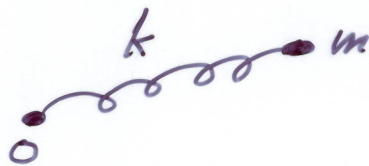
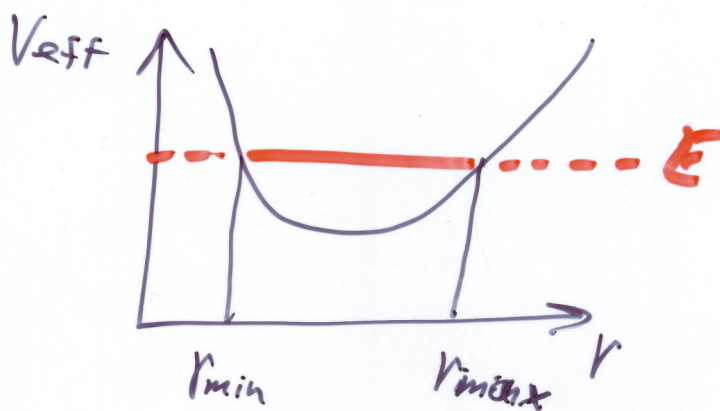
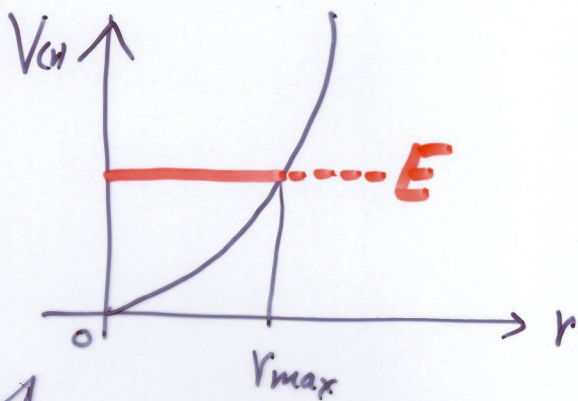


Ideal Hooke's Law Spring



$$\vec{F}(r) = -k\vec{r} \quad V(r) = \frac{1}{2}kr^2$$



↑ also V_{eff} if $l=0$
straight line through origin

$$\vec{F} = -k\vec{r} \Rightarrow F_x = -kx \quad F_y = -ky$$

Two orthogonal independent simple harmonic oscillators - same frequency, different phases.

$$x = A \cos(\omega t)$$

$$y = B \cos(\omega t + \varphi)$$



$$A=B \\ \varphi = \pi/2$$



$$A \neq B$$



$$\varphi = 0$$

Virial Theorem

Newton 2nd; $\dot{\vec{p}}_i = \vec{F}_i$

$$G \equiv \sum_{i=1}^N \vec{p}_i \cdot \vec{v}_i$$

$$\frac{dG}{dt} = \sum_i \dot{\vec{p}}_i \cdot \vec{v}_i + \sum_i \vec{p}_i \cdot \dot{\vec{v}}_i$$

$$\downarrow \\ \sum_i \vec{F}_i \cdot \vec{v}_i$$

$$\hookrightarrow \sum_i m_i \dot{\vec{v}}_i \cdot \dot{\vec{v}}_i = \sum_i m_i v_i^2 = 2T$$

Time average $Z(t)$

$$\langle Z \rangle = \frac{1}{\tau} \int_{t=0}^{\tau} Z(t) dt$$

$$\left\langle \frac{dG}{dt} \right\rangle = \frac{1}{\tau} \int_{t=0}^{\tau} \frac{dG}{dt} dt = \frac{1}{\tau} [G(\tau) - G(0)] = 0$$

① G is periodic, choose $\tau = n$ periods

② \vec{p}_i, \vec{r}_i bounded, let $\tau \rightarrow \infty$

$$= 2\langle T \rangle + \left\langle \sum_i \vec{F}_i \cdot \vec{r}_i \right\rangle$$

$$\Rightarrow \langle T \rangle = -\frac{1}{2} \left\langle \sum_i \vec{F}_i \cdot \vec{r}_i \right\rangle \quad \text{Virial - Rudolf Clausius}$$

e.g. Ideal Gas Law

Equipartition theorem \Rightarrow each atom has $\frac{1}{2} k_B T$ per degree of freedom. $\langle T \rangle = \frac{3}{2} k_B T N$

ideal gas - ignore interatomic forces.
only forces of constraint from walls.

$$\sum_i \vec{F}_i \cdot \vec{r}_i = -P \oint_{\text{walls}} \vec{r}_0 \cdot d\vec{A} = -P \iiint (\vec{\nabla}_0 \cdot \vec{r}) dV = -P \nabla V$$

$$\vec{\nabla}_0 \cdot \vec{r} = \left(\hat{i} \frac{\partial}{\partial x} + \hat{j} \frac{\partial}{\partial y} + \hat{k} \frac{\partial}{\partial z} \right) \cdot (x\hat{i} + y\hat{j} + z\hat{k}) = 3$$

$$\langle T \rangle = -\frac{1}{2} \left\langle \sum_i \vec{F}_i \cdot \vec{r}_i \right\rangle \Rightarrow \frac{3}{2} N k_B T = \frac{3PV}{2} \Rightarrow PV = N k_B T$$

If \vec{F} is central and comes from a potential energy

$$V(r) \Rightarrow \vec{F} = -\vec{\nabla} V \Rightarrow |F| = \frac{\partial V}{\partial r}$$

$$\langle T \rangle = +\frac{1}{2} \left\langle r \frac{\partial V}{\partial r} \right\rangle$$

If $V = k r^{n+1}$ so $F \propto r^n$ then $r \frac{\partial V}{\partial r} = (n+1)V$

$$\langle T \rangle = \frac{n+1}{2} \langle V \rangle \Rightarrow \begin{cases} \text{Inverse square law force } n=-2 \\ \langle T \rangle = \frac{1}{2} \langle V \rangle \\ \text{Hooke's law } n=+1 \\ \langle T \rangle = \langle V \rangle \end{cases}$$

Last time, we found $r(t)$ and $\theta(t)$ parametrized by time t . Often we want to know the shape of the orbit: $r(\theta)$

$$l = m r^2 \dot{\theta} = m r^2 \frac{d\theta}{dt} \Rightarrow dt = \frac{m r^2}{l} d\theta$$

$$\frac{d}{dt} = \frac{l}{m r^2} \frac{d}{d\theta}$$

$$\dot{r} = \frac{dr}{dt} = \sqrt{\frac{2}{m} \left(E - V(r) - \frac{l^2}{2m r^2} \right)}$$

V_{eff}

$$dt = \frac{dr}{\sqrt{\frac{2}{m} \left(E - V(r) - \frac{l^2}{2m r^2} \right)}} \Rightarrow d\theta = \frac{l}{m r^2} \frac{dr}{\sqrt{\frac{2}{m} \left(E - V(r) - \frac{l^2}{2m r^2} \right)}}$$

Convenient Change of Variable

$$u = \frac{1}{r}$$

$$du = -\frac{1}{r^2} dr$$

$$u_0 = \frac{1}{r_0}$$

$$V(r) = k r^{n+1} \rightarrow k u^{-(n+1)}$$

$$\frac{du}{dr} = -\frac{1}{r^2} = -u^2$$

$$\theta - \theta_0 = - \int_{u^a = u_0}^u \frac{du}{\sqrt{\frac{2mE}{l^2} - \frac{2mk}{l^2} u^{-(n+1)} - u^{-2}}}$$

Bertrand's theorem on closed orbits.

Closed orbits



not closed



Euler-Lagrange for r : $m \ddot{r} = f(r) + \frac{l^2}{m r^3}$

Convert to u

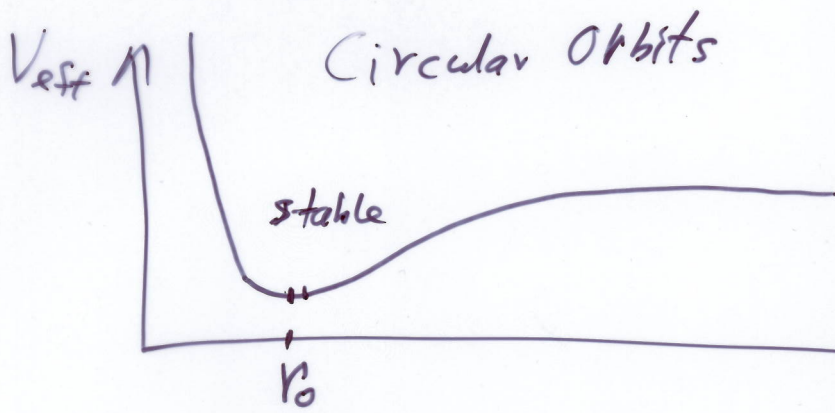
$$m \frac{l}{m r^2} \frac{d}{d\theta} \left[\frac{l}{m r^2} \frac{dr}{d\theta} \right] = - \frac{dV(r)}{dr} + \frac{l^2}{m r^3}$$

$$= \frac{l^2}{m} u^2 \frac{d}{d\theta} \left[u^2 \frac{d}{d\theta} \left(\frac{1}{u} \right) \right] = - \frac{dV}{du} \frac{du}{dr} + \frac{l^2}{m} u^3$$

$\underbrace{\frac{d}{d\theta} \left(\frac{1}{u} \right)}_{-\frac{1}{u^2} \frac{du}{d\theta}}$

$$= - \frac{l^2 u^2}{m} \frac{d^2 u}{d\theta^2} = u^2 \frac{dV(\frac{1}{u})}{du} + \frac{l^2 u^3}{m}$$

$$\frac{d^2 u}{d\theta^2} = - \frac{m}{l^2} \frac{dV(\frac{1}{u})}{du} - u$$



If $r = r_0 \Rightarrow \dot{r} = 0, \ddot{r} = 0 \Rightarrow f(r_0) = -\frac{l^2}{m r_0^3}$

$\dot{r} = 0 \Rightarrow E = V(r_0) + \frac{l^2}{2m r_0^2} = \frac{1}{2} m v_\theta^2$ (no $\frac{1}{2} m v_r^2$)

For stability, $V_{\text{eff}}(r)$ must be concave up

$$\frac{d^2 V_{\text{eff}}(r)}{dr^2} > 0 \Rightarrow \left. \frac{df}{dr} \right|_{r=r_0} + \frac{3l^2}{m r_0^4} > 0$$

$$\left. \frac{df}{dr} \right|_{r=r_0} < -\frac{3f(r_0)}{r_0}$$

If $f(r) = -k r^n$ this is $-k n r^{n-1} < 3 k r^{n-1}$

$$\Rightarrow \boxed{n > -3}$$



Simple harmonic motion for small deviations away from $r = r_0$ circular orbits

$$U(\theta) = \underset{\substack{\uparrow \\ r_0 \text{ circles}}}{U_0} + a \cos(\beta \theta) \quad \text{small } a \ll U_0$$