

#1. $M_{\odot} = 2 \times 10^{30} \text{ kg}$
 $R_{\odot} = 7 \times 10^5 \text{ km}$

$$\bar{\rho} = \frac{M_{\odot}}{\frac{4}{3} \pi R_{\odot}^3}$$

$$\bar{\rho} = 1.4 \times 10^3 \text{ kg/m}^3$$

$$\rho = 1.4 = 1.4 \text{ g/cm}^3$$

AVERAGE DENSITY ABOUT THE SAME AS PVC PIPE.

#2. $E = pc = \hbar kc$ RELATIVISTIC CASE

$$P_{\text{deg}} = - \frac{\partial E_{\text{TOT}}}{\partial V}$$

THE KEY IS TO FIND E_{TOT} .

FOLLOW EXAMPLE IN CLASS THAT USES SHOEBOX POTENTIAL. WE WILL USE A CUBICAL SHOEBOX

$$k = \left[\left(\frac{n_x \pi}{L} \right)^2 + \left(\frac{n_y \pi}{L} \right)^2 + \left(\frac{n_z \pi}{L} \right)^2 \right]^{1/2}$$

$$E_F = \frac{\hbar \pi c}{L} (n_x^2 + n_y^2 + n_z^2)^{1/2} = \frac{\hbar \pi c}{L} n \quad (1)$$

$n = \frac{L E_F}{\hbar \pi c}$ n is a KIND OF RADIUS IN SPACE THAT LABELS STATES.

Pr. FIND # OF e^- 'S YOU CAN FIT IN THIS SPACE AS A FUNCTION OF E_F .

$$N_e = 2 \cdot \frac{1}{8} \cdot \frac{4}{3} \pi R^3 = \frac{4}{3} \pi \frac{L^3}{8} = \frac{\pi}{3} \frac{E_F^3 L^3}{\hbar^3 \pi^3 c^3}$$

↑
SPIN

ONLY POS n_x, n_y, n_z ALLOWED?

$$N_e = \frac{1}{3\pi^2} \frac{E_F^3 L^3}{\hbar^3 c^3} \quad (2)$$

$$\Rightarrow E_F = 3\pi^2 \hbar^3 c^3 N_e / L^3$$

$$E_F = \hbar c (3\pi^2 n_e)^{1/3} \quad \text{w/ } n_e = e^- \text{ NUMBER DENSITY} \quad (3)$$

COMPUTE E_{TOT} USING TECHNIQUE IN TEXT. FIND E_F FOR A SHELL IN n -SPACE, COUNT # OF e^- 'S IN THE SHELL, MULTIPLY TOGETHER AND THEN SUM OVER ALL SHELLS.

$$E_{TOT} = \int_0^{\infty} \frac{\hbar c}{L} n d^n \quad \text{w/ } d^n = n^2 dn \text{ in } n\text{-space}$$

$$E_{TOT} = \frac{1}{4} \frac{h \pi^2 c}{L} \cdot 4\pi \int_0^R n^3 dn$$

$$= \frac{h \pi^2 c}{L} R^4 / 4 \quad \text{w/ } R = \text{max } n \text{ VALUE}$$

FIND R: $E_F = \frac{h \pi^2 c}{L} n_{max}$

$$\Rightarrow R = n_{max} = \frac{L E_F}{h \pi^2 c}$$

$$R^3 = \left(\frac{L E_F}{h \pi^2 c} \right)^3 = \frac{L}{\pi^2} \left(\frac{L E_F}{h c} \right)^3 = \frac{3 \pi^2}{\pi^3} N_e$$

using (1)

$$R^4 = (R^3)^{4/3} = \left(\frac{3 N_e}{\pi} \right)^{4/3}$$

(5)

HENCE

$$E_{TOT} = \frac{h \pi^2 c}{4L} \left(\frac{3 N_e}{\pi} \right)^{4/3}$$

$$E_{TOT} = \frac{h \pi^2 c}{4} V^{-1/3} \left(\frac{3 N_e}{\pi} \right)^{4/3} \quad (6)$$

$$= \frac{h \pi^2 c}{4} \frac{L^3}{L^4} \left[\frac{3 N_e}{\pi} \right]^{4/3}$$

$$E_{TOT} = \frac{h \pi^2 c}{4} V \left(\frac{3 n_e}{\pi} \right)^{4/3}$$

w/ $n_e = e^-$ density.

(7)

~~PLEASE~~ NOTE THAT $E_{TOT} \propto N_e^{4/3}$ FOR
 RELATIVISTIC CASE, EQ. (6) WHILE
 FOR NON-REL CASE, $E_{TOT} \propto N_e^{5/3}$.
 HENCE WE EXPECT DEGENERACY PRESSURE
 TO BE LESS FOR "COLD" RELATIVISTIC
 e^- 'S THAN FOR NON-RELATIVISTIC e^- 'S.

$$P_{DEG} = - \frac{\partial E_{TOT}}{\partial V}$$

$$P_{DEG} = + \frac{4\pi^2 c}{12} \left(\frac{3 n_e}{\pi} \right)^{4/3}$$

RELATIVISTIC CASE

#3. FIND R_{EQ} AT WHICH

$$P_{REF} = P_{GRAVITY}$$

USE NR P_{REF} . FROM TEXT.

FIND $P_{GRAVITY}$

$$P_{GR} = - \frac{dU_G}{dV} \quad \text{w/ } U_G = \text{GRAV POT. E.}$$

USE TECHNIQUE FROM PHYS 3344:

COMPUTE GRAV PE FROM SPHERE w/
MASS SHELL WRAPPED AROUND.



$$dU_G = - G \frac{\left(\frac{4}{3}\pi \rho r^3\right) \cdot 4\pi \rho r^2 dr}{r}$$

$$= - \frac{(4\pi)^2 G \rho^2 r^4 dr}{3}$$

w/ $\rho = \text{MASS DENSITY}$
ASSUME CONST.

$$U_G = - \frac{(4\pi)^2 G \rho^2}{3} \int_0^R r^4 dr = - \frac{(4\pi)^2}{15} G \rho^2 R^5$$

BUT $\frac{4}{3}\pi \rho R^3 = M = N m_n$ w/ $m_n = \text{NUCLEON MASS}$.

#3.

$$U_G = -\frac{3}{5} \left(\frac{4\pi}{3}\right)^{1/3} G (N m_m)^2 V^{-1/3}$$

$$\Rightarrow P_G = -\frac{dU}{dV} = -\frac{1}{5} \left(\frac{4\pi}{3}\right)^{1/3} G (N m_m)^2 V^{-4/3}$$

- SIGN MEANS P_G INWARD.

WHEN DO 2 PRESSURES BALANCE?

$$\frac{1}{5} \left(\frac{4\pi}{3}\right)^{1/3} G (N m_m)^2 V^{-4/3} = \frac{t^2 \pi^3}{15 m_e} \left(\frac{3 N e}{4\pi}\right)^{5/3} V^{-5/3} \quad (1)$$

$$\text{NOW } P_{eg} = \left(\frac{3}{4\pi}\right)^{1/3} V^{1/3}$$

SO FIND $V^{1/3}$ FROM EQ (1)

$$V^{1/3} = \frac{t^2 \pi^3}{G m_e m_m^2} \cdot \frac{1}{3} 2^{5/3} \left(\frac{3}{4\pi}\right)^{7/3} N^{-1/3}$$

$$= \frac{t^2}{G m_e m_m^2} \cdot \frac{2}{3} 2^{2/3} \cdot \pi \left(\frac{9}{16}\right) \left(\frac{3}{4\pi}\right)^{1/3} N^{-1/3}$$

$$= \frac{t^2}{G m_e m_m^2} \cdot \frac{3\pi}{8} 2^{2/3} \frac{3^{4/3}}{2^{4/3} \cdot \pi^{1/3}} N^{-1/3}$$

$$= \frac{t^2}{G m_e m_m^2} \frac{3\pi}{8} \frac{3^{1/3}}{\pi^{1/3}} N^{-1/3}$$

#3.

$$\text{So, } R_{EQ} = \left[\frac{t^2}{G M_e m_n^2} \right] N^{-1/3} \left(\frac{3}{4\pi} \right)^{1/3} \frac{3\pi}{\rho} \left(\frac{3}{4\pi} \right)^{1/3}$$
$$= \left[\right] N^{-1/3} \left[\frac{9 \cdot 3^3 \pi^3}{4\pi^2} \right]^{1/3} \frac{1}{\rho}$$

$$R_{EQ} = \frac{t^2}{G M_e m_n^2} \cdot N^{-1/3} \left[\frac{9\pi}{4} \right]^{1/3} \frac{3}{\rho}$$

$$R_{EQ} = \frac{3}{\rho} \left[\frac{9\pi}{4} \right]^{1/3} \frac{t^2}{G M_e m_n^2} N^{-1/3}$$

$$R_{EQ} \approx 1.76 \times 10^7 \text{ m}$$

$$R_{EQ} \approx 1.8 \times 10^4 \text{ km}$$

$$\omega N = 10^{57}$$

$$m = 0.9 m_p$$

#4. (a) From Procs 3.

$$R_{ED} = \frac{3}{\rho} \left(\frac{9\pi}{4} \right)^{1/3} \frac{t_s^2}{6M_{\odot}^3} \nu^{-1/3}$$

$$= \frac{m_e}{m_n} R_{ED} \text{ (WHITE DWARF)}$$

$$\approx \frac{1}{1800} R_{ED} \text{ (WHITE DWARF)}$$

$$\approx \frac{1.76 \times 10^7 \text{ m}}{1800}$$

$$R_{ED} \approx 9.8 \times 10^3 \text{ m}$$

$$R_{ED} \approx 10 \text{ km}$$

$$(b) \bar{\rho} = \frac{m}{V} = \frac{5 \times 2 \times 10^{30} \text{ kg}}{\frac{4}{3}\pi (10^4)^3 \text{ m}^3}$$

$$\bar{\rho} \approx 2.4 \times 10^{18} \text{ kg/m}^3$$

$$\approx 2.4 \times 10^{15} \text{ g/cm}^3$$

$$\rho_{PB} \approx 11.75 \text{ g/cm}^3$$

$$\rho_{NS} / \rho_{COAL} \approx 2 \times 10^{14}$$

HS- SEE PROB #2, EQ (7):

$$E_F = \frac{1}{2} c (\frac{3}{2} \pi^2 n_e)^{2/3}$$

MASSLESS
FERMIONS

w/ n_e = # DENSITY OF MASSLESS
FERMIONS.