



**CTEQ School on
QCD Analysis and Electroweak Phenomenology**

LECTURE 4

Introduction to the Parton Model and Perturbative QCD

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University of Pittsburgh, PA

18-28 July 2017

We already studies

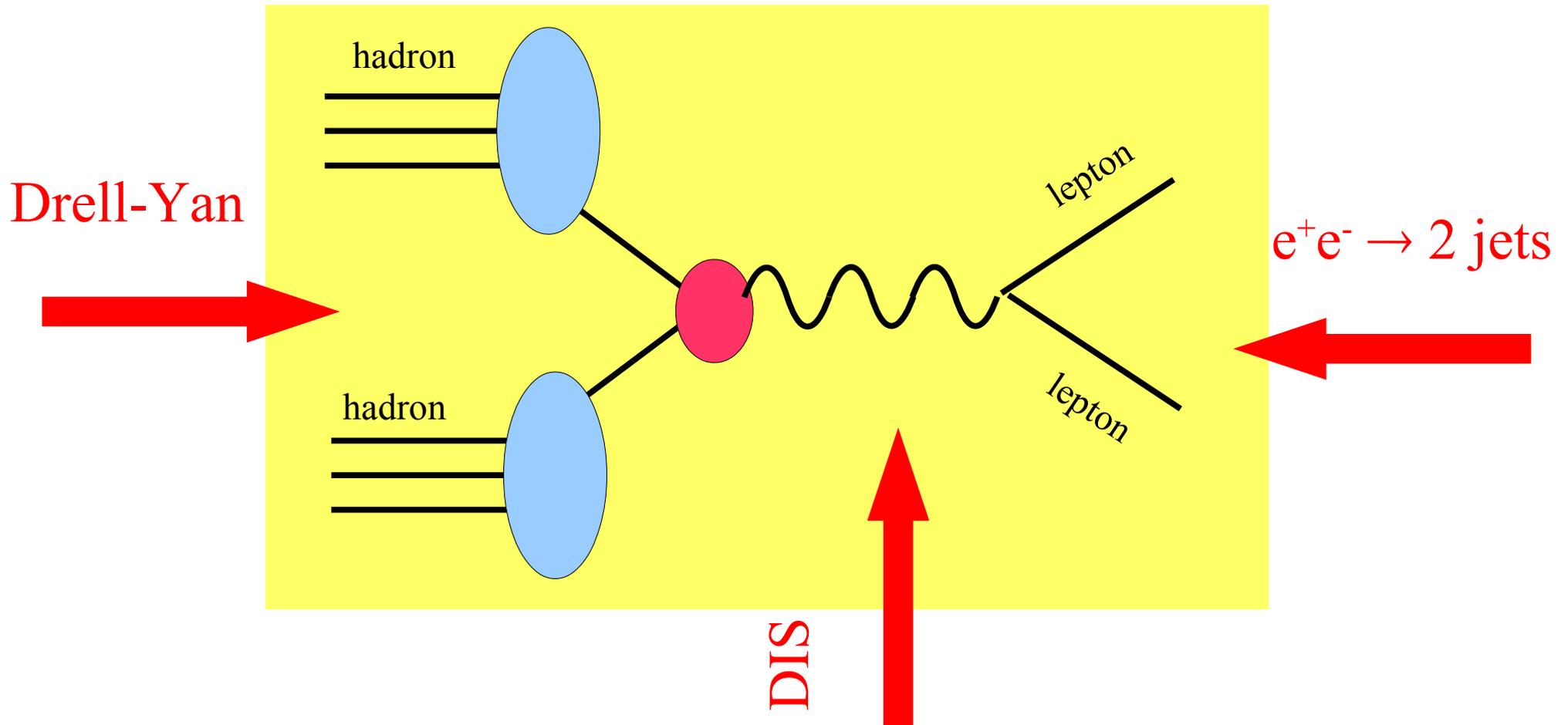
DIS

Now we consider

Drell-Yan Process

$$e^+e^-$$

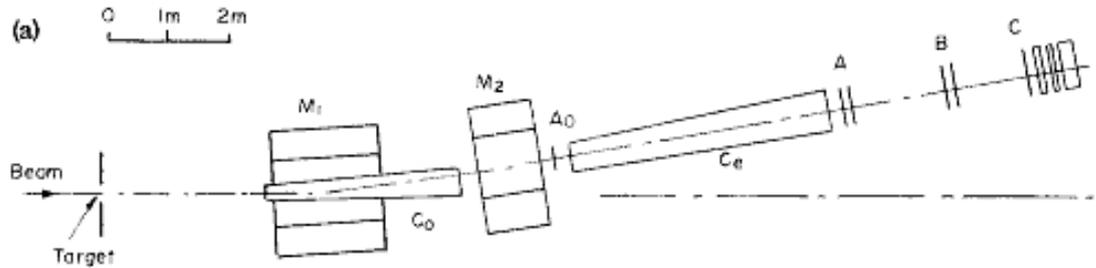
Important for Tevatron and LHC



Drell-Yan and e^+e^- have an interesting historical relation

The Process: $p + Be \rightarrow e^+ e^- X$

very narrow width
 \Rightarrow long lifetime



at BNL AGS

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PHYSICAL REVIEW LETTERS

2 DECEMBER 1974

Experimental Observation of a Heavy Particle J/ψ

J. J. Aubert, U. Becker, P. J. Biggs, J. Burger, M. Chen, G. Everhart, P. Goldhagen, J. Leong, T. McCorriston, T. G. Rhoades, M. Rohde, Samuel C. C. Ting, and Sau Lan Tsou
Laboratory for Nuclear Science and Department of Physics, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139

and

Y. Y. Lee

Brookhaven National Laboratory, Upton, New York 11973

(Received 12 November 1974)

We report the observation of a heavy particle J , with mass $m = 3.1$ GeV and width approximately zero. The observation was made from the reaction $p + Be \rightarrow e^+ + e^- + X$ by measuring the e^+e^- mass spectrum with a precise pair spectrometer at the Brookhaven National Laboratory's 30-GeV alternating-gradient synchrotron.

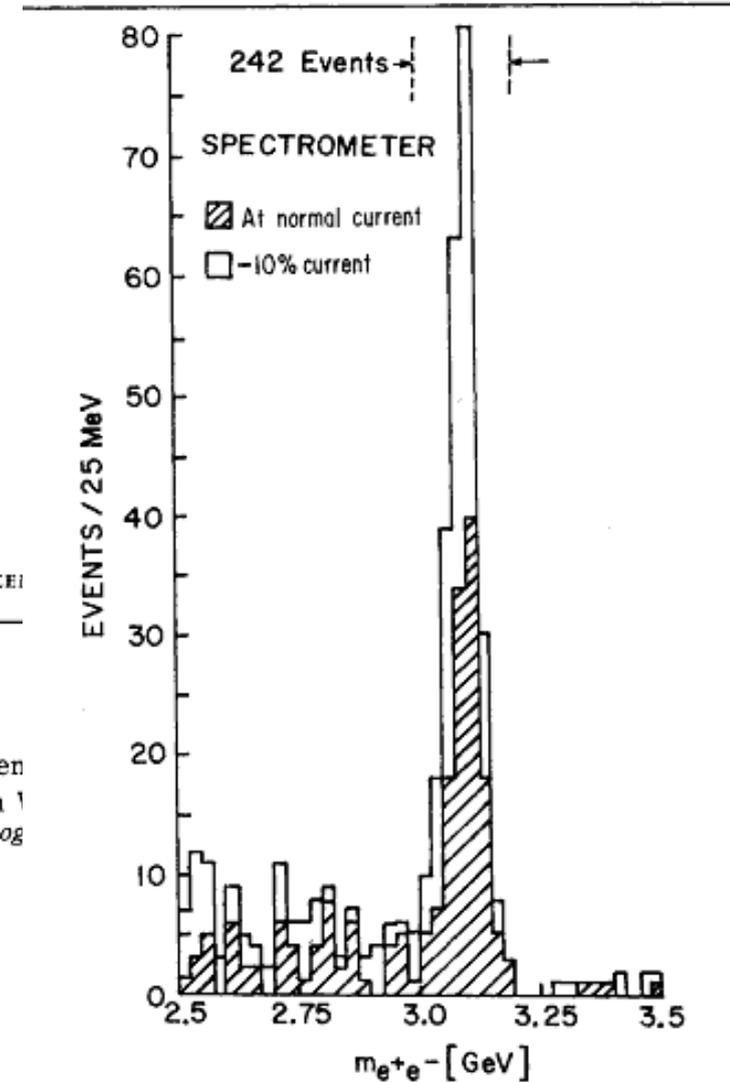


FIG. 2. Mass spectrum showing the existence of J/ψ . Results from two spectrometer settings are plotted showing that the peak is independent of spectrometer currents. The run at reduced current was taken two months later than the normal run.

This experiment is part of a large program to

daily with a thin Al foil. The beam spot

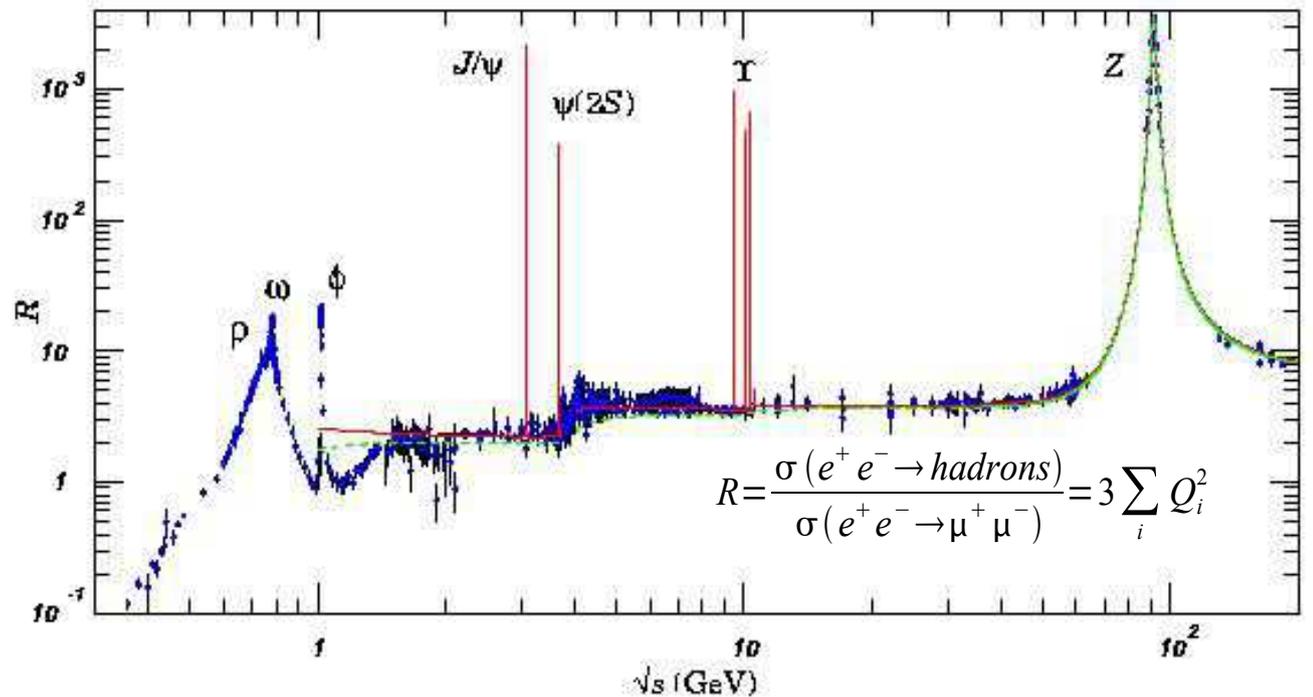
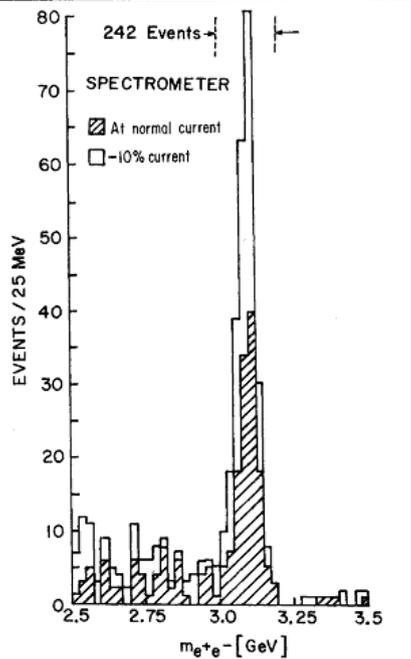
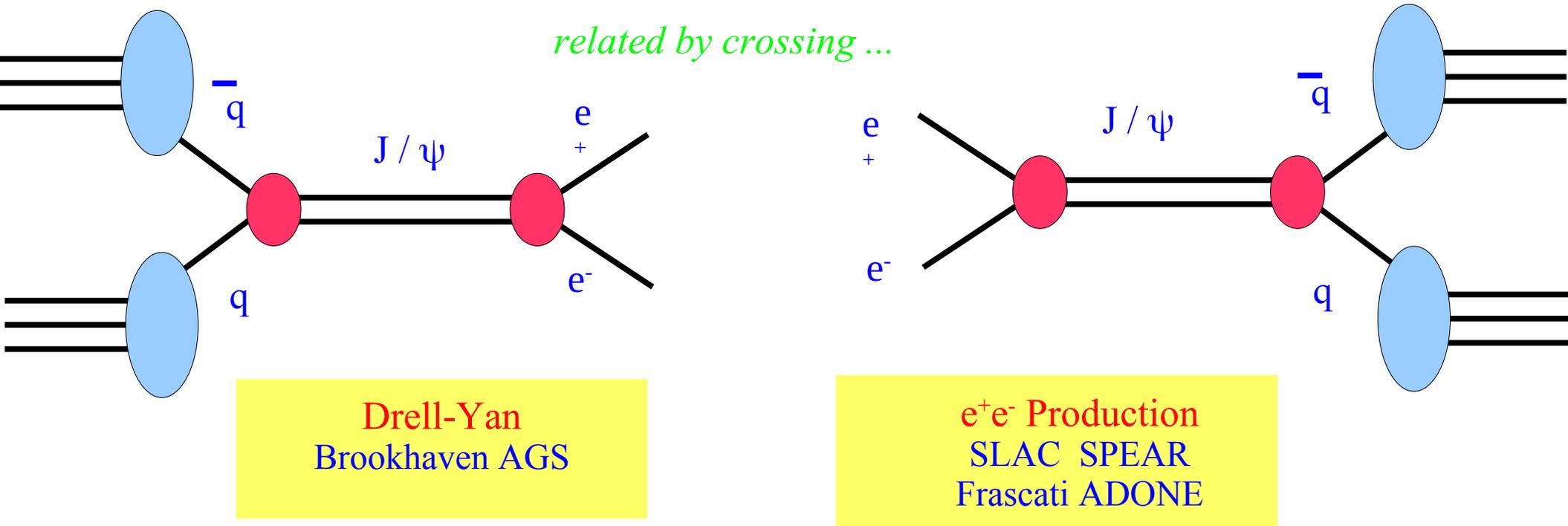
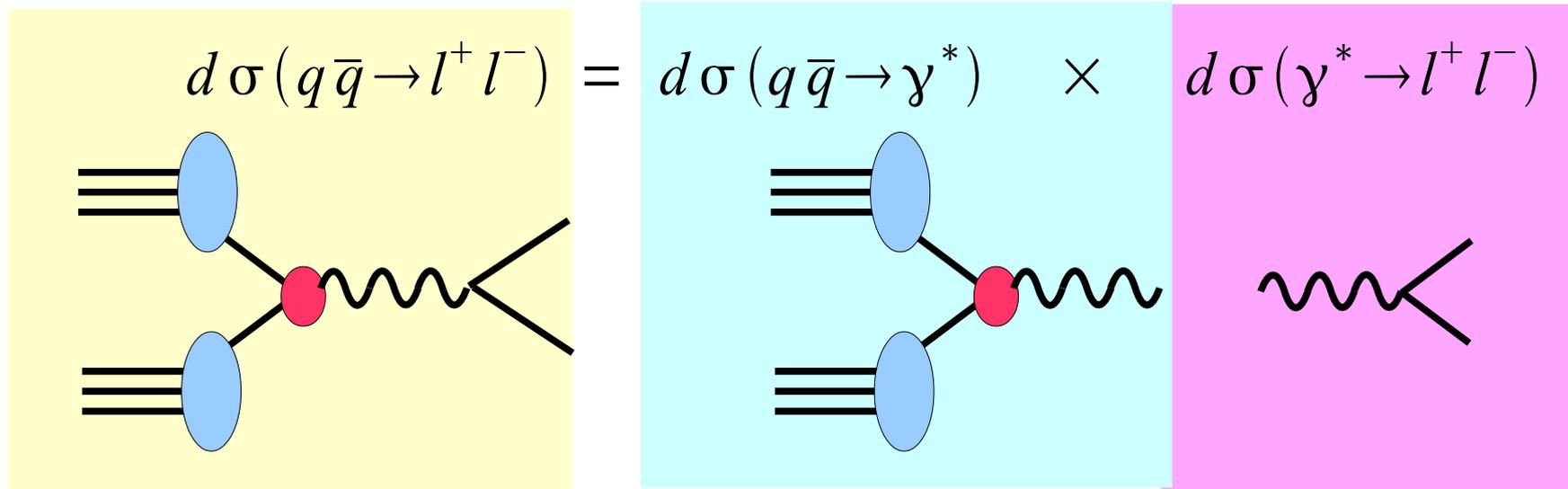


FIG. 2. Mass spectrum showing the existence of J . Results from two spectrometer settings are plotted showing that the peak is independent of spectrometer currents. The run at reduced current was taken two months later than the normal run.

We'll look at Drell-Yan

Specifically W/Z production

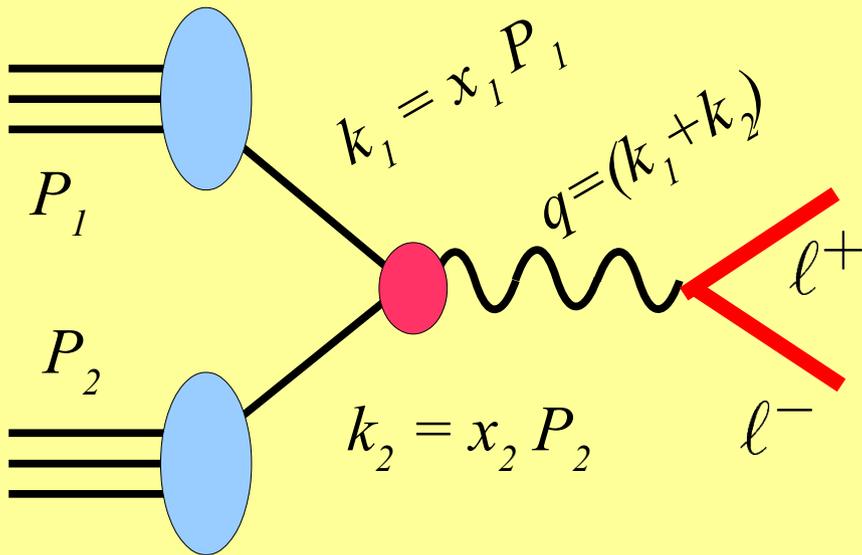
Schematically:



For example:

$$\frac{d\sigma}{dQ^2 d\hat{t}}(q\bar{q} \rightarrow l^+ l^-) = \frac{d\sigma}{d\hat{t}}(q\bar{q} \rightarrow \gamma^*) \times \frac{\alpha}{3\pi Q^2}$$

Kinematics in the hadronic CMS



$$P_1 = \frac{\sqrt{s}}{2} (1, 0, 0, +1) \quad P_1^2 = 0$$

$$P_2 = \frac{\sqrt{s}}{2} (1, 0, 0, -1) \quad P_2^2 = 0$$

$$k_1 = x_1 P_1 \quad k_1^2 = 0$$

$$k_2 = x_2 P_2 \quad k_2^2 = 0$$

$$\frac{d\sigma}{dx_1 dx_2} = \sum_{q, \bar{q}} \left\{ q(x_1) \bar{q}(x_2) + q(x_2) \bar{q}(x_1) \right\} \hat{\sigma}$$

Hadronic
cross
section

Parton
distribution
functions

Partonic
cross
section

Trade $\{x_1, x_2\}$ variables for $\{\tau, y\}$

$$x_{1,2} = \sqrt{\tau} e^{\pm y}$$

$$y = \frac{1}{2} \ln \left(\frac{x_1}{x_2} \right)$$

$$\tau = x_1 x_2$$

$$s = (P_1 + P_2)^2 = \frac{\hat{s}}{x_1 x_2} = \frac{\hat{s}}{\tau}$$

Therefore

$$\tau = x_1 x_2 = \frac{\hat{s}}{s} \equiv \frac{Q^2}{s}$$

Fractional energy² between
partonic and hadronic system

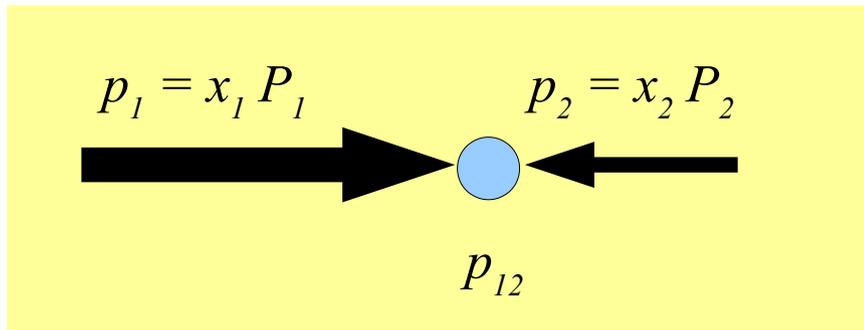
Using: $d x_1 d x_2 = d \tau d y$

$$\frac{d \sigma}{d \tau d y} = \sum_{q, \bar{q}} \left\{ q(x_1) \bar{q}(x_2) + q(x_2) \bar{q}(x_1) \right\} \hat{\sigma}$$

The rapidity is defined as:

$$y = \frac{1}{2} \ln \left\{ \frac{E_{12} + p_L}{E_{12} - p_L} \right\}$$

Partonic CMS has longitudinal momentum **w.r.t.** the hadron frame



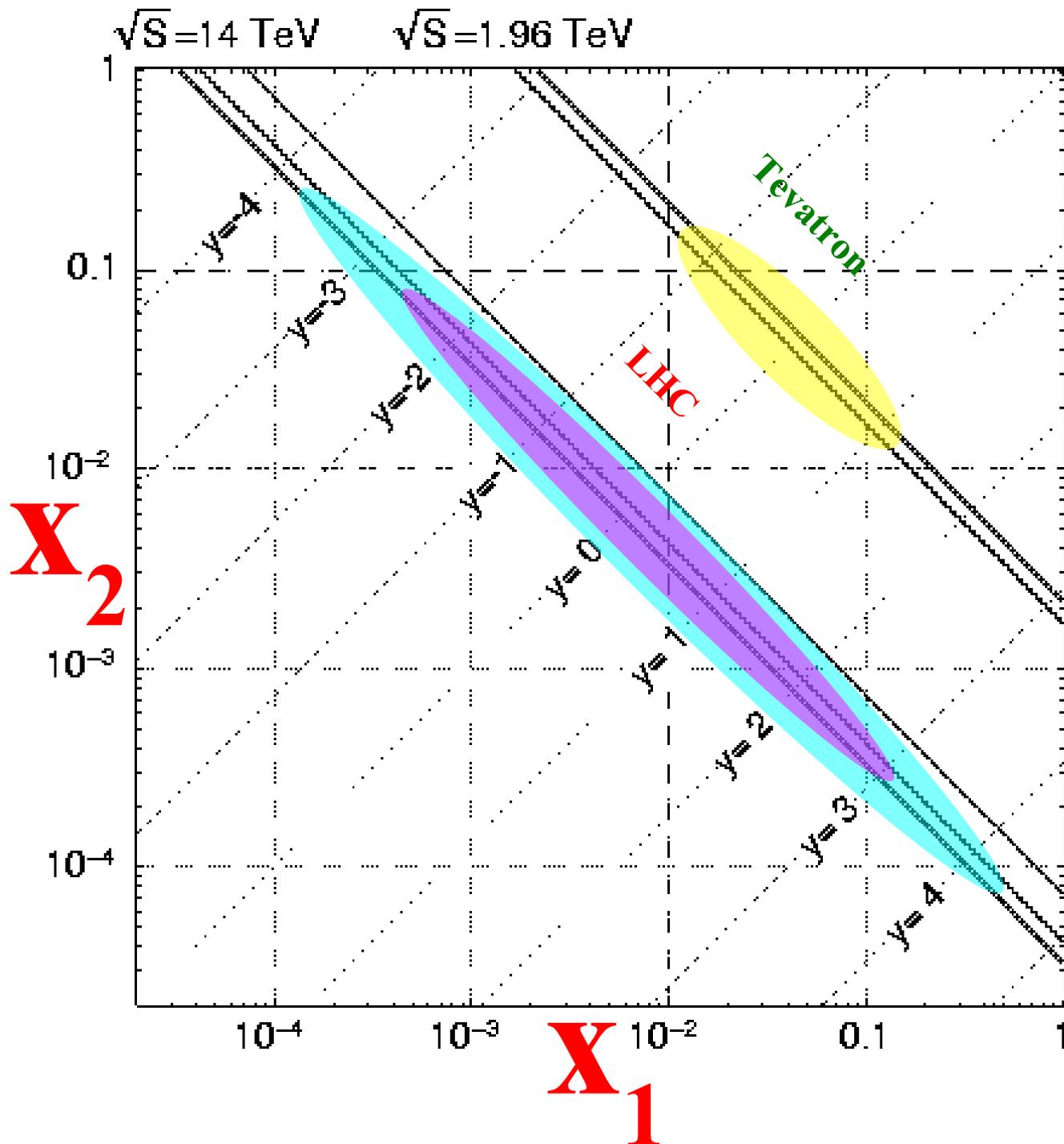
$$p_{12} = (p_1 + p_2) = (E_{12}, 0, 0, p_L)$$

$$E_{12} = \frac{\sqrt{s}}{2} (x_1 + x_2)$$

$$p_L = \frac{\sqrt{s}}{2} (x_1 - x_2) \equiv \frac{\sqrt{s}}{2} x_F$$

x_F is a measure of the longitudinal momentum

$$y = \frac{1}{2} \ln \left\{ \frac{E_{12} + p_L}{E_{12} - p_L} \right\} = \frac{1}{2} \ln \left\{ \frac{x_1}{x_2} \right\}$$



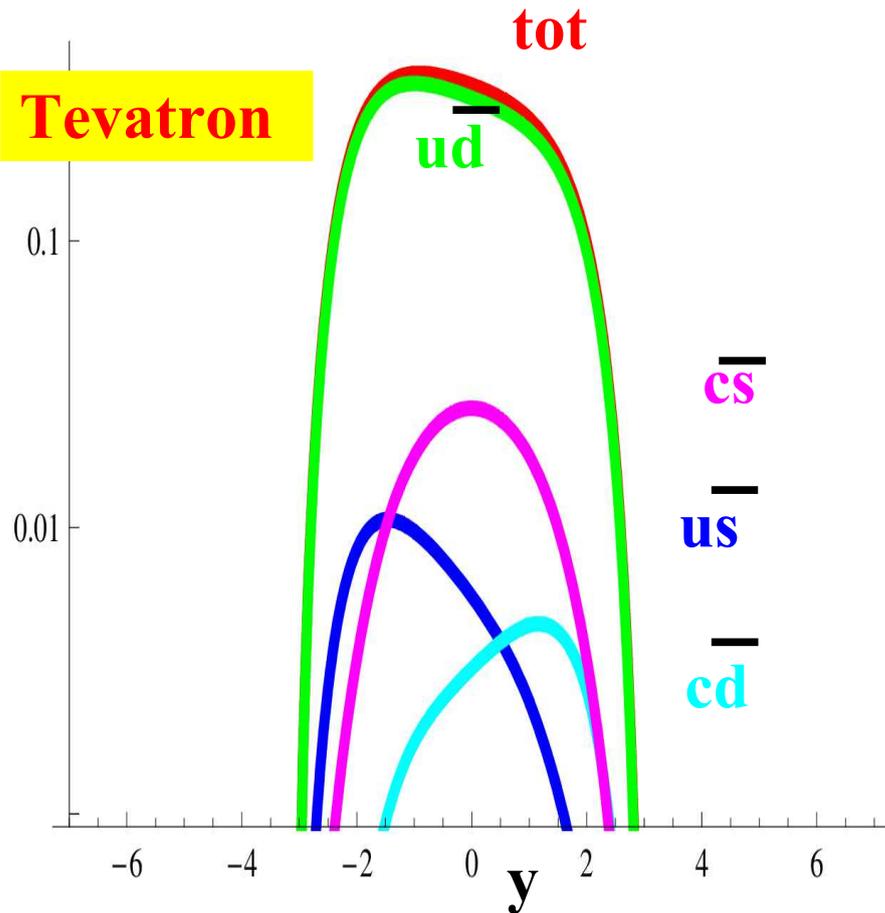
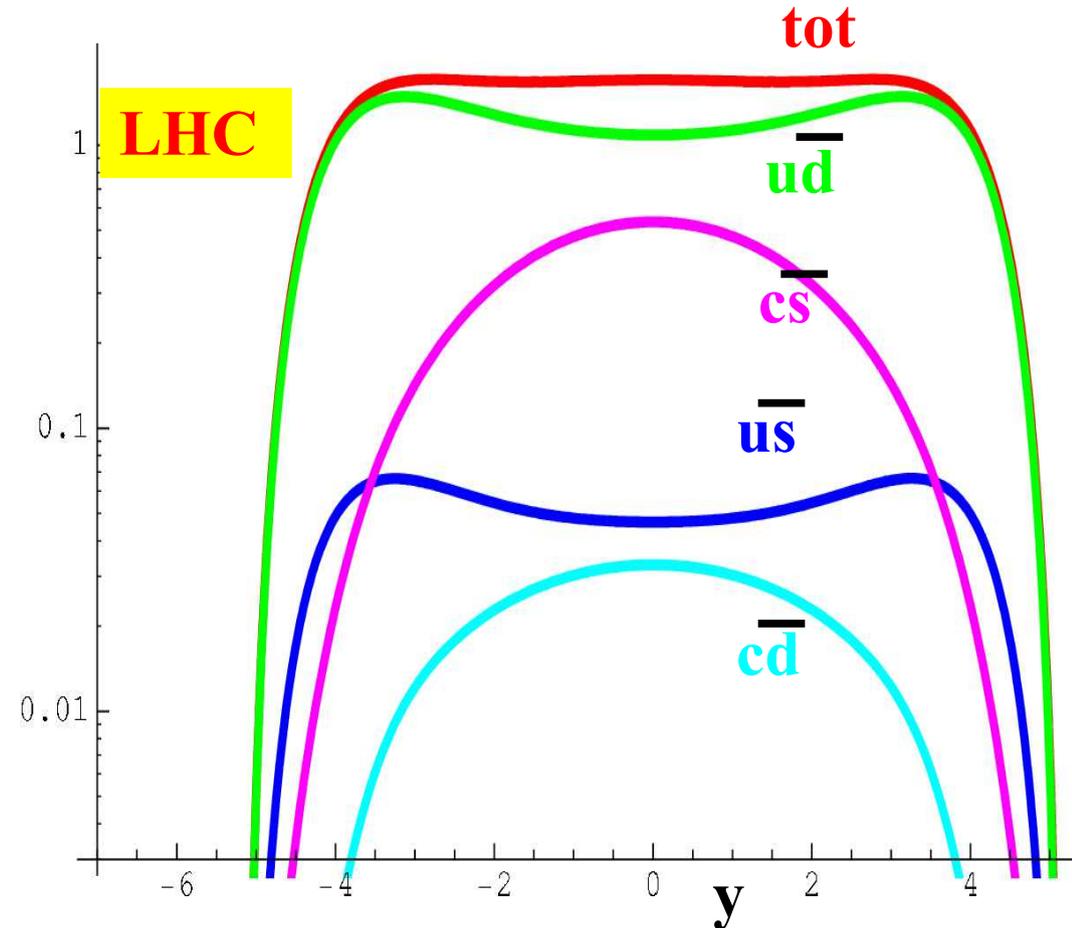
$$y = \frac{1}{2} \ln \left(\frac{x_1}{x_2} \right)$$

$$\tau = x_1 x_2$$

$$x_{1,2} = \sqrt{\tau} e^{\pm y}$$

Z
W

H
Z
W

LO W⁺ LuminositiesLO W⁺ Luminosities

$$d\sigma = \int dx_1 \int dx_2 \int d\tau \left\{ q(x_1)\bar{q}(x_2) + q(x_2)\bar{q}(x_1) \right\} \hat{\sigma} \delta\left(\tau - \frac{M^2}{S}\right)$$

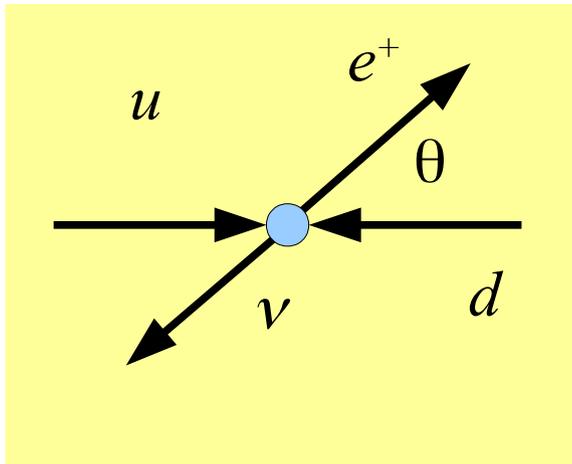
$$\frac{d\sigma}{d\tau} = \frac{dL}{d\tau} \hat{\sigma}(\tau)$$

$$\frac{dL}{d\tau} = f \otimes f$$

The CMS of the parton-parton system is moving longitudinally relative to the hadron-hadron system

How do we measure the W-boson mass?

$$u + \bar{d} \rightarrow W^+ \rightarrow e^+ \nu$$

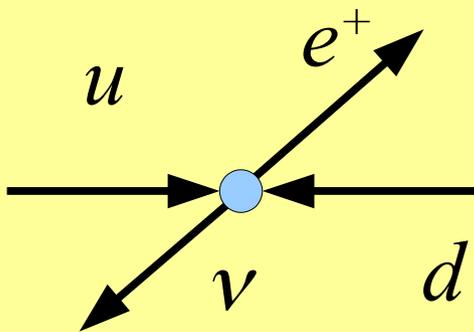


Can't measure W directly

Can't measure ν directly

Can't measure longitudinal momentum

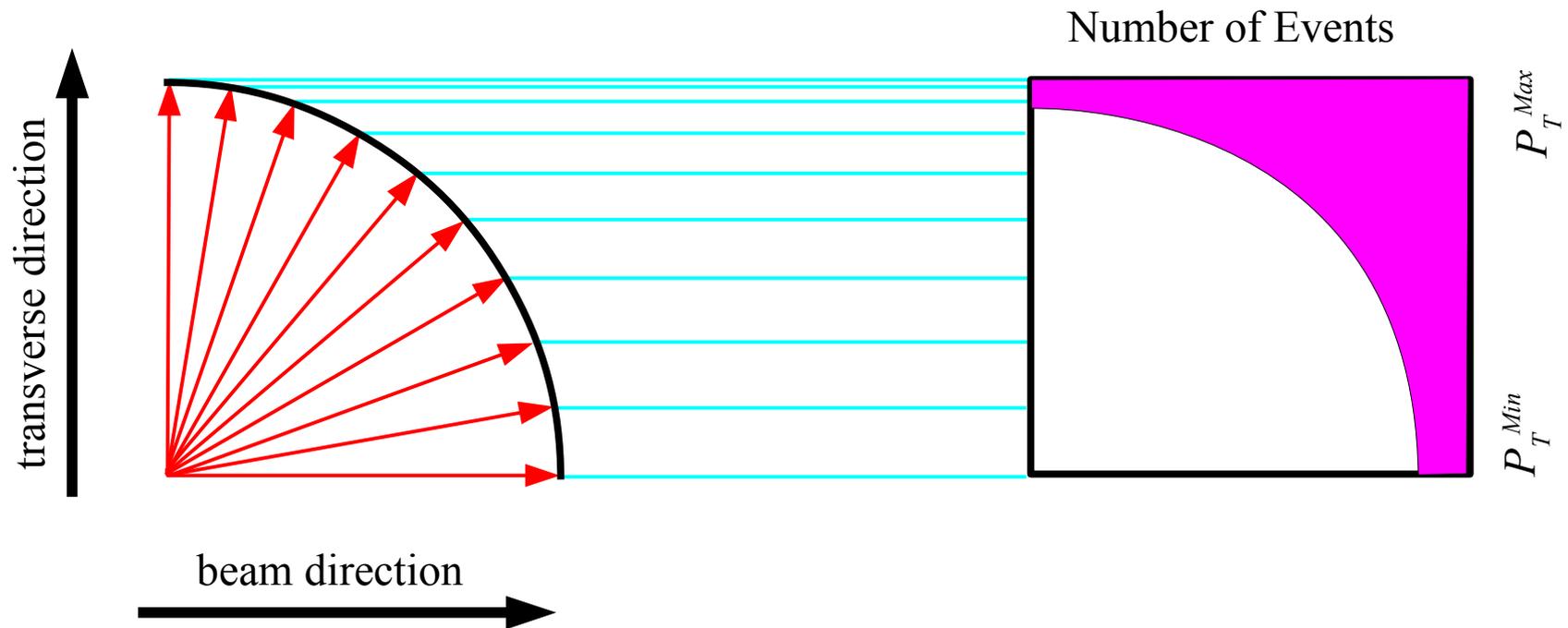
We can measure the P_T of the lepton



Suppose lepton distribution is uniform in θ

The dependence is actually $(1+\cos\theta)^2$, but we'll worry about that later

What is the distribution in P_T ?

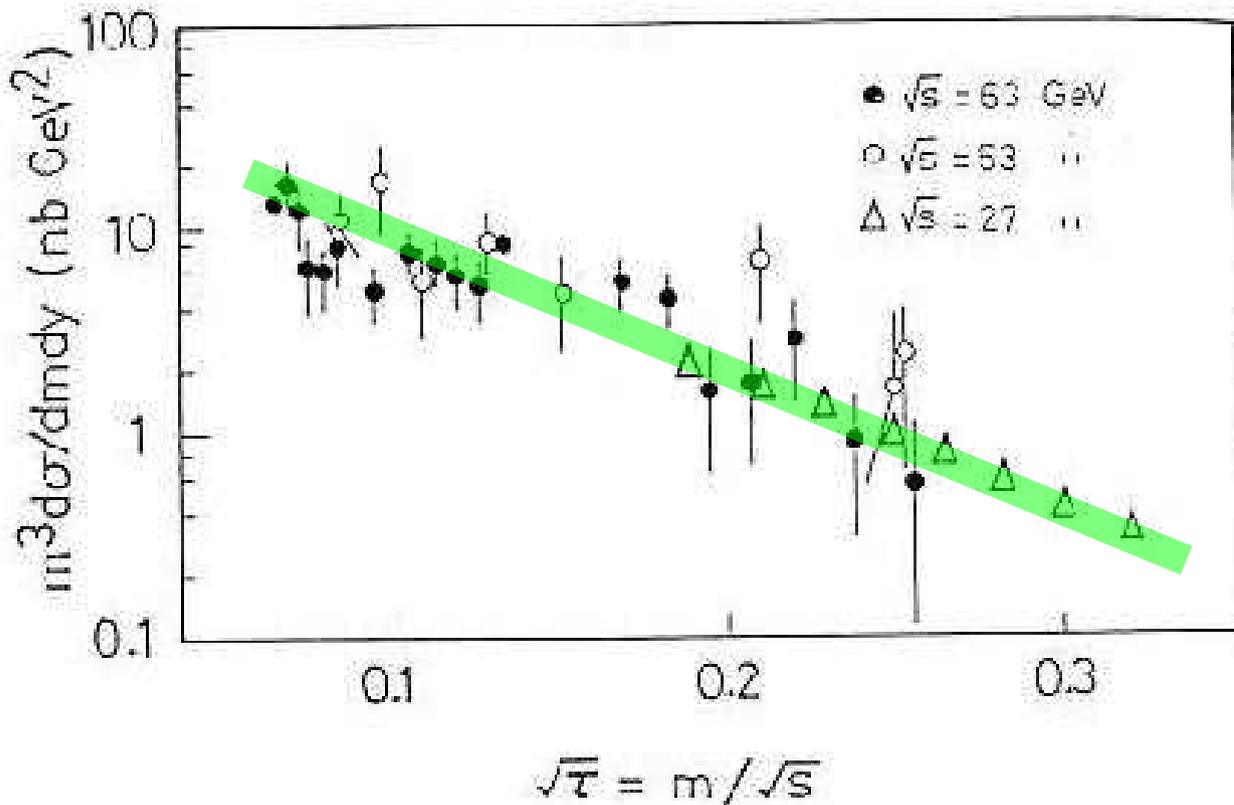


We find a peak at $P_T^{max} \approx M_W/2$

Using: $\hat{\sigma}_0 = \frac{4\pi\alpha^2}{9\hat{s}} Q_i^2$ and $\delta(Q^2 - \hat{s}) = \frac{1}{s x_1} \delta(x_2 - \frac{\tau}{x_1})$

we can write the cross section in the scaling form:

$$Q^4 \frac{d\sigma}{dQ^2} = \frac{4\pi\alpha^2}{9} \sum_{q,\bar{q}} Q_i^2 \int_{\tau}^1 \frac{dx_1}{x_1} \tau \left\{ q(x_1) \bar{q}(\tau/x_1) + \bar{q}(x_1) q(\tau/x_1) \right\}$$



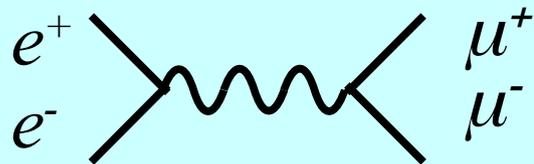
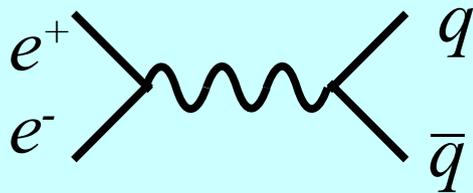
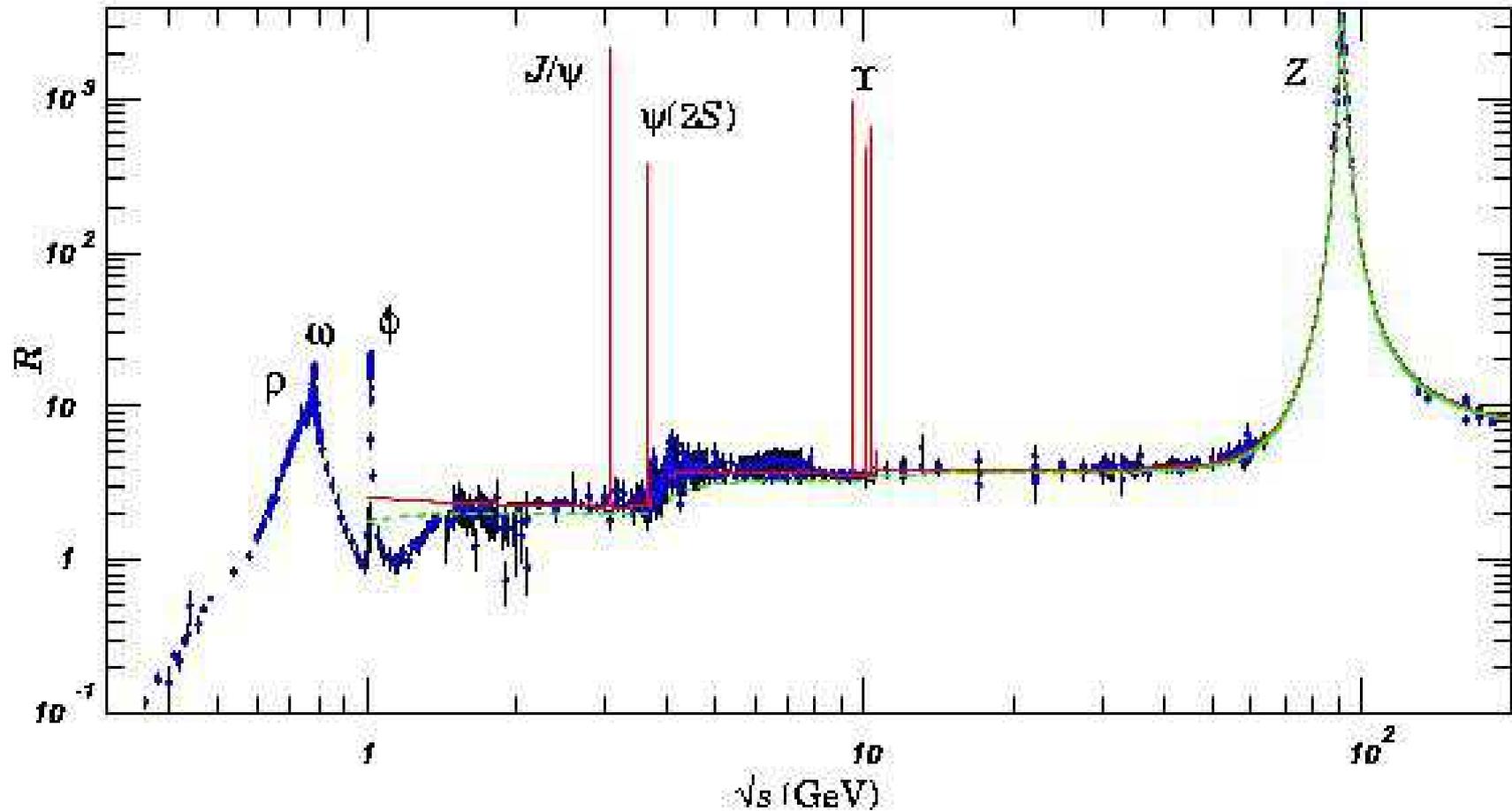
Notice the RHS is a function of only τ , not Q .

This quantity should lie on a universal scaling curve.

Cf., DIS case, & scattering of point-like constituents

e^+e^- R ratio

$$R = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)}$$



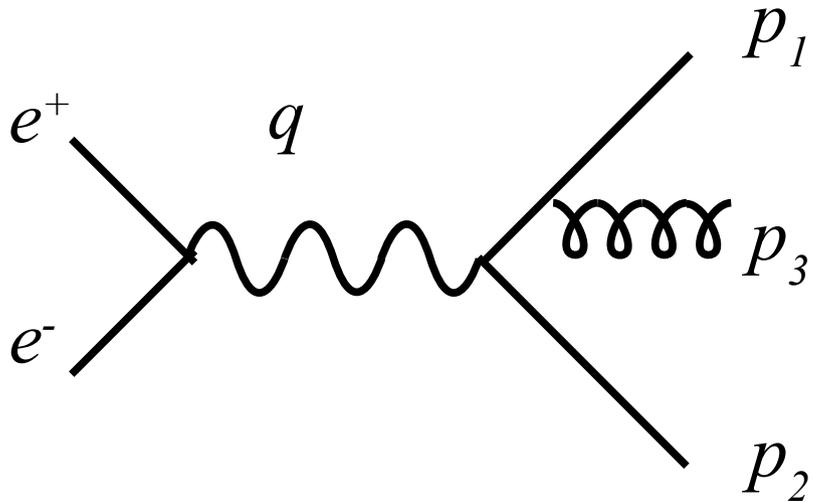
$$R = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)} = 3 \sum_i Q_i^2 \left[1 + \frac{\alpha_s}{\pi} \right]$$

3 quark colors

NLO correction

$$e^+e^-$$

NLO corrections



Define the energy fractions E_i :

$$x_i = \frac{E_i}{\sqrt{s}/2} = \frac{2p_i \cdot q}{s}$$

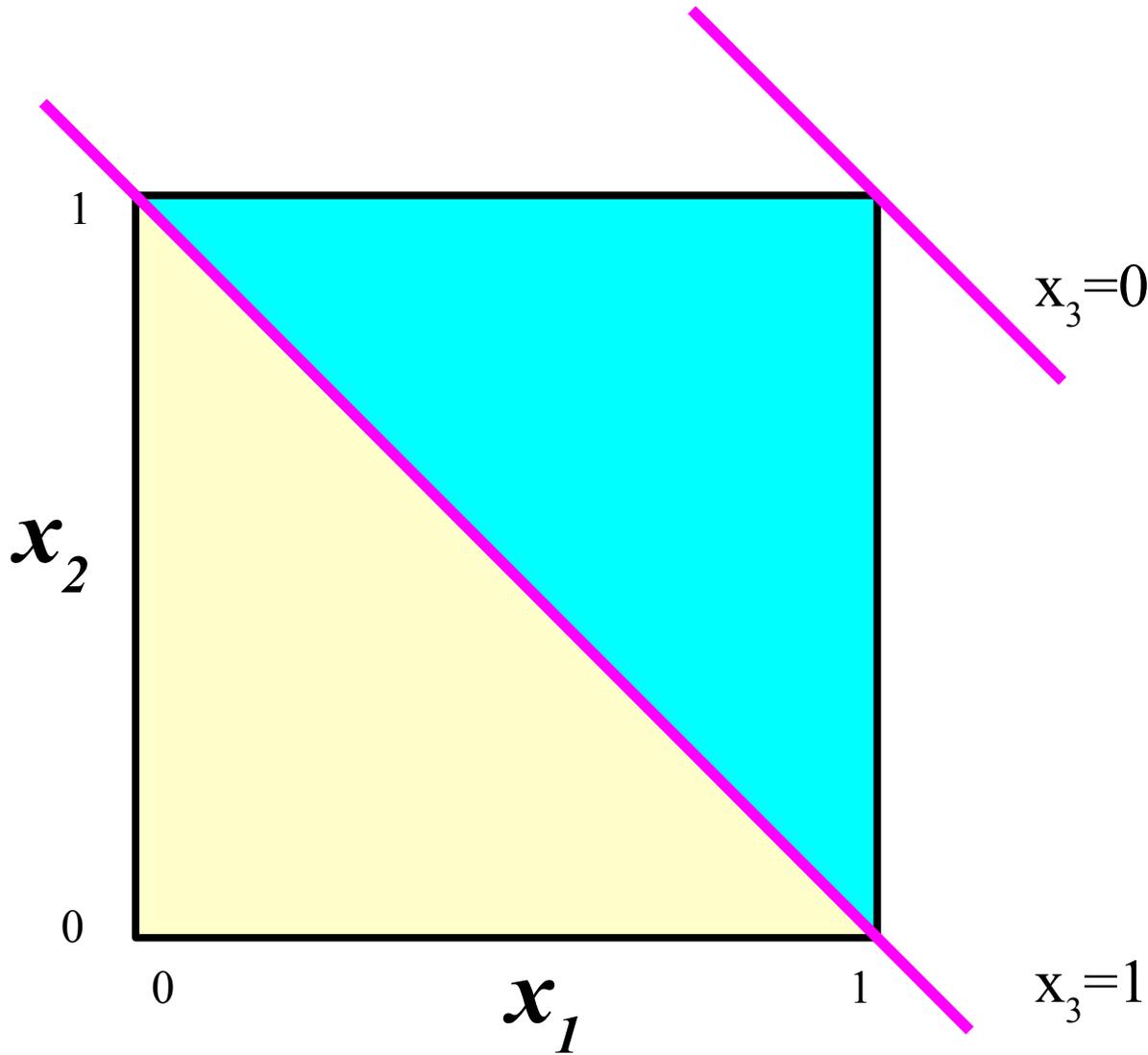
Energy Conservation:

$$\sum_i x_i = 2$$

Range of x:

$$x_i \subset [0, 1]$$

Exercise: show 3-body phase space is flat in $dx_1 dx_2$

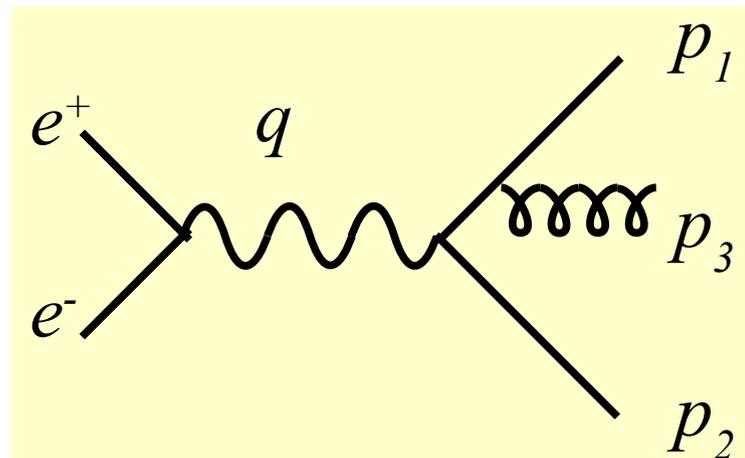


$$x_i = \frac{E_i}{\sqrt{s}/2}$$

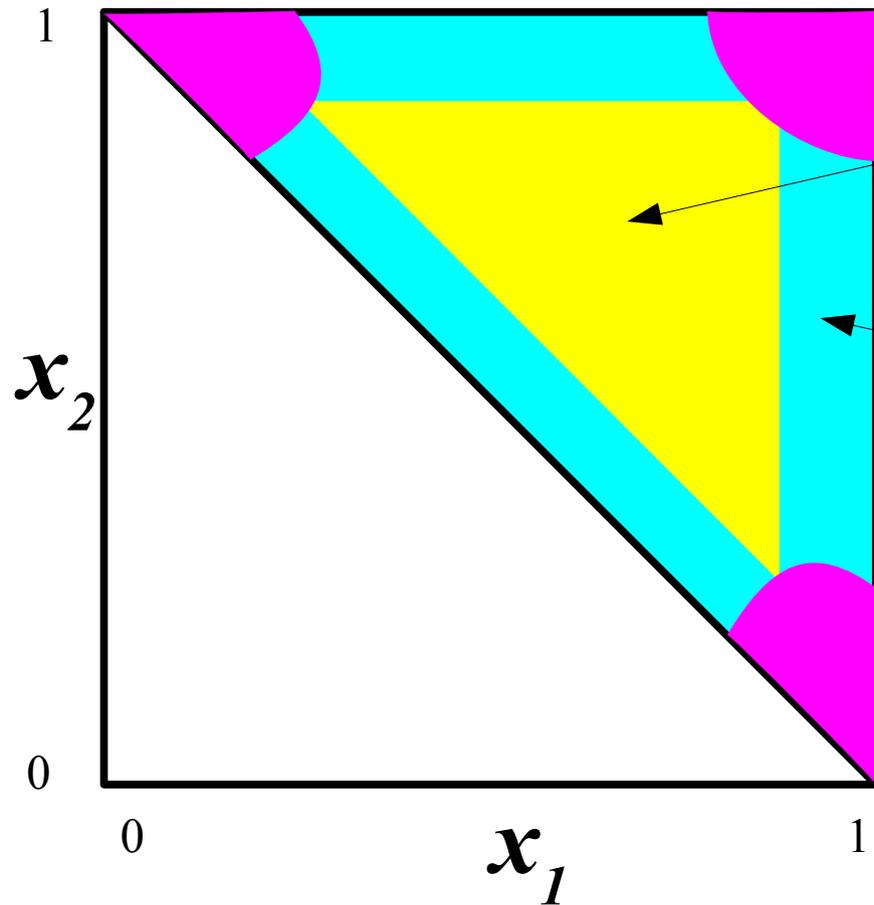
$$x_i \in [0, 1]$$

$$x_1 + x_2 + x_3 = 2$$

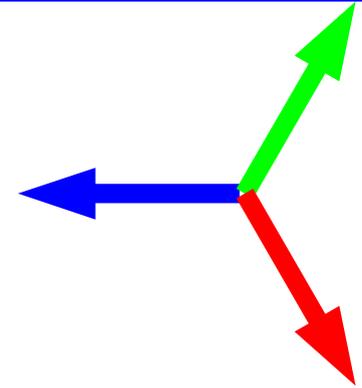
$$d\Gamma \sim dx_1 dx_2$$



$$\sigma_0 = \frac{4\pi\alpha^2}{s} \sum e_q^2 \quad C_F = 4/3$$



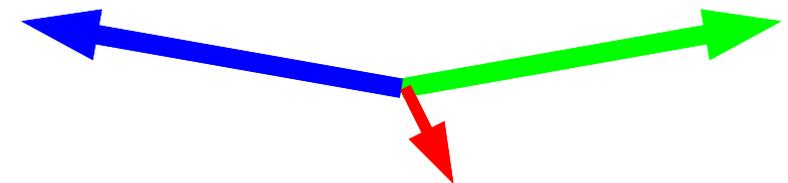
3-Jet



Collinear

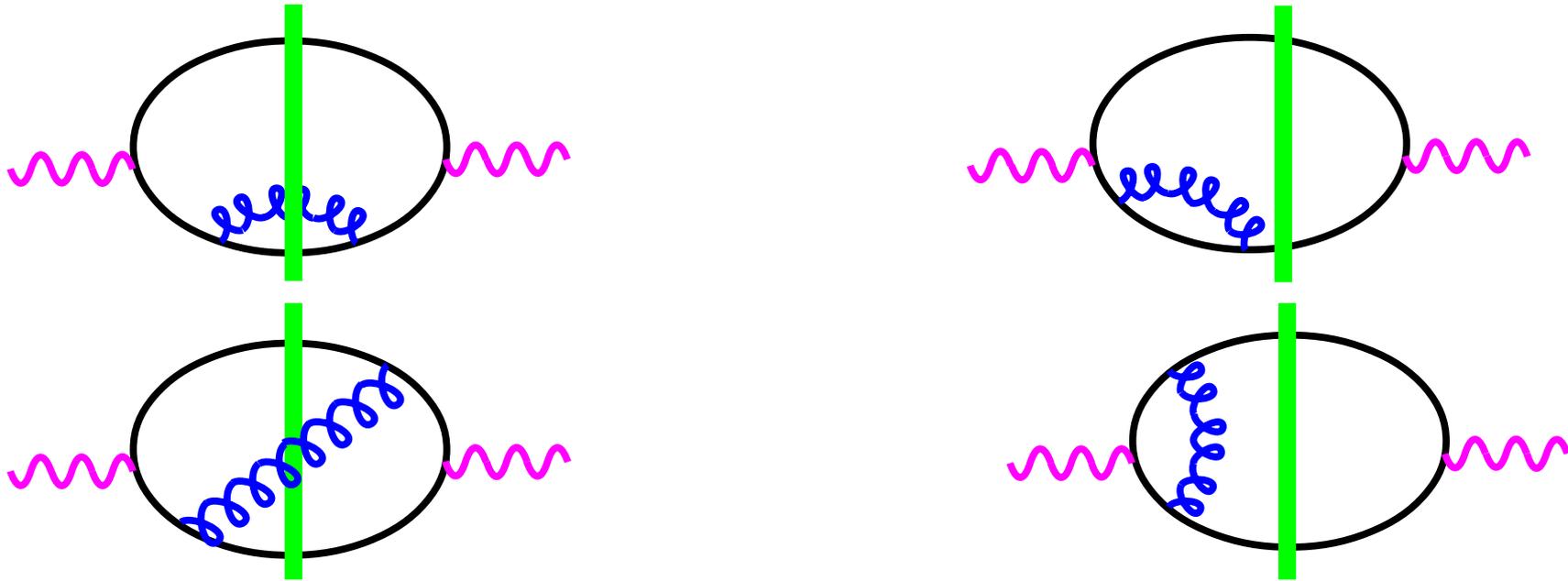


Soft



After symmetrization

$$\frac{1}{\sigma_0} \frac{d\sigma}{dx_1 dx_2} = \frac{\alpha_s}{2\pi} C_F \frac{x_1^2 + x_2^2 + x_3^2}{(1-x_1)(1-x_2)(1-x_3)}$$



$$\sigma_2^{(\epsilon)} = \sigma_0 C_F \frac{\alpha_s}{\pi} (\dots) \left[+\frac{-1}{\epsilon^2} + \frac{-3}{2\epsilon} + \frac{\pi^2}{2} - 4 \right]$$

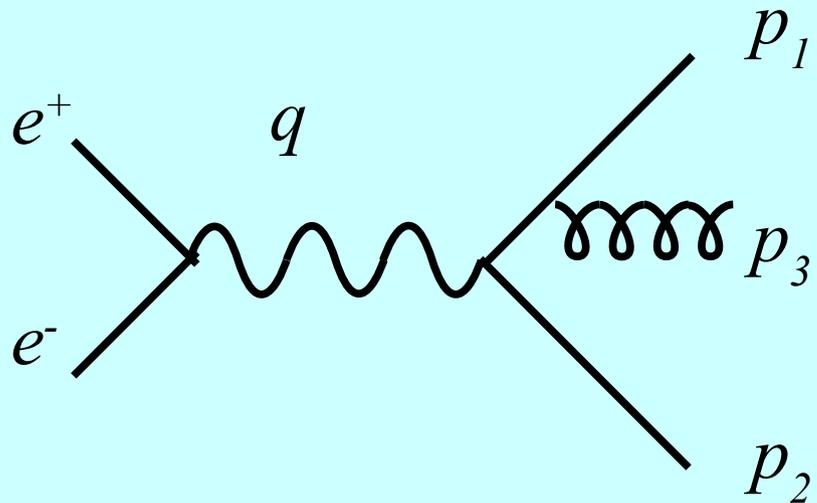
$$\sigma_3^{(\epsilon)} = \sigma_0 C_F \frac{\alpha_s}{\pi} (\dots) \left[+\frac{+1}{\epsilon^2} + \frac{+3}{2\epsilon} + \frac{-\pi^2}{2} - \frac{19}{4} \right]$$

$$\sigma_2^{(\epsilon)} + \sigma_3^{(\epsilon)} = \sigma_0 C_F \frac{\alpha_s}{\pi} (\dots) \left[0 + 0 + 0 + -\frac{35}{4} \right]$$

Same result with
gluon mass
regularization

$$e^+e^-$$

Differential Cross Sections



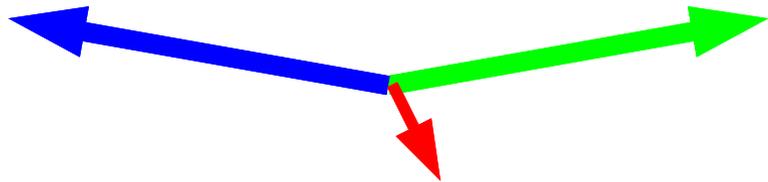
What do we do about soft and collinear singularities???

Introduce the concept of “Infrared Safe Observable”

The soft and collinear singularities will cancel
ONLY
if the physical observables are appropriately defined.

Observables must satisfy the following requirements:

Soft



if $p_s \rightarrow 0$

$$\mathcal{O}_{n+1}(p_1, \dots, p_n, p_s) \longrightarrow \mathcal{O}_n(p_1, \dots, p_n)$$

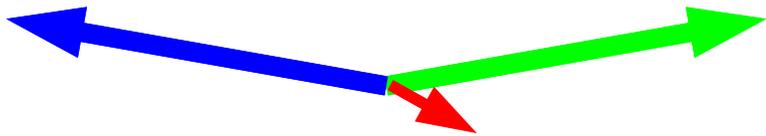
Collinear



if $p_a \parallel p_b$

$$\mathcal{O}_{n+1}(p_1, \dots, p_a, p_b, \dots, p_n) \longrightarrow \mathcal{O}_n(p_1, \dots, p_a + p_b, \dots, p_n)$$

Soft



if $p_s \rightarrow 0$



$$\mathcal{O}_{n+1}(p_1, \dots, p_n, p_s) \longrightarrow \mathcal{O}_n(p_1, \dots, p_n)$$

Collinear



if $p_a \parallel p_b$



$$\mathcal{O}_{n+1}(p_1, \dots, p_a, p_b, \dots, p_n) \longrightarrow \mathcal{O}_n(p_1, \dots, p_a + p_b, \dots, p_n)$$

Infrared Safe Observables:

Event shape distributions

Jet Cross sections

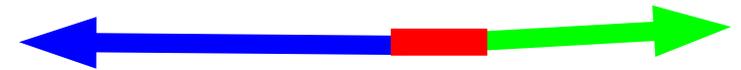
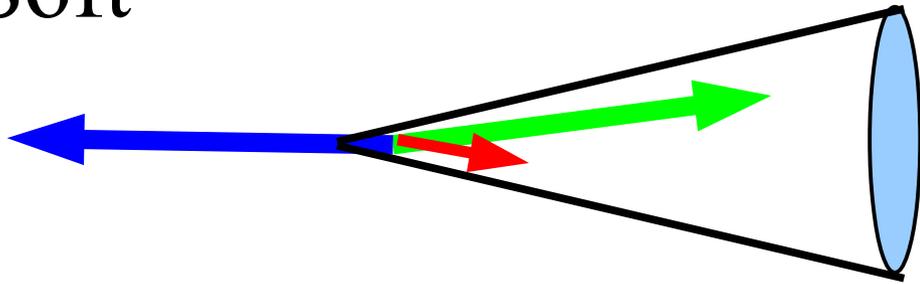
Un-Safe Infrared Observables:

Momentum of the hardest particle
(affected by collinear splitting)

100% isolated particles
(affected by soft emissions)

Particle multiplicity
(affected by both soft & collinear emissions)

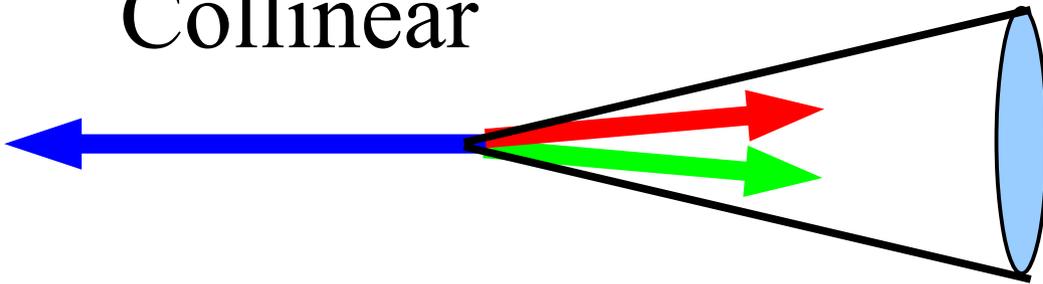
Soft



if $p_s \rightarrow 0$

$$\mathcal{O}_{n+1}(p_1, \dots, p_n, p_s) \longrightarrow \mathcal{O}_n(p_1, \dots, p_n)$$

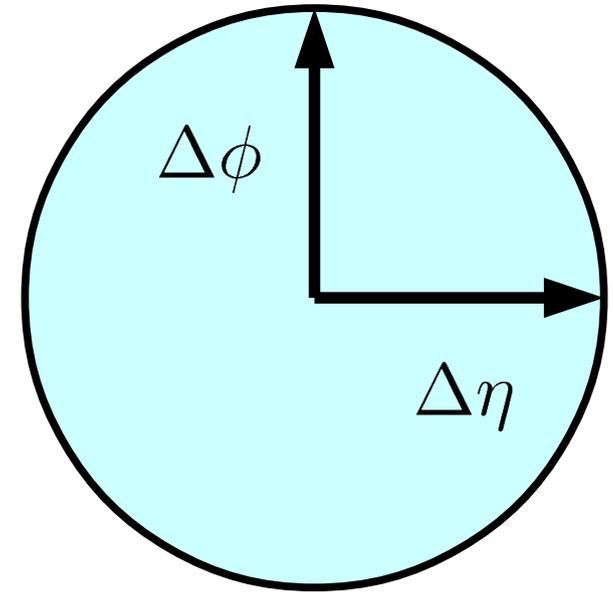
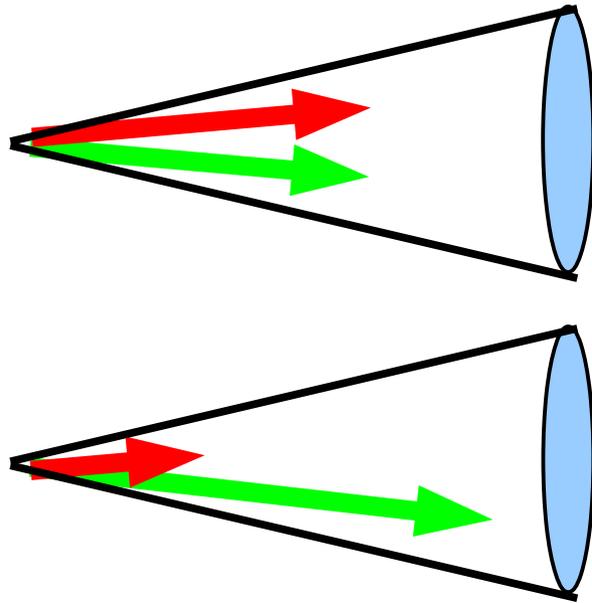
Collinear



if $p_a \parallel p_b$

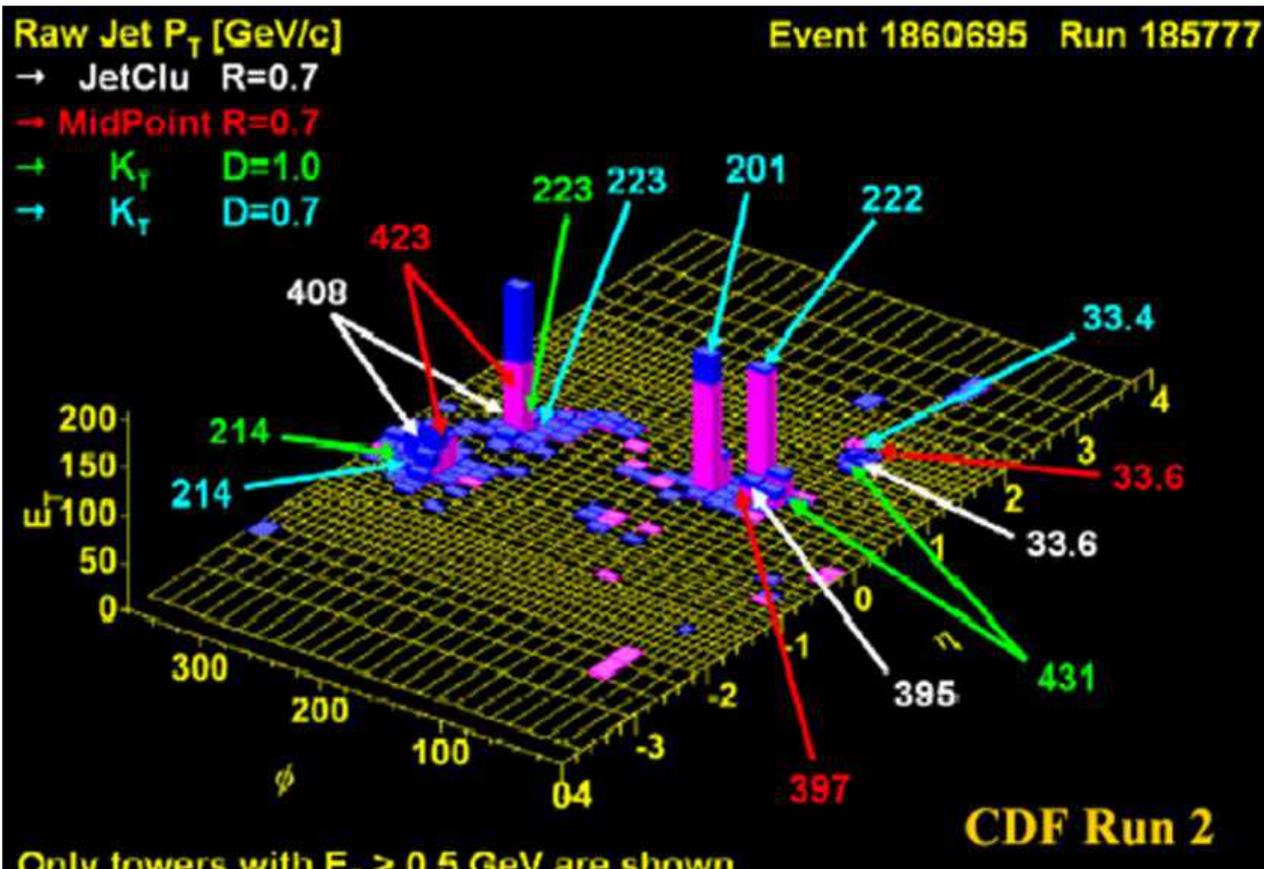
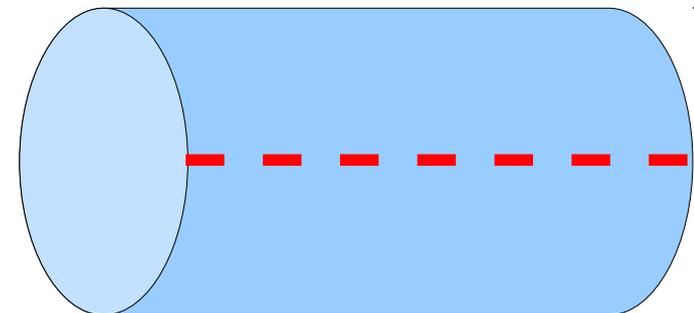
$$\mathcal{O}_{n+1}(p_1, \dots, p_a, p_b, \dots, p_n) \longrightarrow \mathcal{O}_n(p_1, \dots, p_a + p_b, \dots, p_n)$$

Jet Cone

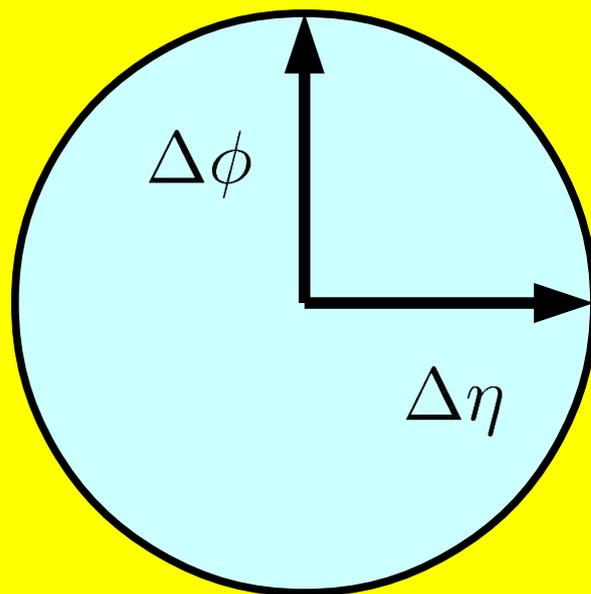


$$R^2 = (\Delta\eta)^2 + (\Delta\phi)^2$$

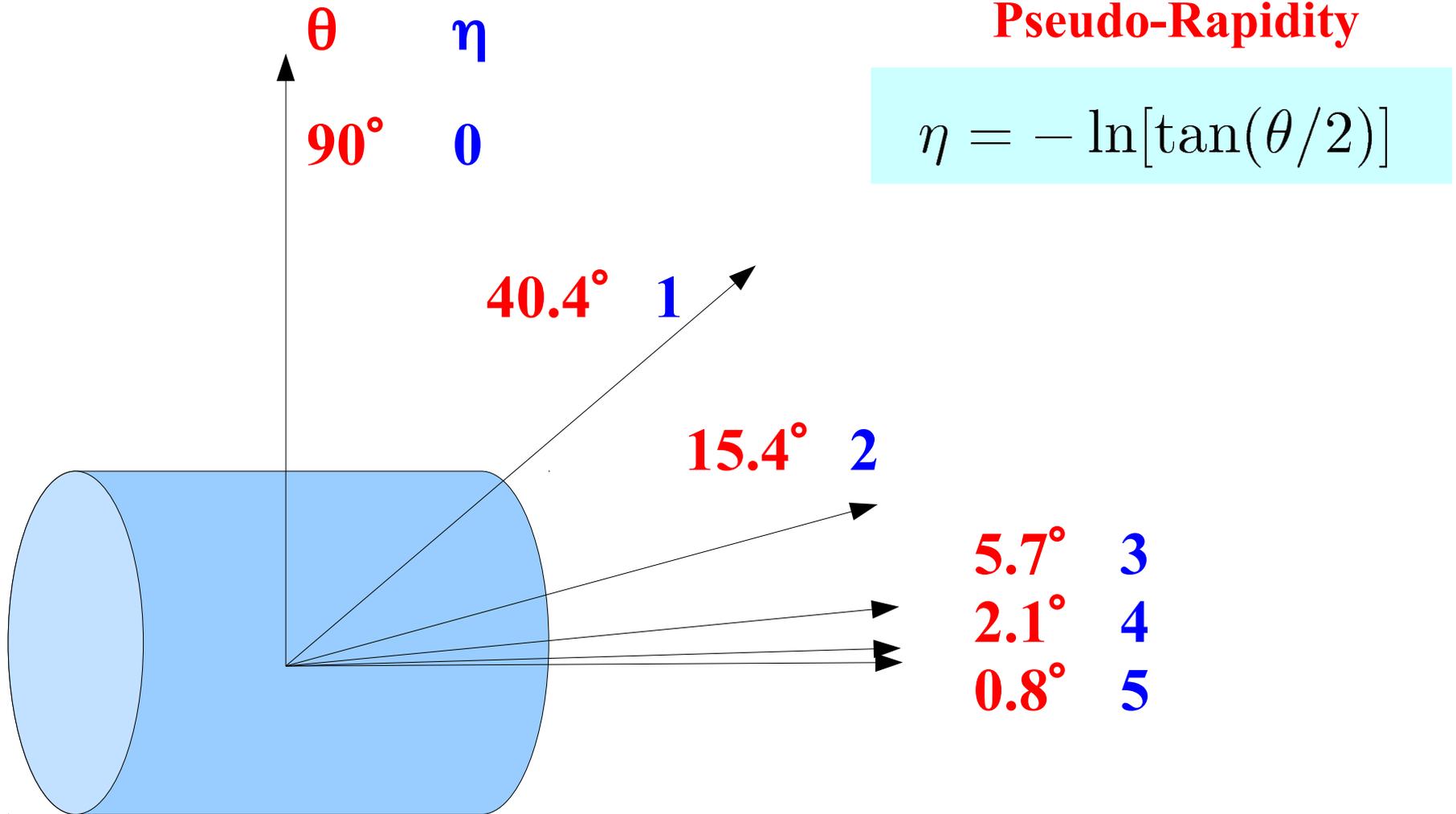
Let's examine this definition a bit more closely

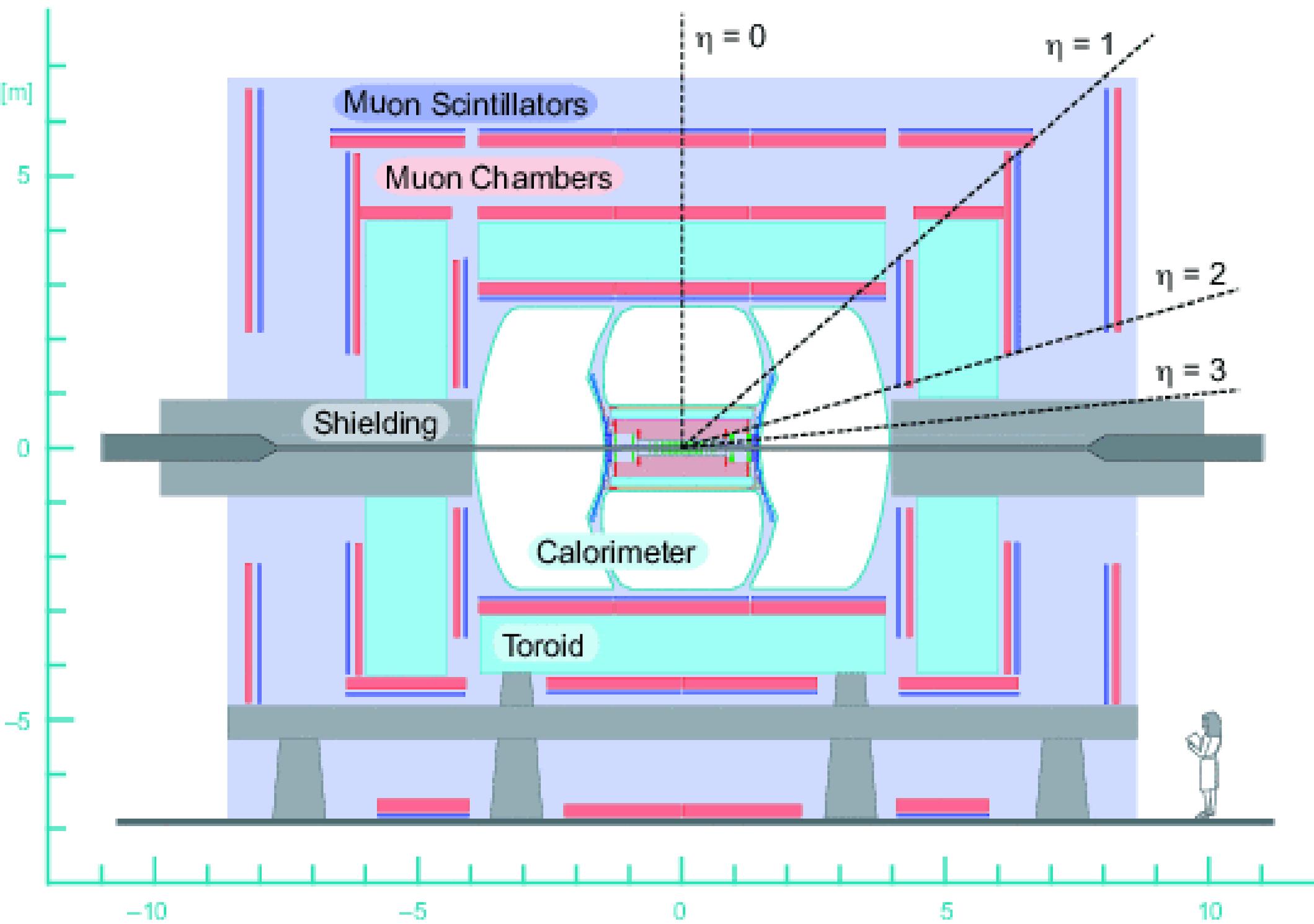


Jet Cone

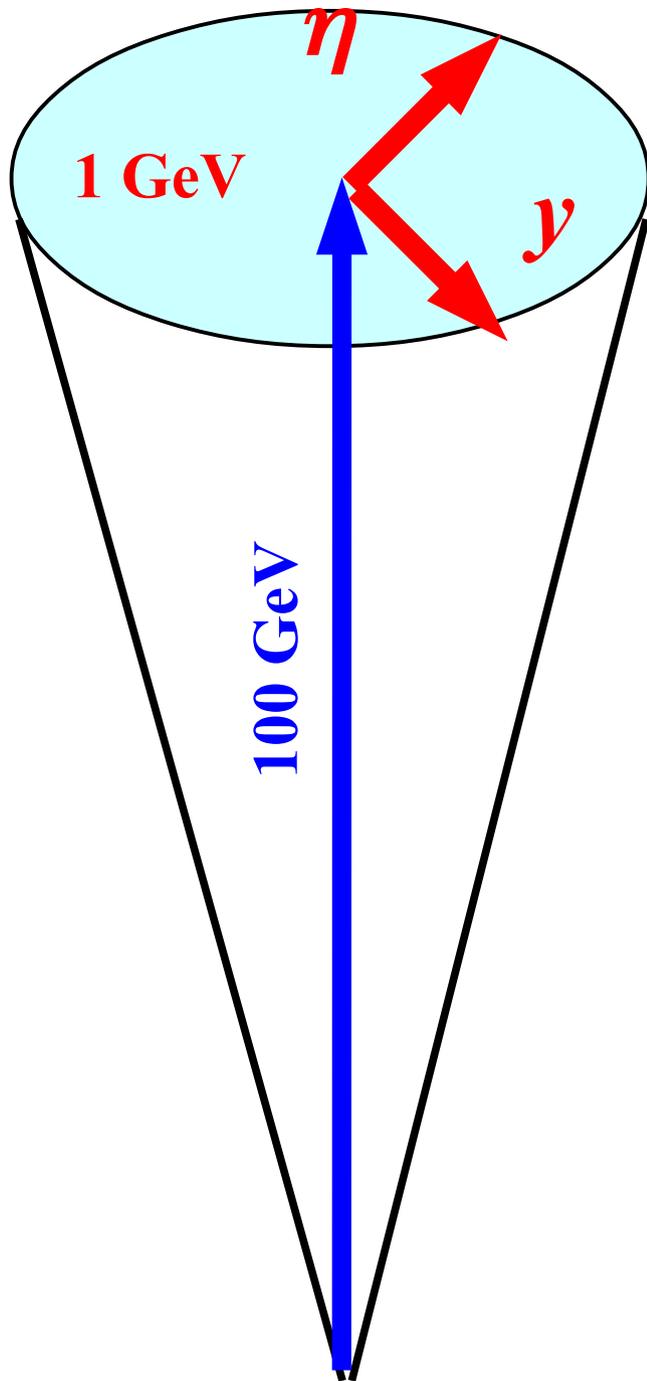


$$R^2 = (\Delta\eta)^2 + (\Delta\phi)^2$$





homework



PROBLEM #2: In a Tevatron detector, consider two particles traveling in the transverse direction:

$$p_1^\mu = \{E, 100, 0, 1\}$$

$$p_2^\mu = \{E, 100, 1, 0\}$$

where the components are expressed in GeV units. E is defined such that the particles are massless.

- a) Compute E .
- b) For each particle, compute the pseudorapidity η and azimuthal angle ϕ .
- c) Explain how the above exercise justifies the correct jet radius definition to be:

$$R = \sqrt{\eta^2 + \phi^2}$$

In particular, why is the above correct and $R = \sqrt{\eta^2 + 2\phi^2}$, for example, incorrect.

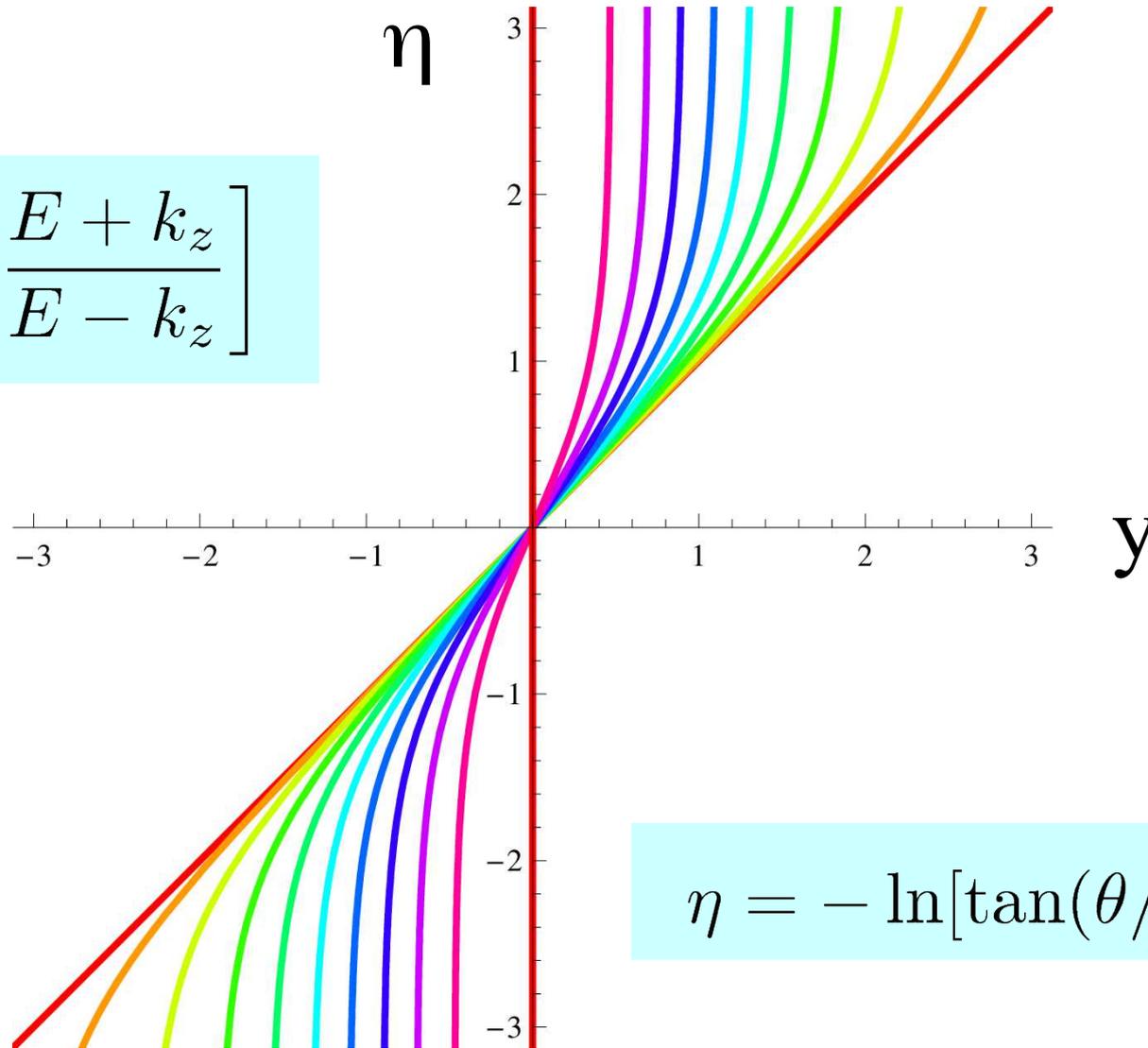
$$P_\mu = \{P_t, P_x, P_y, P_z\}$$

$$P_\mu = \{P_+, \vec{P}_\perp, P_-\} \quad \vec{P}_\perp = \{P_x, P_y\}$$

$$P_\pm = \frac{1}{\sqrt{2}} (P_t \pm P_z)$$

- 1) Compute the metric $g_{\mu\nu}$ in the light-cone frame, and compute $\vec{P}_1 \cdot \vec{P}_2$ in terms of the light-cone components.
- 2) Compute the boost matrix B for a boost along the z -axis, and show the light-cone vector transforms in a particularly simple manner.
- 3) Show that a boost along the z -axis uniformly shifts the rapidity of a vector by a constant amount.

$$y = \frac{1}{2} \ln \left[\frac{E + k_z}{E - k_z} \right]$$



$$\eta = -\ln[\tan(\theta/2)]$$

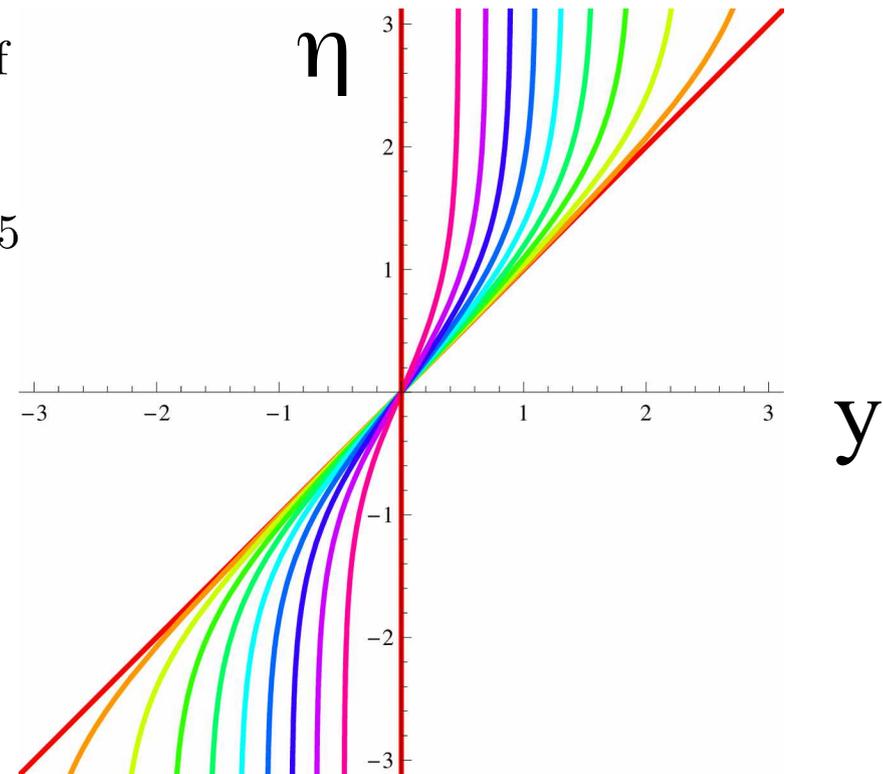
$$\frac{M}{E} = \{0, 0.1, 0.2, \dots\}$$

PROBLEM #1: Consider the rapidity y and the pseudo-rapidity η :

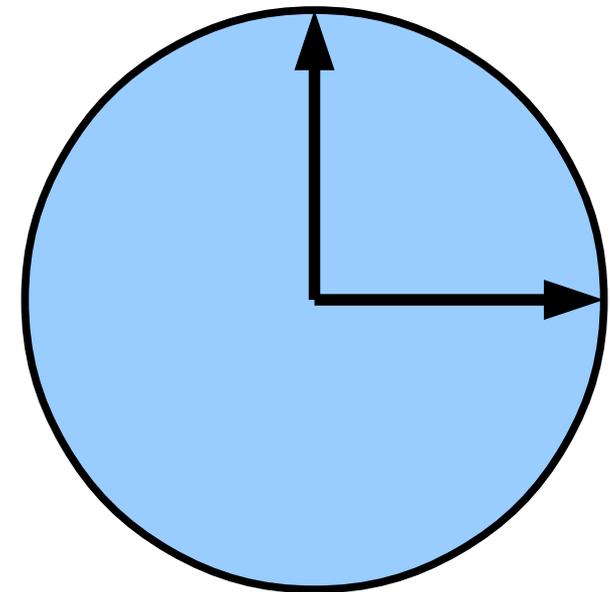
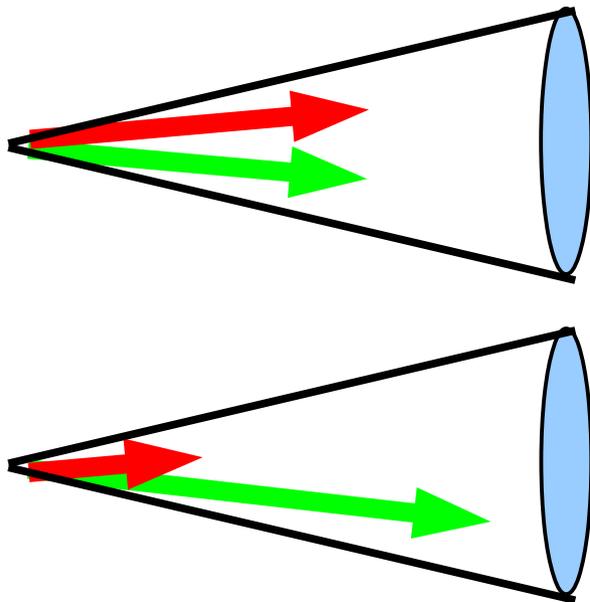
$$y = \frac{1}{2} \ln \left(\frac{E + P_z}{E - P_z} \right)$$

$$\eta = -\ln \left[\tan \left(\frac{\theta}{2} \right) \right]$$

- a) Make a parametric plot of $\{y, \eta\}$ as a function of the particle.
- b) Show that in the limit $m \rightarrow 0$ that $y \rightarrow \eta$.
- c) Make a table of η for $\theta = [0^\circ, 180^\circ]$ in steps of 5°
- d) Make a table of θ for $\eta = [0, 10]$ i

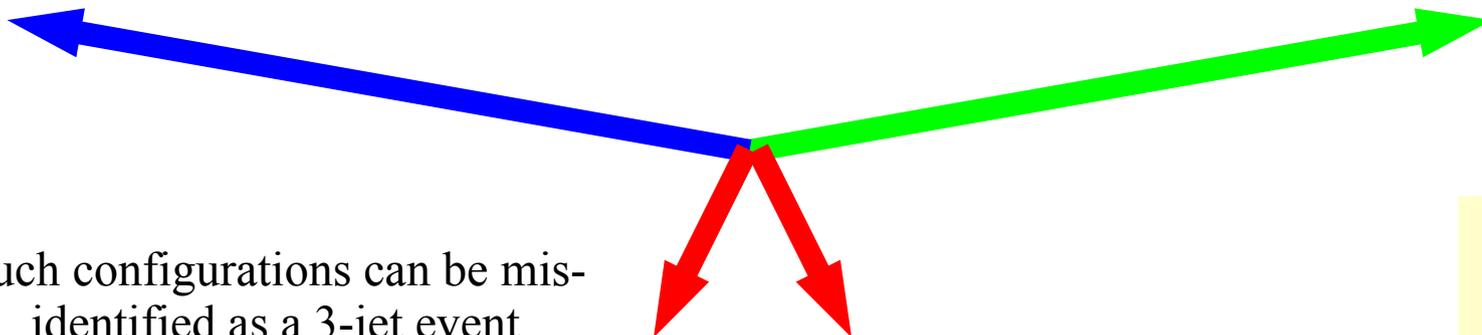


Jet Cone



Problem:
The cone definition is simple,
BUT
it is too simple

$$R^2 = (\Delta\eta)^2 + (\Delta\phi)^2$$



Such configurations can be mis-identified as a 3-jet event

See talk by
Dave Soper &
Andrew Larkoski

Drell-Yan: Tremendous discovery potential

Need to compute 2 initial hadrons

e^+e^- processes:

Total Cross Section:

Differential Cross Section: singularities

Infrared Safe Observables

Stable under soft and collinear emissions

Jet definition

Cone definition is simple:

... it is TOO simple

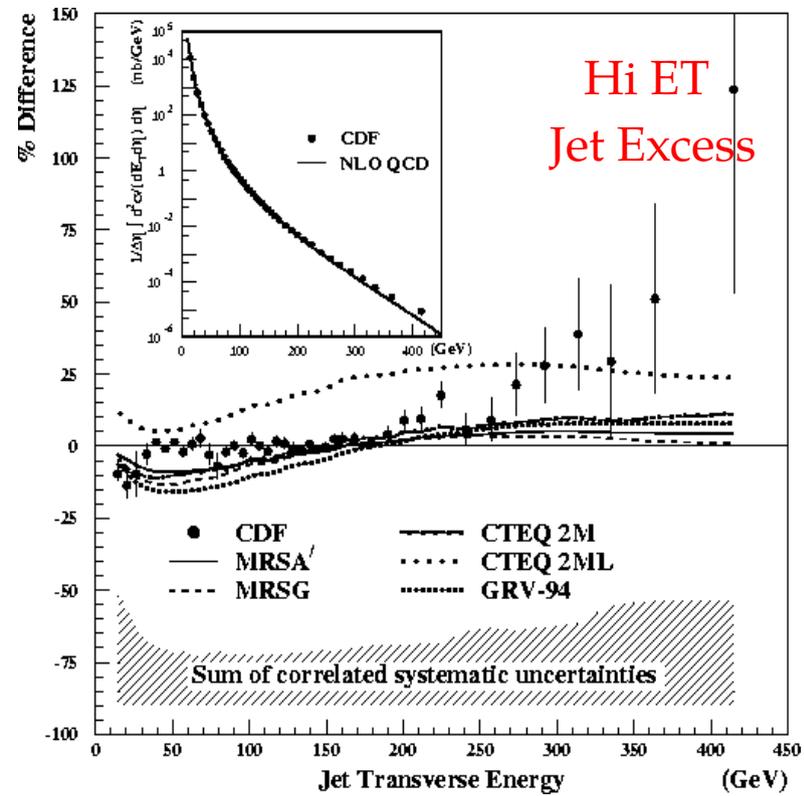
Final Thoughts

$$\begin{aligned} \mathcal{L}_{\text{QCD}} &= \bar{\psi}_i (i\gamma^\mu (D_\mu)_{ij} - m \delta_{ij}) \psi_j - \frac{1}{4} G_{\mu\nu}^a G_a^{\mu\nu} \\ &= \bar{\psi}_i (i\gamma^\mu \partial_\mu - m) \psi_i - g G_\mu^a \bar{\psi}_i \gamma^\mu T_{ij}^a \psi_j - \frac{1}{4} G_{\mu\nu}^a G_a^{\mu\nu} \end{aligned}$$

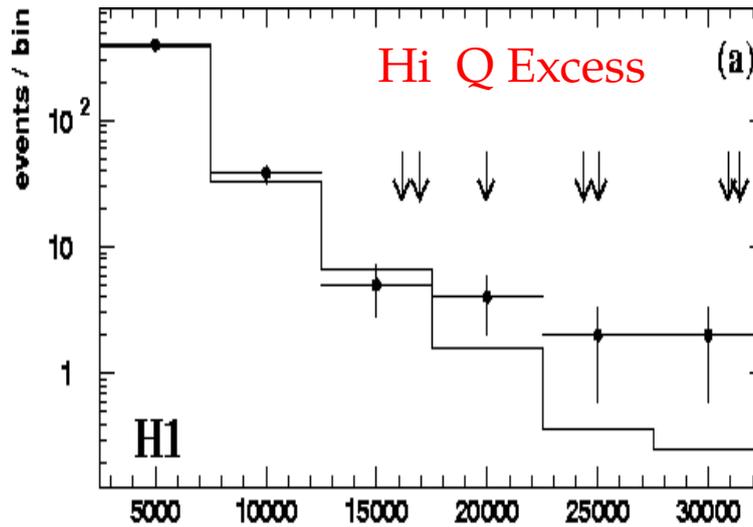
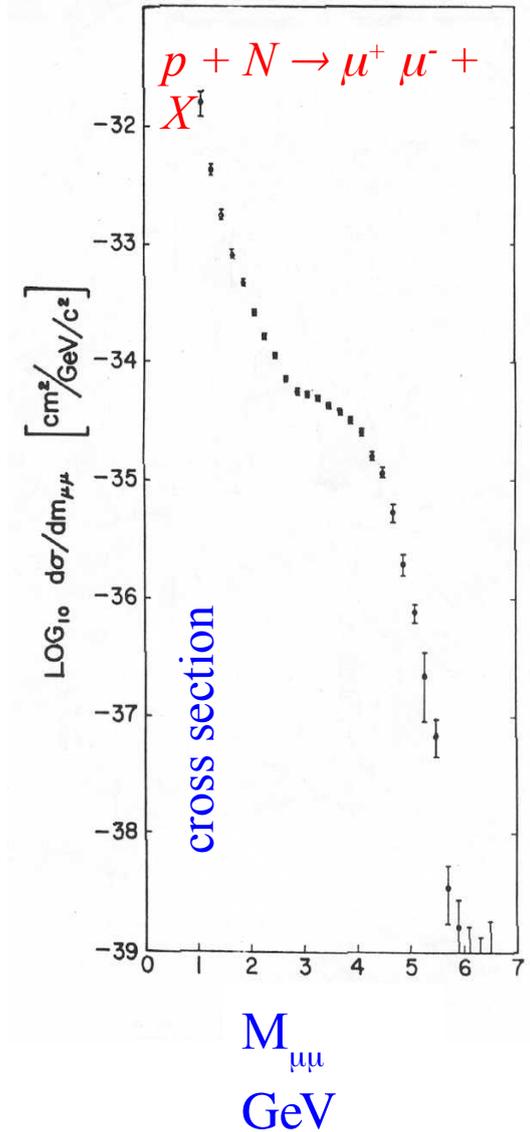
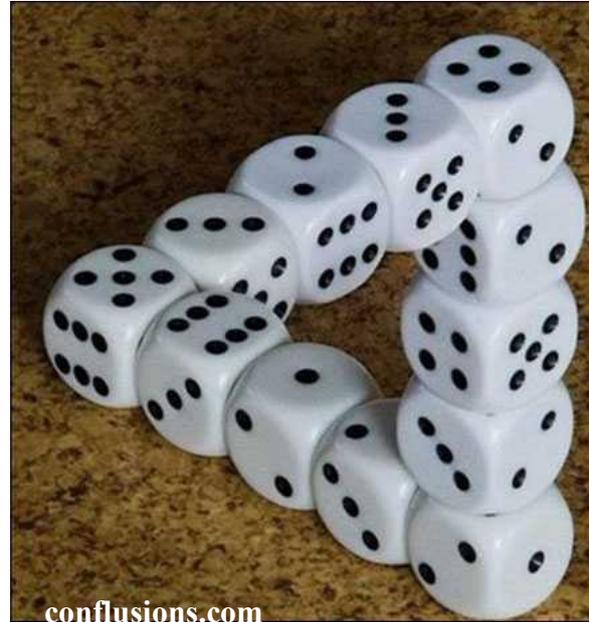
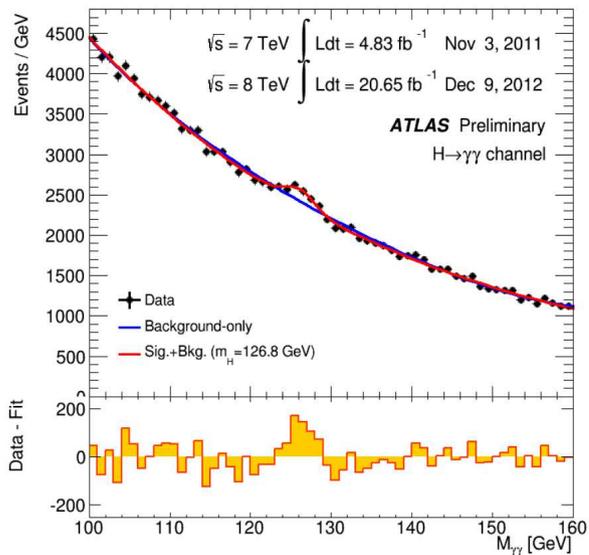
TimePivot

29 30 31 32 33 34 35 36

First player plays the upper staff.



CDF Collaboration, PRL 77, 438 (1996)



H1 Collaboration, ZPC74, 191 (1997) Q_0^2 (GeV²)
 ZEUS Collaboration, ZPC74, 207 (1997)

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for making all this possible.

FRANK AND ERNEST®

THERE ARE SOME THINGS I'VE ALWAYS WONDERED ABOUT...

LIKE WHAT?

LIKE, HOW CAN YOU TELL WHEN IT'S EXACTLY MIDNIGHT?

EASY. THE DARKNESS IS DIRECTLY OVERHEAD.

GEE! AND WHY DO DAYS GET LONGER IN THE SUMMER?

BECAUSE HEAT MAKES THINGS EXPAND!

AND WHY IS AIR SPEED DIFFERENT FROM GROUND SPEED?

SIMPLE. BECAUSE THE EARTH IS ROUND AND THE AIR IS FLAT.

AND WHAT HOLDS THINGS TOGETHER?

VELCRO. NEUTRONS AND PROTONS ARE HELD TOGETHER BY VELCRO.

THANK YOU.

DON'T MENTION IT. I HAVE A NATURAL TALENT FOR SCIENCE.

END OF LECTURE

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