

1. Starting with the definition of the factorial

$$N! = \int_{x=0}^{\infty} x^N e^{-x} dx = \Gamma(N + 1)$$

show all the steps to derive the formula

$$\int_{r=0}^{\infty} r^{N-1} \exp(-r^2) dr = \frac{1}{2} \Gamma\left(\frac{N}{2}\right)$$

that we used in the lecture notes.

2. Remember our convention for n -spheres: $n = 2$ dimensions corresponds to the film of a soap bubble and has surface area $S_2 = 4\pi r^2$; the volume of the solid spherical ball is $V_3 = \frac{4}{3}\pi r^3$; the “volume” of a disk is $V_2 = \pi r^2$. (n is always the power of r).
- (a) For which (non-integer) dimension n is the area of the unit ($r = 1$) hypersphere a maximum?
 - (b) For which (non-integer) dimension n is the volume of the unit hypersphere a maximum?
 - (c) Plot graphs of S_n and V_n versus n .
3. (a) Using last week’s problem on the entropy of a black hole, find the temperature of a black hole as a function of its mass.
- (b) Find the temperature of a solar-mass black hole.
 - (c) Find the temperature of the largest known black hole.
 - (d) Find the temperature of a Planck-mass black hole.
4. Consider a two-state paramagnet with $N = 10^{23}$ elementary dipoles, with the total energy fixed at zero so that exactly half of the dipoles point up and half point down. Give all answers first as a function of N , then numerically.
- (a) How many total states (with any energy) does the system have?
 - (b) How many microstates have zero energy? Give the exact formula in terms of N , then use Stirling’s approximation.
 - (c) What fraction of states have energy zero? (100% is the wrong answer.)
 - (d) Suppose that the microstate can change a googol (10^{100}) times every second. How long will it take the system to explore all the zero-energy microstates. Compare this to the age of the Universe, 14 billion years.
 - (e) Comment on the statement, “If you wait long enough, the system eventually will be found in any of its accessible microstates.”

Bonus

1. According to the Sackur-Tetrode equation (which we will derive soon),

$$S = Nk_B \left[\ln \left(\frac{V}{N} \left(\frac{4\pi mU}{3Nh^2} \right)^{3/2} \right) + \frac{5}{2} \right]$$

the entropy of a monatomic ideal gas can become negative when its temperature (and hence its energy) is sufficiently low. Of course this is absurd, so the Sackur-Tetrode equation must be invalid at very low temperatures. Suppose that you start with a sample of helium at room temperature and atmospheric pressure, then lower the temperature while holding the density fixed. Pretend that the helium remains a gas and does not liquefy. Below what temperature would the Sackur-Tetrode equation predict that S is negative?

2. (a) Consider a horizontal slab of atmosphere whose thickness is dz . Find an expression for $\frac{dP}{dz}$ in terms of the average mass m of an air molecule and the number density $n = N/V$.
- (b) Use the ideal gas law to find a differential equation for the pressure.
- (c) Assume that the temperature does not vary with height and solve the differential equation for the pressure as a function of height.
- (d) Explain the boundary condition at $z=0$.