

P34.2 (a) Since the light from this star travels at 3.00×10^8 m/s
the last bit of light will hit the Earth in $\frac{6.44 \times 10^{18} \text{ m}}{3.00 \times 10^8 \text{ m/s}} = 2.15 \times 10^{10} \text{ s} = 680 \text{ years}$.
Therefore, it will disappear from the sky in the year $2004 + 680 = \boxed{2.68 \times 10^3 \text{ C.E.}}$.
The star is 680 light-years away.

(b) $\Delta t = \frac{\Delta x}{v} = \frac{1.496 \times 10^{11} \text{ m}}{3 \times 10^8 \text{ m/s}} = \boxed{499 \text{ s}} = 8.31 \text{ min}$

(c) $\Delta t = \frac{\Delta x}{v} = \frac{2(3.84 \times 10^8 \text{ m})}{3 \times 10^8 \text{ m/s}} = \boxed{2.56 \text{ s}}$

(d) $\Delta t = \frac{\Delta x}{v} = \frac{2\pi(6.37 \times 10^6 \text{ m})}{3 \times 10^8 \text{ m/s}} = \boxed{0.133 \text{ s}}$

(e) $\Delta t = \frac{\Delta x}{v} = \frac{10 \times 10^3 \text{ m}}{3 \times 10^8 \text{ m/s}} = \boxed{3.33 \times 10^{-5} \text{ s}}$

P34.3 $v = \frac{1}{\sqrt{\kappa\mu_0 \epsilon_0}} = \frac{1}{\sqrt{1.78}} c = 0.750c = \boxed{2.25 \times 10^8 \text{ m/s}}$

P34.5 (a) $f\lambda = c$
or $f(50.0 \text{ m}) = 3.00 \times 10^8 \text{ m/s}$
so $\boxed{f = 6.00 \times 10^6 \text{ Hz} = 6.00 \text{ MHz}}$.

(b) $\frac{E}{B} = c$
or $\frac{22.0}{B_{\max}} = 3.00 \times 10^8$
so $\mathbf{B}_{\max} = \boxed{-73.3 \hat{\mathbf{k}} \text{ nT}}$.

(c) $k = \frac{2\pi}{\lambda} = \frac{2\pi}{50.0} = 0.126 \text{ m}^{-1}$
and $\omega = 2\pi f = 2\pi(6.00 \times 10^6 \text{ s}^{-1}) = 3.77 \times 10^7 \text{ rad/s}$
 $\mathbf{B} = \mathbf{B}_{\max} \cos(kx - \omega t) = \boxed{-73.3 \cos(0.126x - 3.77 \times 10^7 t) \hat{\mathbf{k}} \text{ nT}}$.

P34.11 $S = I = \frac{U}{At} = \frac{Uc}{V} = uc$ $\frac{\text{Energy}}{\text{Unit Volume}} = u = \frac{I}{c} = \frac{1000 \text{ W/m}^2}{3.00 \times 10^8 \text{ m/s}} = \boxed{3.33 \mu\text{J/m}^3}$

P34.15 Power output = (power input)(efficiency).

Thus,
$$\text{Power input} = \frac{\text{Power out}}{\text{eff}} = \frac{1.00 \times 10^6 \text{ W}}{0.300} = 3.33 \times 10^6 \text{ W}$$

and
$$A = \frac{\mathcal{P}}{I} = \frac{3.33 \times 10^6 \text{ W}}{1.00 \times 10^3 \text{ W/m}^2} = \boxed{3.33 \times 10^3 \text{ m}^2}.$$

P34.26 The pressure P upon the mirror is

$$P = \frac{2S_{\text{av}}}{c}$$

where A is the cross-sectional area of the beam and

$$S_{\text{av}} = \frac{\mathcal{P}}{A}.$$

The force on the mirror is then

$$F = PA = \frac{2}{c} \left(\frac{\mathcal{P}}{A} \right) A = \frac{2\mathcal{P}}{c}.$$

Therefore,

$$F = \frac{2(100 \times 10^{-3})}{(3 \times 10^8)} = \boxed{6.67 \times 10^{-10} \text{ N}}.$$

P34.59 Think of light going up and being absorbed by the bead which presents a face area πr_b^2 .

The light pressure is $P = \frac{S}{c} = \frac{I}{c}$.

(a)
$$F_t = \frac{I\pi r_b^2}{c} = mg = \rho \frac{4}{3} \pi r_b^3 g \quad \text{and} \quad I = \frac{4\rho g c}{3} \left(\frac{3m}{4\pi\rho} \right)^{1/3} = \boxed{8.32 \times 10^7 \text{ W/m}^2}$$

(b)
$$\mathcal{P} = IA = (8.32 \times 10^7 \text{ W/m}^2) \pi (2.00 \times 10^{-3} \text{ m})^2 = \boxed{1.05 \text{ kW}}$$