

# Energy Transport + The Poynting Vector 23.1

We have already encountered potential energy stored in electric + magnetic fields.

For EM wave, how much energy is delivered?

instantaneous value →

$$\vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B}$$

Poynting Vector

→ energy - per area per time  
→ in direction of travel of EM wave

or  
Power per area

Magnitude always

$$S = \frac{1}{\mu_0} E B$$

since  $\vec{E} \perp \vec{B}$  for EM waves

- since  $c = E/B$ , we also have

$$S = \frac{E^2}{2\mu_0 c} = \frac{c B^2}{2\mu_0}$$

## Intensity of EM Wave

23.2

Generally we do not care about the rapid variation in power for oscillations of fields in EM wave.

- Calculate average quantities

$$\underline{E_{rms} = \frac{E_{max}}{\sqrt{2}}}, \quad \underline{B_{rms} = \frac{B_{max}}{\sqrt{2}}}$$

Substituting gives:

$$I(\text{Savys}) = \frac{1}{\mu_0} \left( \frac{E_{max}}{\sqrt{2}} \right) \left( \frac{B_{max}}{\sqrt{2}} \right)$$

$$\boxed{I = \frac{E_{max} B_{max}}{2\mu_0}}$$

Since  $c = E/B$ , we can also write

$$\underline{\underline{I = \frac{c B_{max}^2}{2\mu_0} = \frac{E_{max}^2}{2\mu_0 c}}}$$

## Energy Density

22.3

We would like to compare the energy per volume in the  $E$  and  $B$  field components.

$$u_E = \frac{1}{2} \epsilon_0 E^2 \quad (\text{from our study of capacitors})$$

$$u_B = \frac{1}{2} \frac{B^2}{\mu_0} \quad (\text{inductors})$$
$$= \frac{(E/c)^2}{2\mu_0} = \underline{\underline{\frac{1}{2} \epsilon_0 E^2}}$$

So  $|u_E| = |u_B|$  and the total is

$$u = u_E + u_B = \epsilon_0 E^2 = B^2/\mu_0$$

The average is

$$\boxed{u_{\text{avg}} = \epsilon_0 E_{\text{max}}^2/2}$$
$$= B_{\text{max}}^2/2\mu_0$$

$$\therefore \underline{\underline{I = c u_{\text{avg}}}}$$

# Variation of Intensity with Distance 23,4

How does power get distributed at a distance from source of EM waves?

Consider point source:

- have a sphere at radius,  $r$

- all energy in wave @ source

→ must be deposited on surface of sphere

- energy is 'spread out' as  $r$  gets larger



$$I = \frac{\text{power delivered}}{\text{area of sphere}} = \frac{P_s}{4\pi r^2}$$

This is why lights get dim the further away you are.

Example:

Calculate the energy density in sunlight. Its intensity is  $1 \text{ kW/m}^2$ .

$$\underline{\underline{u_{\text{avg}} = I/c}}$$

$$= \frac{10^3 \text{ W/m}^2}{3 \times 10^8 \text{ m/s}}$$

$$= \frac{1}{3} \times 10^{-5} \frac{\text{W/m}^2}{\text{m/s}}$$

$$= \boxed{3.3 \times 10^{-6} \text{ J/m}^3}$$

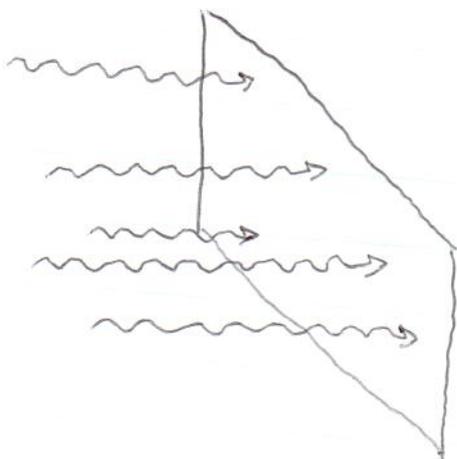
# Radiation Pressure

23.5

Light has energy

Consider surface  $\perp$  to direction of travel of wave

there is momentum associated with light



Total energy transferred to surface in time interval,  $\Delta t$ , is  $\Delta U$ .

If complete absorption:  
- momentum transferred by radiation

$$\Delta U = c \Delta p$$

$$\therefore \boxed{\Delta p = \frac{\Delta U}{c}}$$

## Radiation Pressure (cont.) 23.6

To calculate the force on the area, consider intensity

$$I = \frac{\text{energy/time}}{\text{area}}$$

$$\hookrightarrow \Delta U = I A \Delta t$$

Since force =  $dp/dt$ ,

$$\underline{\underline{F}} = \frac{I A \Delta t}{c \Delta t} = \underline{\underline{\frac{I A}{c}}}$$

the pressure is then

$$\boxed{p_r = F/A = I/c}$$

For total reflection,

$$\boxed{\Delta p = 2 \Delta U/c}$$

and

$$\boxed{p_r = 2I/c}$$

Example:

An intense light delivers  $25 \text{ W/m}^2$  to a facing wall.

What pressure is exerted on the wall?

$$\begin{aligned}
 \underline{p_r} &= S_{\text{avg}}/c \\
 &= \frac{2.5 \times 10^2 \text{ W}}{3 \times 10^8 \text{ m}^2 \cdot \text{m/s}} \\
 &= \frac{2.5}{3} \times 10^{-7} \frac{\text{N}}{\text{m}^2} \\
 &\approx \boxed{8 \times 10^{-8} \text{ N/m}^2}
 \end{aligned}$$