

# General Theory of Relativity (not Theory of General Relativity)

(Opposed to the Special Theory of Relativity.)

What's the difference?

Misconception #1

X Special Rel. can not handle accelerations  
(non-inertial reference frames.)

Sure it can! Later: Relativistic Rocket.

The difference: In S.R. the 4-dimensional spacetime is "flat". In G.R., the spacetime is "curved."

$$G_{\mu\nu} = 8\pi T_{\mu\nu}$$

Einstein Field Equation (EFE)  
geometrized units  
 $G=1$   $c=1$   $\frac{\text{light year}}{\text{year}}$

start unpacking

$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

$$G = \text{Newton's constant} = 6.67 \times 10^{-11} \frac{\text{N}\cdot\text{m}^2}{\text{kg}^2}$$

$$c = \text{Speed of Light in Vacuum} = 2.9979 \times 10^8 \frac{\text{m}}{\text{s}}$$

$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu} \quad \text{— tensor equation}$$

↑  
Einstein tensor

↑ energy-momentum tensor = stress energy tensor  
energy (including mass-energy) density, pressures, stresses.

Curvature of 4-dim spacetime

John Archibald Wheeler:

"Matter tells spacetime how to curve."

$$G_{\mu\nu} = 8\pi T_{\mu\nu}$$

Spacetime tells matter how to move."

↑ Geodesic Law - later.

tensor equations - form-invariant  $\equiv$  covariant - describe transformations

under:  $\vec{F} = m\vec{a}$   $\xrightarrow{\text{vector rotation}}$   $\vec{F}' = m\vec{a}'$

Maxwell  $\rightarrow \partial_\mu F^{\mu\nu} = 0$   $\xrightarrow{\text{4-vector Lorentz boost}}$   $\partial'_\mu F'^{\mu\nu} = 0$

$$G_{\mu\nu} = 8\pi T_{\mu\nu} \quad \xrightarrow{\text{general coordinate transformations}} \quad G'_{\mu\nu} = 8\pi T'_{\mu\nu}$$

Rotations

$$\begin{pmatrix} x' \\ y' \\ z' \\ t' \end{pmatrix} = \begin{pmatrix} \dots \\ \dots \\ \dots \\ \dots \end{pmatrix} \begin{pmatrix} x \\ y \\ z \\ t \end{pmatrix}$$

Boosts

$$\Lambda = \begin{pmatrix} \dots \\ \dots \\ \dots \\ \dots \end{pmatrix}$$

GR

$$\begin{cases} x' = f(x, y, z, t) \\ y' = g(x, y, z, t) \\ z' = \dots \\ t' = \dots \end{cases}$$

## Field Equation

Charges & currents tell the field what to be.

$$\vec{\nabla} \cdot \vec{E}(\vec{r}) = \frac{\rho(\vec{r})}{\epsilon_0}$$

$$\vec{\nabla} \times \vec{B}(\vec{r}) = \mu_0 \vec{J}(\vec{r})$$

## Equation of Motion

Fields ( $\vec{E}, \vec{B}$ ) tell charges how to move

$$m \frac{d^2 \vec{x}}{dt^2} = \vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$$

Matter tells spacetime how to curve

$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu}$$

$\tau$  = proper time

Spacetime tells matter how to move

$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\alpha\beta}^{\mu} \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau} = 0$$

$R_{\alpha\beta\gamma\delta}$

Riemann tensor  
256 components

test for ~~curvature~~ curvature.

$$\text{Cartesian } g = \begin{matrix} & x & y & z \\ \begin{matrix} x \\ y \\ z \end{matrix} & \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \end{matrix}$$

$$R_{\alpha\beta\gamma\delta} = 0$$

$$R_{\mu\nu} = \sum_{\rho=\sigma}^3 R^{\rho}{}_{\mu\rho\nu}$$

$\uparrow$  Ricci tensor  
 $\uparrow$  Riemann Curvature tensor

$\leftarrow 4 \times 4 \times 4 \times 4$  object  
 $R_{0000} =$   
 $R_{0001} =$   
 $R_{0002} =$

$\uparrow$  256 components  $\rightarrow$  20 independent components

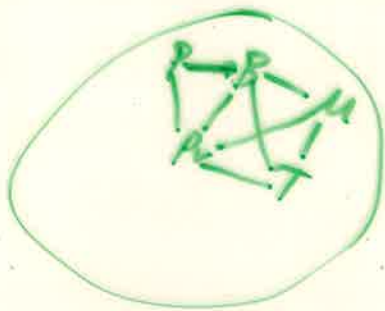
$$R^{\alpha}{}_{\beta\rho\sigma} = \frac{\partial \Gamma^{\alpha}{}_{\beta\rho}}{\partial x^{\sigma}} - \frac{\partial \Gamma^{\alpha}{}_{\beta\sigma}}{\partial x^{\rho}} + \Gamma^{\alpha}{}_{\lambda\rho} \Gamma^{\lambda}{}_{\beta\sigma} - \Gamma^{\alpha}{}_{\lambda\sigma} \Gamma^{\lambda}{}_{\beta\rho}$$

$\Gamma^{\alpha}{}_{\beta\rho}$  = Christoffel symbols

$$\Gamma^{\alpha}{}_{\beta\rho} = \frac{1}{2} \left( \frac{\partial g_{\alpha\rho}}{\partial x^{\beta}} + \frac{\partial g_{\alpha\beta}}{\partial x^{\rho}} - \frac{\partial g_{\beta\rho}}{\partial x^{\alpha}} \right)$$

$g_{\mu\nu}$  = metric tensor

tells about distances between points  
(and other things)



$T_{\mu\nu} =$

	0	1	2	3	0	$t$
0	$\infty$	-	-	-	1	$x$
1	-	$\lambda$	$\lambda$	$\lambda$	2	$r$
2	-	$\lambda$	$\lambda$	$\lambda$	3	$z$
3	-	$\lambda$	$\lambda$	$\lambda$		

Reproduces Newton's Law.

Einstein knew  $G_{00} = 8\pi T_{00}$

$\Rightarrow$  must be true for all components

$G_{11} = 8\pi T_{11} \quad | \quad G_{15} = 8\pi T_{15}$

$G_{\mu\nu} = 8\pi T_{\mu\nu}$

Symmetric tensors  
16 comp.  $\rightarrow$  10 independent comp.

10 2nd-order, coupled non-linear inhomogeneous Partial differential Equations.

Einstein  $\downarrow$

$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu}$

$\uparrow$  Ricci tensor       $\uparrow$  metric tensor

Ricci scalar  $R = \sum_{\nu=0}^3 R^{\nu}_{\nu} = R^{\nu}_{\nu}$

$\sum_{\nu=0}^3 g^{\mu\nu} R_{\mu\nu} = R^{\mu}_{\mu}$

contraction = trace