

# Physics 3305 - Home Work - Chapter 4

## Exercise 4.17

$$\lambda = \frac{h}{p} = \frac{h}{m_e c} = \frac{6.63 \cdot 10^{-34} \text{ J}\cdot\text{s}}{(9.11 \cdot 10^{-31} \text{ kg})(3 \cdot 10^8 \text{ m/s})} = 2.43 \cdot 10^{-12} \text{ m}$$

## Exercise 4.18

$$E = \frac{p^2}{2m} = \frac{3}{2} k_B T \quad \text{with } p = \frac{h}{\lambda}$$
$$\frac{(h/\lambda)^2}{2m} = \frac{3}{2} k_B T \Rightarrow \frac{h^2}{\lambda^2 m} = 3 k_B T \Rightarrow$$

$$\boxed{\lambda = \frac{h}{\sqrt{3mk_B T}}}$$

## Exercise 4.19

$$\lambda = \frac{h}{\sqrt{3mk_B T}}, \quad T(^{\circ}\text{K}) = T(^{\circ}\text{C}) + 273$$

a) For an electron:

$$\lambda = \frac{6.63 \cdot 10^{-34} \text{ J}\cdot\text{s}}{\sqrt{3(9.11 \cdot 10^{-31}) (1.38 \cdot 10^{-23} \text{ J/K}) (295 \text{ K})}} = 6.29 \text{ nm}$$

b) For a proton

$$\lambda = \frac{6.63 \cdot 10^{-34} \text{ J}\cdot\text{s}}{\sqrt{3(1.67 \cdot 10^{-27} \text{ kg}) (1.38 \cdot 10^{-23} \text{ J/K}) (295 \text{ K})}} = 0.147 \text{ nm}$$

Although the proton's speed would be smaller, its mass is so much larger that its momentum is large thus smaller wave-length.

In situations in which dimensions are smaller or comparable to  $\sim 6.29 \text{ nm}$  the electron will exhibit its wave nature.

At the same(T) dimensions would have to be smaller by a factor (42) for a proton to exhibit a wave nature.

### Exercise 2.2

If the maximum nonrelativistic speed is taken to be  $c/10$ , the wavelength would be

$$\lambda = \frac{h}{p} = \frac{6.63 \cdot 10^{-34} \text{ J}\cdot\text{s}}{(9.11 \cdot 10^{-31} \text{ kg})(3 \cdot 10^7 \text{ m/s})} = 2.43 \cdot 10^{-11} \text{ m}$$

Wavelengths this small or smaller would imply relativistic motion. For the accelerating potential. Use formula example 4.3

$$V = \frac{h^2}{2mq\lambda^2} = \frac{(6.63 \cdot 10^{-34} \text{ J}\cdot\text{s})^2}{2(9.11 \cdot 10^{-31} \text{ kg})(1.6 \cdot 10^{-19} \text{ C})(2.43 \cdot 10^{-11} \text{ m})^2}$$

$$\boxed{V = 2500 \text{ V}}$$

### Exercise 2.4

Find the speed of the moon  $F = ma$   $a = \frac{v^2}{R}$

$$F = G \frac{m_{\text{earth}} m_{\text{moon}}}{r^2} = m_{\text{moon}} \frac{v^2}{r}$$

$$\Rightarrow v = \sqrt{\frac{G m_{\text{earth}}}{r}} = \sqrt{\frac{(6.67 \cdot 10^{-11} \text{ N}\cdot\text{m}^2/\text{kg}^2)(5.98 \cdot 10^{24} \text{ kg})}{3.84 \cdot 10^8 \text{ m}}}$$

$$\boxed{v = 1.02 \cdot 10^3 \text{ m/s}}$$

$$\text{Thus } \lambda = \frac{h}{p} = \frac{h}{mv} = \frac{6.63 \cdot 10^{-34} \text{ J}\cdot\text{s}}{(7.35 \cdot 10^{22} \text{ kg})(1.02 \cdot 10^3 \text{ m/s})} = 8.85 \cdot 10^{-60} \text{ m}$$

This is much smaller than the dimensions of the region in which it moves. In fact, ~~it is~~ it is smaller than the atomic nucleus.

The moon certainly orbits as a classical particle

### Example 4.41

Heisenberg principle:  $\Delta x \Delta p \gg \frac{\hbar}{2}$

$$\Delta x = 1 \mu\text{m} = 10^{-6} \text{ m}, m = 0.145 \text{ kg}, \hbar = \frac{h}{2\pi} = 1.055 \cdot 10^{-34} \text{ J}\cdot\text{s}$$

$$\Delta x \Delta p \gg \frac{\hbar}{2} \Rightarrow (10^{-6} \text{ m})(0.145 \text{ kg}) \Delta v \gg \frac{1}{2}(1.055 \cdot 10^{-34} \text{ J}\cdot\text{s})$$

$$\Delta v \gg 3.6 \cdot 10^{-28} \text{ m/s}$$

It is theoretically impossible to say whether it might not be moving at  $\sim 10^{-22} \text{ m/s}$ .

### Example 4.43

$$\Delta x \Delta p \gg \frac{1}{2} \hbar \rightarrow (5 \cdot 10^{-15} \text{ m})(1.67 \cdot 10^{-27} \text{ kg}) \Delta v \gg \frac{1}{2}(1.055 \cdot 10^{-34} \text{ J}\cdot\text{s})$$

$$\Rightarrow \Delta v \gg 6.3 \cdot 10^6 \text{ m/s}$$

$$\text{its KE} = \frac{1}{2} (1.67 \cdot 10^{-27} \text{ kg}) (6.3 \cdot 10^6 \text{ m/s})^2 = 0.2 \text{ MeV}$$

### Exercise 4.48

$$\Delta E = 150 \text{ MeV} = 150 \cdot 10^6 \times 1.6 \cdot 10^{-19} \text{ J}$$

$$\Delta t \Delta E \gg \frac{\hbar}{2}$$

$$\Delta t \gg \frac{\hbar/2}{\Delta E}$$

$$\Delta t \gg \frac{\frac{1}{2}(1.055 \cdot 10^{-34} \text{ J}\cdot\text{s})}{150 \cdot 10^{-13} \text{ J}}$$

$$\Delta t \gg 2.2 \cdot 10^{-24} \text{ s}$$

### Exercise 4.62

Eq. 4-27

$$\Delta\omega \Delta t \gg \frac{1}{2} \Rightarrow \Delta\omega \gg \frac{1}{2\Delta t}$$

$$\omega = 2\pi f$$

Angular  
frequency

$$\Delta f = \frac{1}{2\pi} \Delta\omega$$

$$, \Delta t = 1 \text{ ns} = 10^{-9} \text{ s}$$

$$\Delta f = \frac{1}{2\pi} \Delta\omega \approx \frac{1}{2\pi} \left( \frac{1}{2\Delta t} \right) = \frac{1}{4\pi} \left( \frac{1}{10^{-9}} \right)$$

$$\Delta f = 7.96 \times 10^7 \text{ Hz}$$

(C) For the 1060 nm laser

$$f = \frac{c}{\lambda} = \frac{3 \times 10^8 \text{ m/s}}{1060 \times 10^{-9} \text{ m}} = 2.83 \times 10^{14} \text{ Hz}$$

The relative uncertainty is significant for 100 MHz radiowave, but is very small compared to the frequency of the light.

### Exercise 4.63

$$\Delta x = 0.3 \mu\text{m}$$

$$\Delta k \approx \frac{1}{2\Delta x} = \frac{1}{2 \times 0.3 \times 10^{-6}} = 1.67 \times 10^6 \text{ m}^{-1}$$

$$k = \frac{2\pi}{\lambda} \quad (\text{wave vector}) \Rightarrow \Delta k = \frac{2\pi}{\lambda^2} \Delta\lambda =$$

$$\Delta\lambda = \frac{\lambda^2 \Delta k}{2\pi} = \frac{(6 \times 10^{-7} \text{ m})^2 \times 1.67 \times 10^6}{2\pi}$$

$$\Delta\lambda = 95 \text{ nm}$$

A 1 femtosecond pulse of 600 nm light is not just 600 nm light