1. Along an adiabat, the pressure and volume of an ideal gas are related by  $PV^{\gamma} = \text{constant.}$  Show that

$$\frac{dT}{dP} = F(\gamma)\frac{T}{P}$$

and find the function  $F(\gamma)$ .

- 2. Helium gas is heated in a process for which the molar heat capacity  $c_X$  (neither  $c_V$  for constant volume nor  $c_P$  for constant pressure) is 2R, where R is the universal gas constant. During the process, the volume of the gas quadruples.
  - (a) How does the absolute temperature of the gas change?
  - (b) How does the pressure of the gas change?
- 3. See the attached problem with a PV diagram.
- 4. See the attached problem on the last page.

## Bonus

- 1. (a) Describe and draw all of the quadratic degrees of freedom of a carbon dioxide molecule.
  - (b) What is f at high temperature, so that none of the modes are frozen out?
  - (c) Treating  $CO_2$  as an ideal gas, what is the molar heat capacity at constant volme in terms of the gas constant R?
  - (d) What is the molar heat capacity at constant pressure?

The adjacent *p-V* diagram shows the so-called Stirling cooling cycle (refrigerator). Its working fluid is a monoatomic gas (for instance, helium). Processes  $1\rightarrow 2$  and  $3\rightarrow 4$  are isothermal: the fluid is held at constant temperature by thermal baths at temperatures  $T_{\rm h}$  ("hot") and  $T_{\rm c}$  ("cold"), respectively. Processes  $2\rightarrow 3$  and  $4\rightarrow 1$  are isochoric: they take place at constant volumes  $V_2$  and  $V_1$ , respectively. The heat capacity of the working fluid at constant volume, per mole, is  $C_v$ .

2)



Determine the heat absorbed by the fluid in ....

- a. ... the process  $1 \rightarrow 2$
- *b.* ... the process  $2 \rightarrow 3$
- c. ... the process  $3 \rightarrow 4$
- d. ... the process  $4 \rightarrow 1$

e. Determine the net work done on the fluid per cycle.

f. Is the cylle reversible? Explain.

**Problem 1.55.** Heat capacities are normally positive, but there is an important class of exceptions: systems of particles held together by gravity, such as stars and star clusters.

- (a) Consider a system of just two particles, with identical masses, orbiting in circles about their center of mass. Show that the gravitational potential energy of this system is −2 times the total kinetic energy.
- (b) The conclusion of part (a) turns out to be true, at least on average, for any system of particles held together by mutual gravitational attraction:

$$\overline{U}_{\text{potential}} = -2\overline{U}_{\text{kinetic}}.$$

Here each  $\overline{U}$  refers to the total energy (of that type) for the entire system, averaged over some sufficiently long time period. This result is known as the **virial theorem**. (For a proof, see Carroll and Ostlie (1996), Section 2.4.) Suppose, then, that you add some energy to such a system and then wait for the system to equilibrate. Does the average total kinetic energy increase or decrease? Explain.

- (c) A star can be modeled as a gas of particles that interact with each other only gravitationally. According to the equipartition theorem, the average kinetic energy of the particles in such a star should be  $\frac{3}{2}kT$ , where T is the average temperature. Express the total energy of a star in terms of its average temperature, and calculate the heat capacity. Note the sign.
- (d) Use dimensional analysis to argue that a star of mass M and radius R should have a total potential energy of  $-GM^2/R$ , times some constant of order 1.
- (e) Estimate the average temperature of the sun, whose mass is  $2 \times 10^{30}$  kg and whose radius is  $7 \times 10^8$  m. Assume, for simplicity, that the sun is made entirely of protons and electrons.