Apparent magnitude	Brightness relative to magnitude 0	Example	Apparent magnitude	Brightness relative to magnitude 0	Example	Apparent magnitude	Brightness relative to magnitude 0	Pogson's ratio: $\sqrt[5]{100} \approx 2.512$ Example
-27	6.31 × 10 ¹⁰	Sun	-7	631	SN 1006 supernova	13	6.31 × 10 ^{−6}	3C 273 quasar / limit of 4.5–6" (11–15 cm) telescopes
-26	2.51 × 10 ¹⁰		-6	251	ISS (max)	14	2.51 × 10 ^{−6}	Pluto (max) / limit of 8–10" (20–25 cm) telescopes
-25	1 × 10 ¹⁰		-5	100	Venus (max)	15	1 × 10 ^{−6}	
-24	3.98 × 10 ⁹		-4	39.8		16	3.98 × 10 ^{−7}	Charon (max)
-23	1.58 × 10 ⁹		-3	15.8	Jupiter (max), Mars (max)	17	1.58 × 10 ^{−7}	
-22	6.31 × 10 ⁸		-2	6.31	Mercury (max)	18	6.31 × 10 ^{−8}	
-21	2.51 × 10 ⁸		-1	2.51	Sirius	19	2.51 × 10 ^{−8}	
-20	1 × 10 ⁸		0	1	Vega, Saturn (max)	20	1 × 10 ⁻⁸	
-19	3.98 × 10 ⁷		1	0.398	Antares	21	3.98 × 10 ^{−9}	Callirrhoe (satellite of Jupiter)
-18	1.58 × 10 ⁷		2	0.158	Polaris	22	1.58 × 10 ⁻⁹	
-17	6.31 × 10 ⁶		3	0.0631	Cor Caroli	23	6.31 × 10 ⁻¹⁰	
-16	2.51 × 10 ⁶		4	0.0251	Acubens	24	2.51 × 10 ⁻¹⁰	
-15	1 × 10 ⁶		5	0.01	Vesta (max), Uranus (max)	25	1 × 10 ⁻¹⁰	Fenrir (satellite of Saturn)
-14	3.98 × 10 ⁵		6	3.98 × 10 ^{−3}	typical limit of naked eye ^[note 2]	26	3.98 × 10 ⁻¹¹	
-13	1.58 × 10 ⁵	Full moon	7	1.58 × 10 ^{−3}	Ceres (max)	27	1.58 × 10 ^{−11}	visible light limit of 8m telescopes
-12	6.31 × 10 ⁴		8	6.31 × 10 ⁻⁴	Neptune (max)	28	6.31 × 10 ⁻¹²	
-11	2.51 × 10 ⁴		9	2.51 × 10 ^{−4}		29	2.51 × 10 ⁻¹²	
-10	1 × 10 ⁴		10	1 × 10 ⁻⁴	typical limit of 7x50 binoculars	30	1 × 10 ⁻¹²	
-9	3.98 × 10 ³	Iridium flare	11	3.98 × 10 ^{−5}		31	3.98 × 10 ⁻¹³	
-8	1.58 × 10 ³		12	1.58 × 10 ^{−5}		32	1.58 × 10 ^{−13}	visible light limit of HST



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Stellar Astrophysics

• Stefan-Boltzmann Law:

$$F_{bol} = \sigma T^4; \sigma = \frac{2\pi^5 k^4}{15c^2 h^3} = 5.67 \times 10^{-5} \, ergs^{-1} \, cm^{-2} \, K^{-4}$$

- Effective temperature of a star: Temp. of a black body with the same luminosity per surface area
- Stars can be treated as black body radiators to a good approximation
- Effective surface temperature can be obtained from the B-V color index with the Ballesteros equation:

$$T = 4600 \left(\frac{1}{0.92(B-V) + 1.70} + \frac{1}{0.92(B-V) + 0.62} \right)$$

• Luminosity:

$$L = 4\pi r_*^2 \sigma T_E^4$$



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