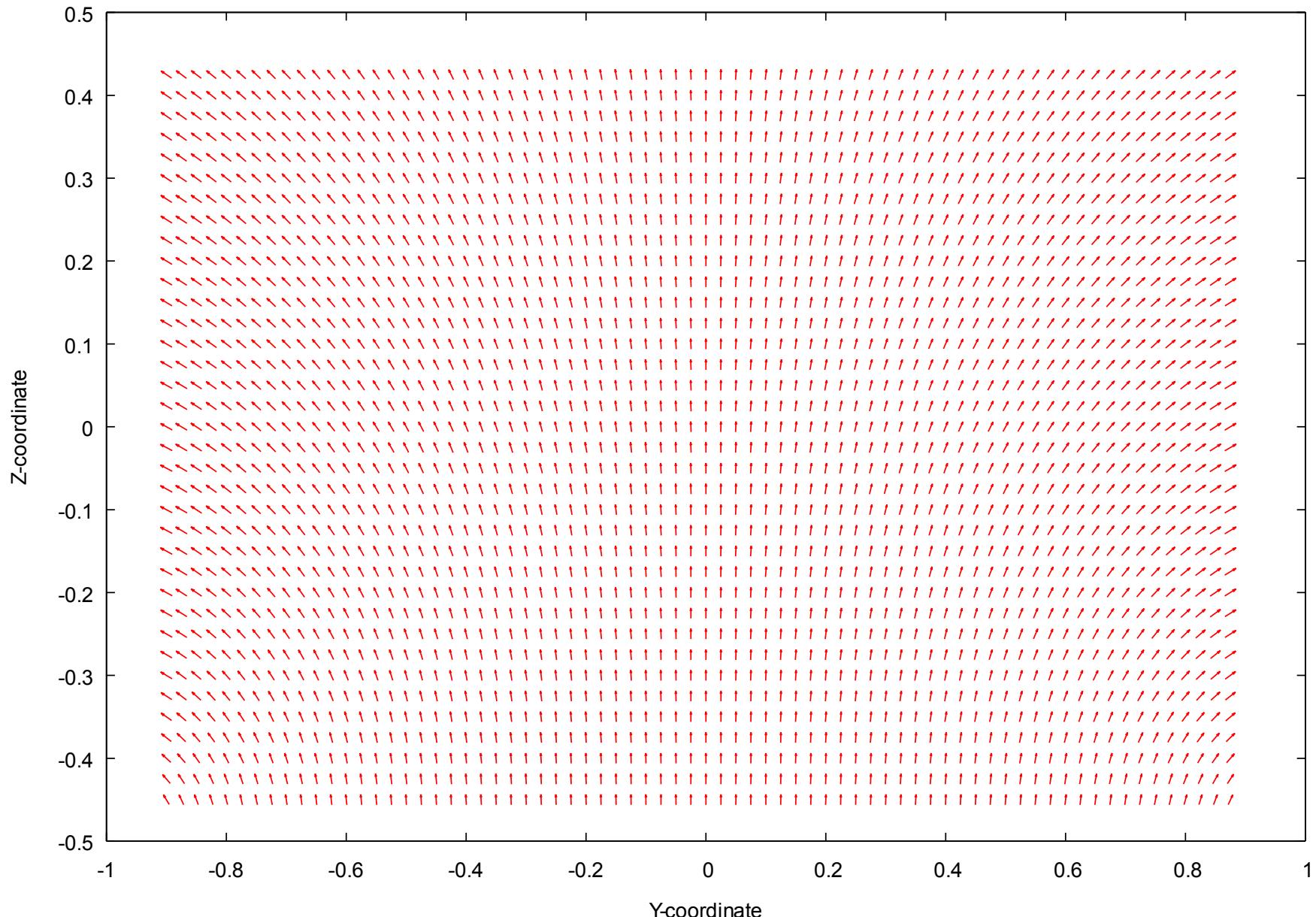
# Magnetic Field Off Of z-Axis

R=1m I=1A y=0



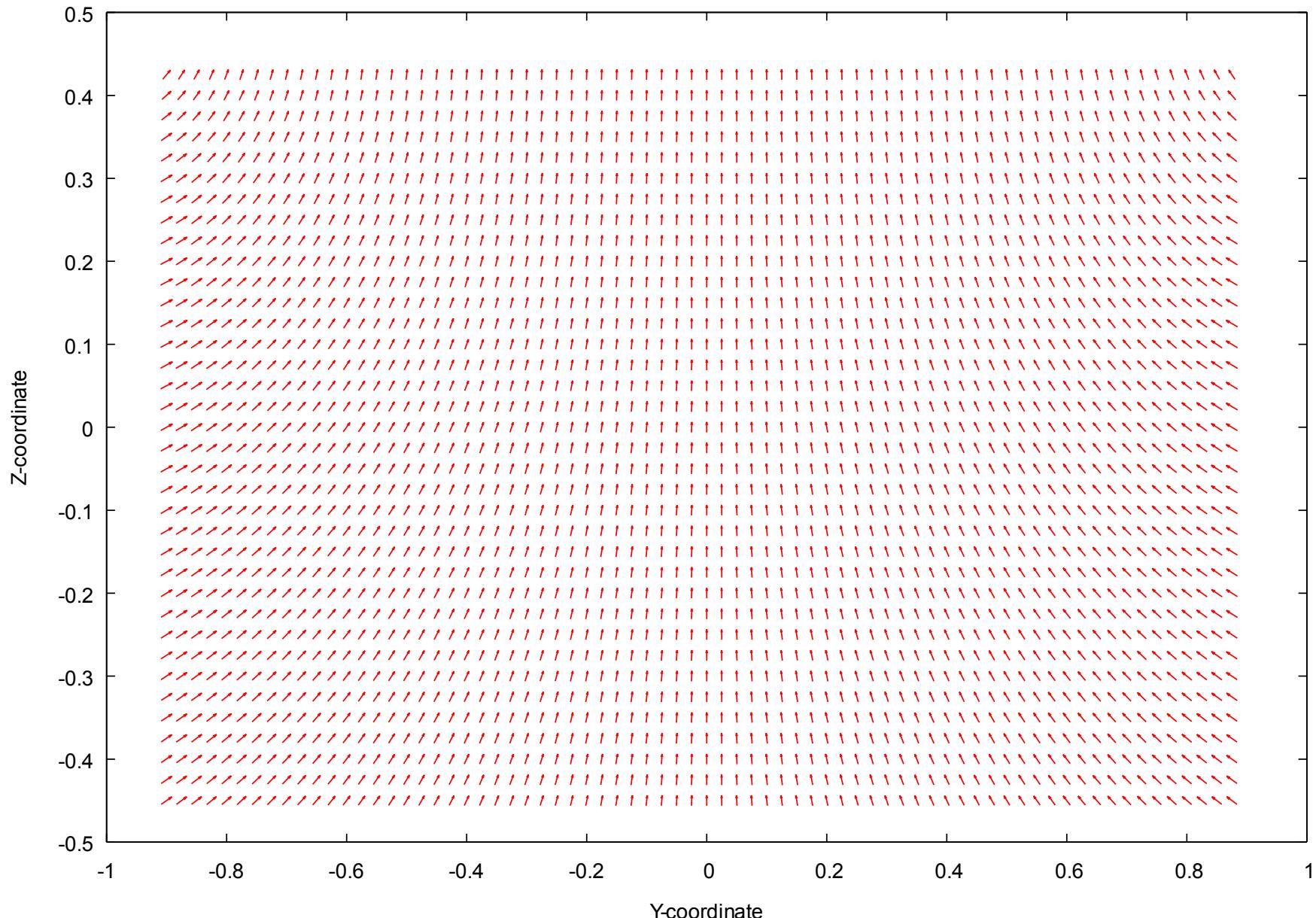
# Vector Plot of Magnetic Field in Y-Z Plane

Bottom Coil Only



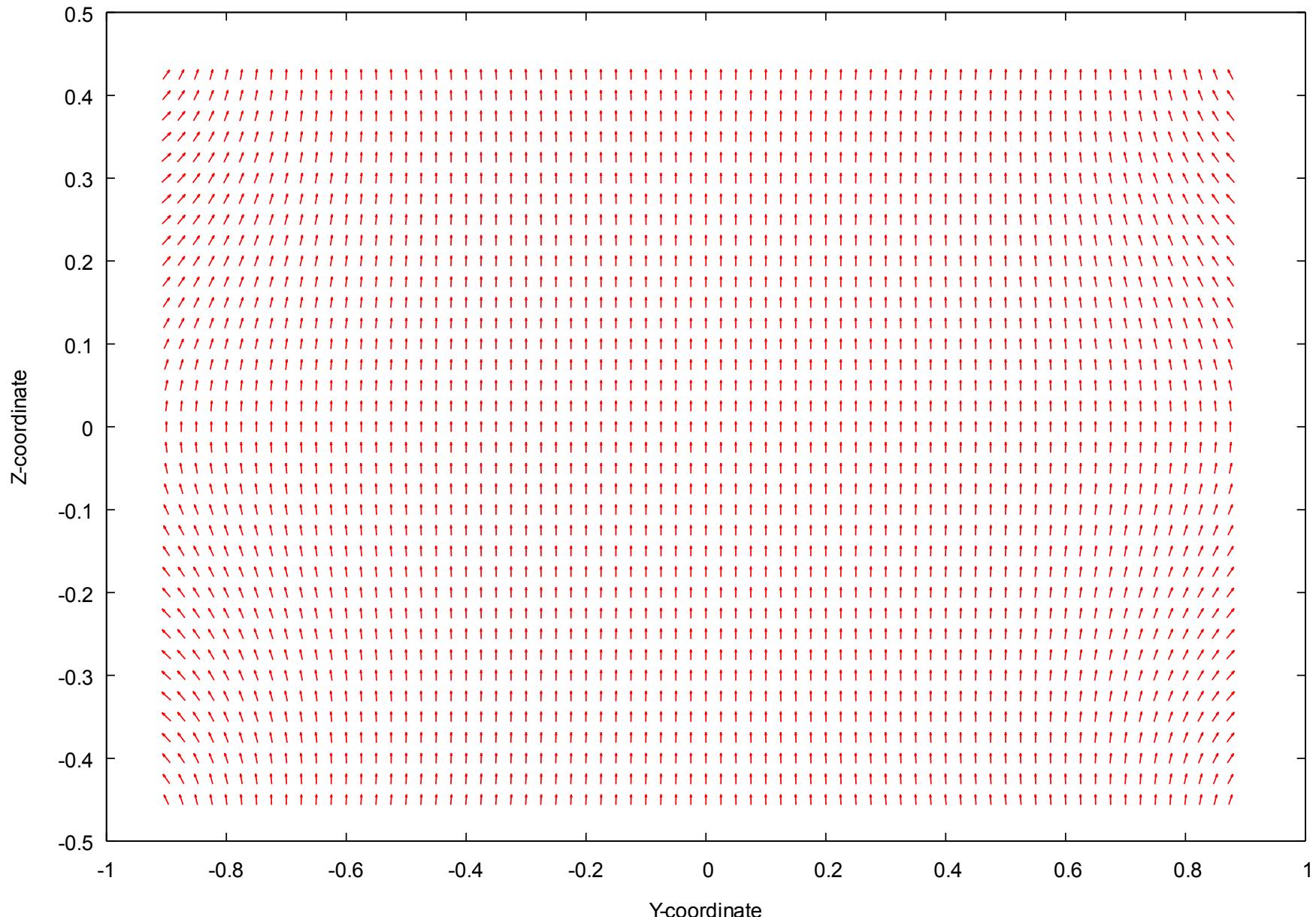
# Vector Plot of Magnetic Field in Y-Z Plane

Top Coil Only



# Vector Plot of Magnetic Field in Y-Z Plane

Both Coils



# Magnetic Field of a Real Solenoid

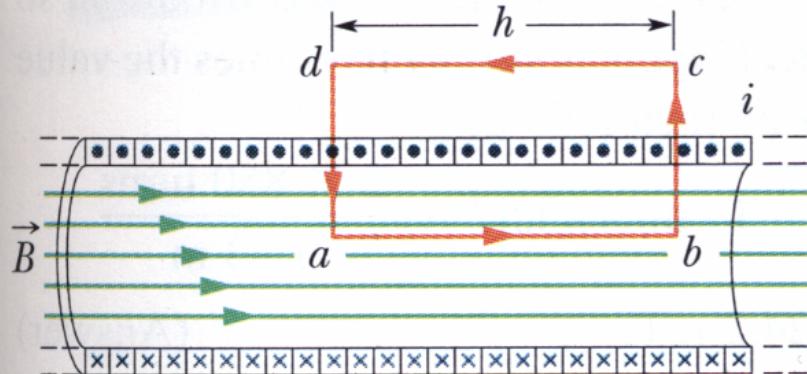


Fig. 30-19 Application of Ampere's law to a section of a long ideal solenoid carrying a current *i*. The Amperian loop is the rectangle *abcd*.

$$B = \mu_0 i n$$

where *n* = turns per unit length

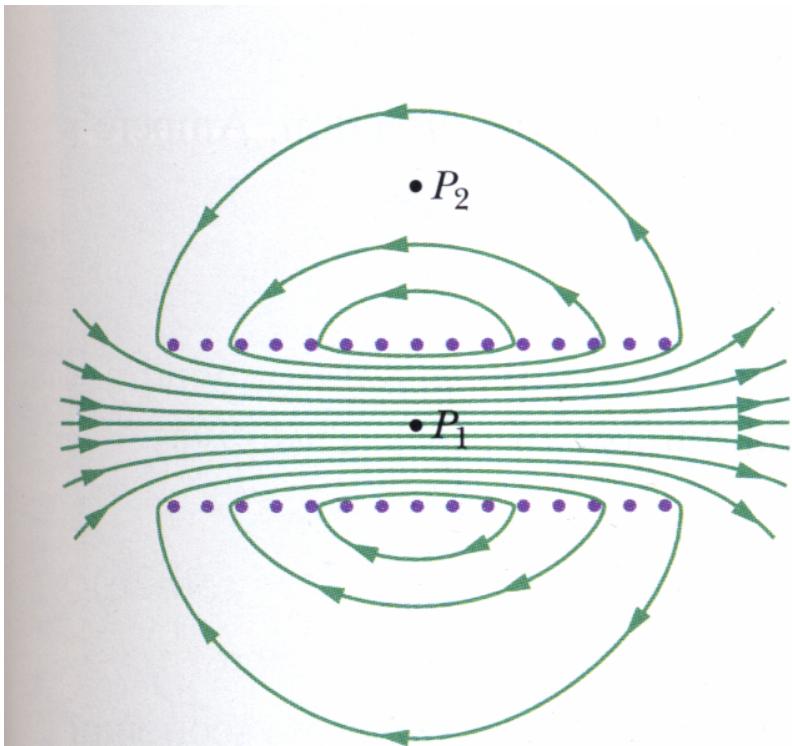


Fig. 30-18 Magnetic field lines for a real solenoid of finite length. The field is strong and uniform at interior points such as *P<sub>1</sub>* but relatively weak at external points such as *P<sub>2</sub>*.

# Magnetic Field of a Real Solenoid

Coordinates of solenoid conductor follow the parameterized equations

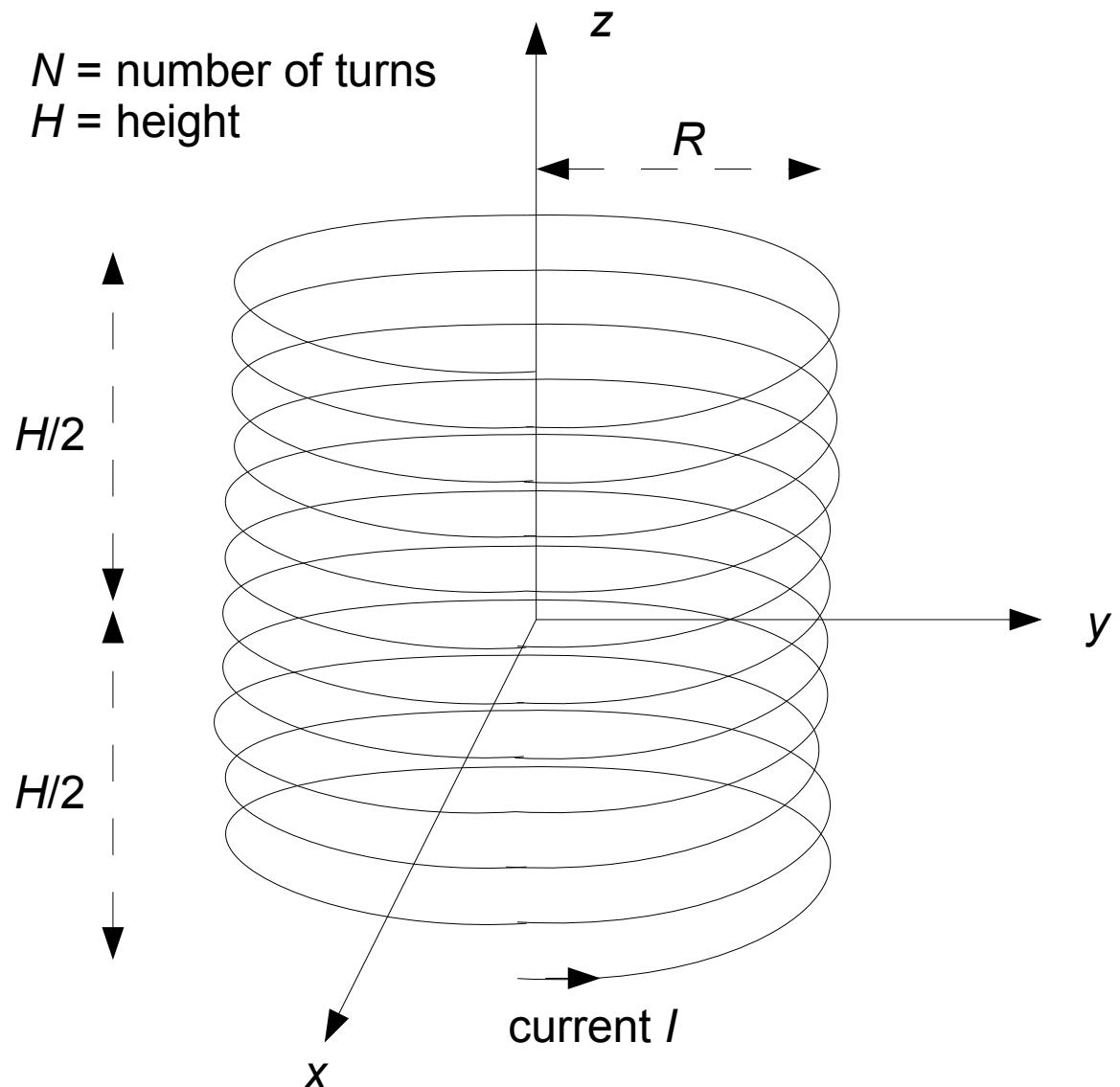
$$x = R \cos(\theta)$$

$$y = R \sin(\theta)$$

$$z = a\theta$$

for  $-N\pi \leq \theta \leq N\pi$

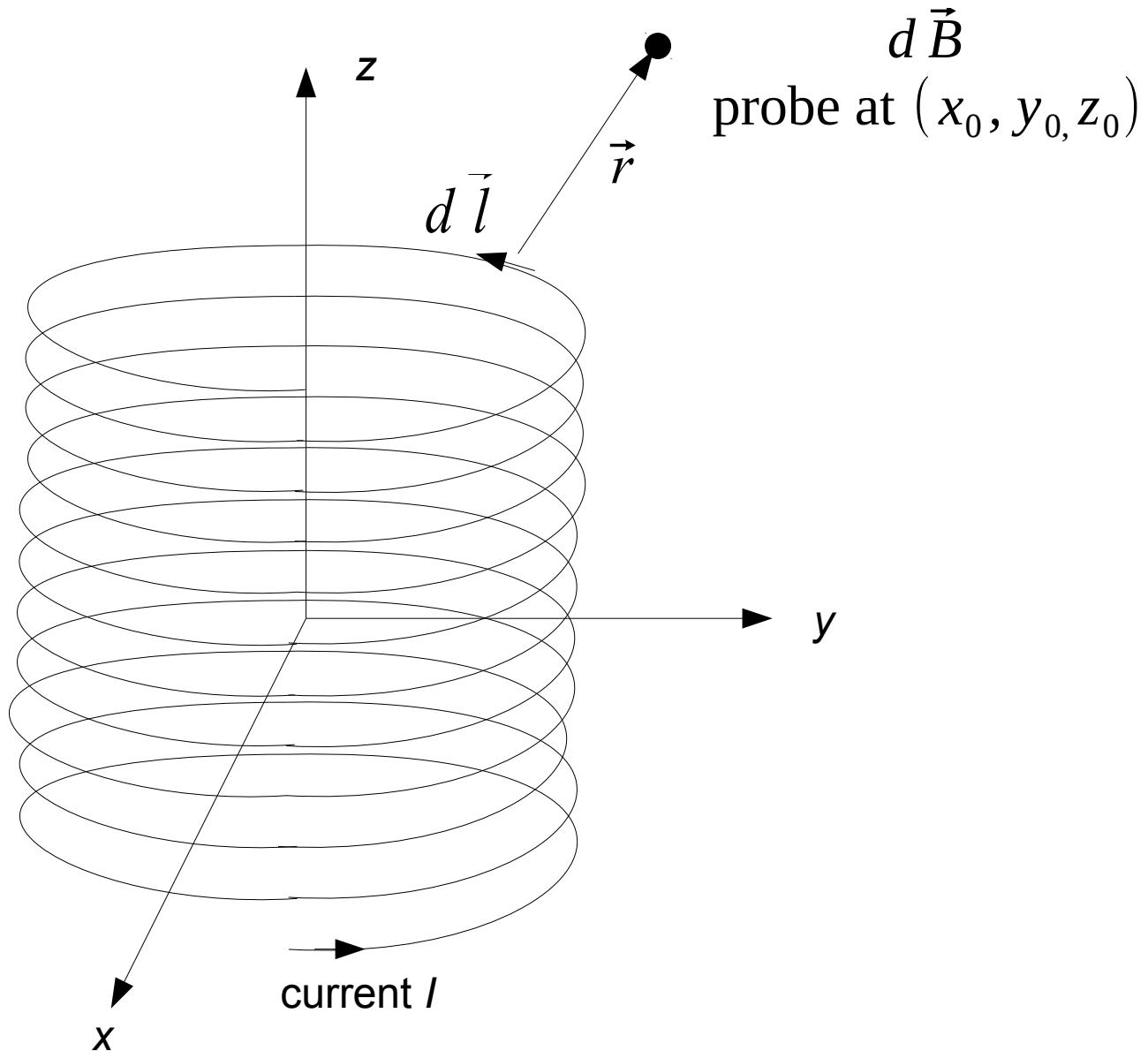
and where  $a = \frac{H}{2\pi N}$



# Finding Magnetic Field with Biot-Savart Law

$$d\vec{B} = \frac{\mu_0}{4\pi} \frac{I d\vec{l} \times \vec{r}}{r^3}$$

$$\vec{B}(\vec{r}) = \int d\vec{B}$$



$d\vec{B}$   
probe at  $(x_0, y_0, z_0)$

# Beware of Random Walks of Roundoff Error!

Example: edit integration lab program to try to compute pi:

$$\pi = \int_0^1 4\sqrt{1-x^2} dx$$

# Beware of Random Walks of Roundoff Error!

```
#include <stdio.h>
#include <stdlib.h>
#include <math.h>
#include "integrate.h"

double func(double x) {
    return(4.0 * sqrt(1.0 - x*x));
}

int main(int argc,char *argv[])
{
    double xinc;

    if (argc != 2) {
        fprintf(stderr,"%s <xinc>\n",argv[0]);
        exit(1);
    }
    xinc = atof(argv[1]);
    printf("Trapezoidal rule:\n");
    trapezoidal(func,0.0,1.0,xinc);
    printf("Simpson's rule:\n");
    simpson(func,0.0,1.0,xinc);
    exit(0);
}
```



# Beware of Random Walks of Roundoff Error!

```
$ ./pi_integrate 0.01
```

```
Trapezoidal rule:
```

```
-nan
```

```
Simpson's rule:
```

```
-nan
```

```
$ ./pi_integrate 0.0025
```

```
Trapezoidal rule:
```

```
3.14144566781
```

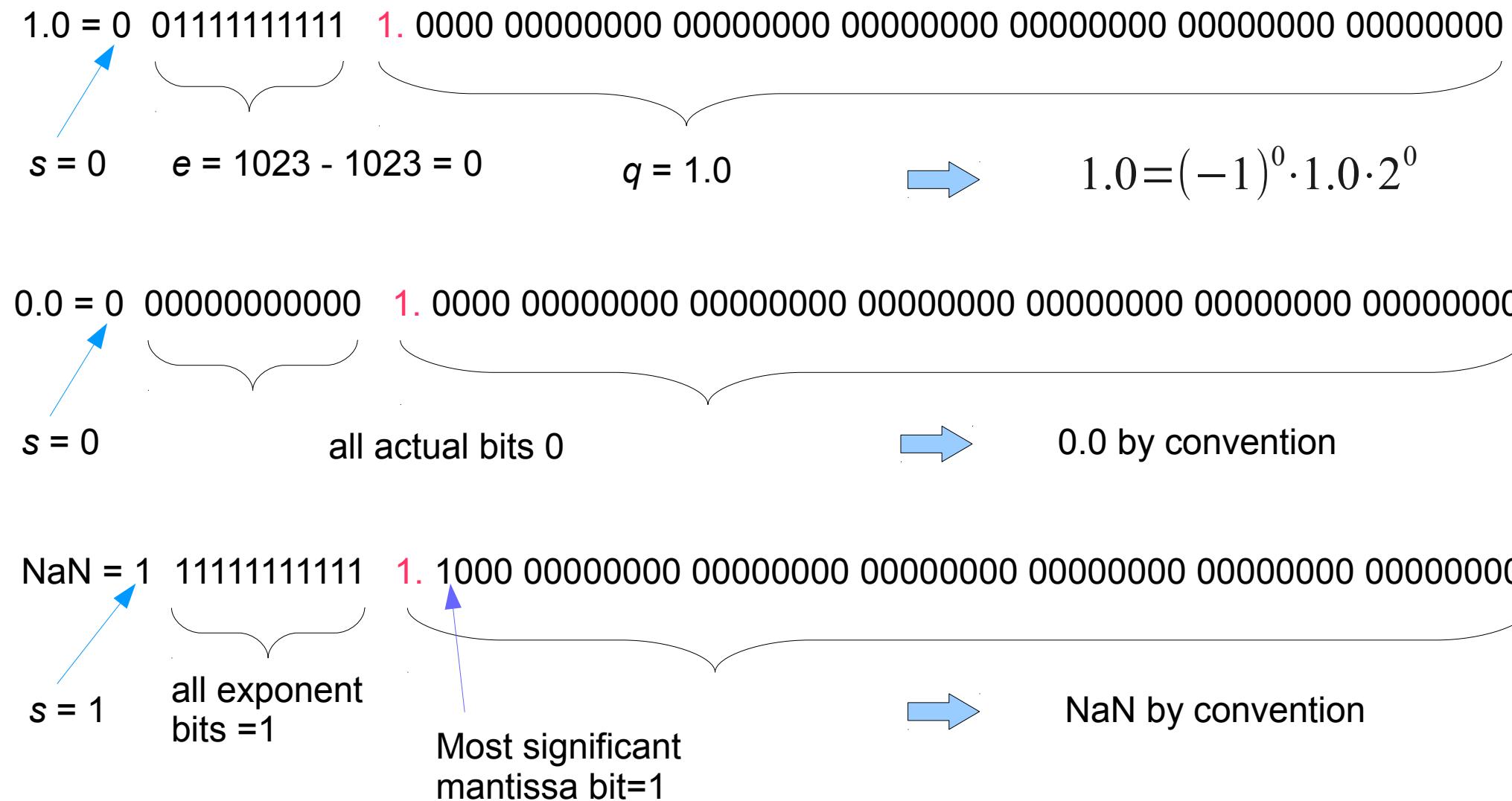
```
Simpson's rule:
```

```
3.14153524166
```

Try two different step sizes through the integration routines

So what is this nan business???

# NaN Binary Storage Convention



# The man Page for sqrt()

SQRT(3)

Linux Programmer's Manual

SQRT(3)

## NAME

`sqrt, sqrtf, sqrtl` - square root function

## SYNOPSIS

```
#include <math.h>

double sqrt(double x);
float sqrtf(float x);
long double sqrtl(long double x);
```

Link with `-lm`.

## DESCRIPTION

The `sqrt()` function returns the nonnegative square root of `x`.

## RETURN VALUE

On success, these functions return the square root of `x`.

If `x` is a NaN, a NaN is returned.

If `x` is +0 (-0), +0 (-0) is returned.

If `x` is positive infinity, positive infinity is returned.

If `x` is less than -0, a domain error occurs, and a NaN is returned.

!!!