

Tests of Lorentz and CPT violation with Neutrinos

09/24/12

Teppei Katori, MIT



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Tests of Lorentz and CPT violation with Neutrinos

outline

- 1. Spontaneous Lorentz symmetry breaking
- 2. What is Lorentz and CPT violation?
- 3. Modern test of Lorentz and CPT violation
- 4. Lorentz violation with neutrino oscillation
- 5. MiniBooNE experiment
- 6. Test for Lorentz violation with MiniBooNE data
- 7. Future test of Lorentz violation with neutrinos
- 8. Conclusion

Teppei Katori Massachusetts Institute of Technology SMU HEP seminar, Dallas, TX, Sep. 24, 2012

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1. Spontaneous Lorentz symmetry breaking

- 2. What is Lorentz and CPT violation?
- **3. Modern tests of Lorentz and CPT violation**
- 4. Lorentz violation with neutrino oscillation
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1. Spontaneous symmetry breaking

Every fundamental symmetry needs to be tested, including Lorentz symmetry.

After the recognition of theoretical processes that create Lorentz violation, testing Lorentz invariance becomes very exciting

Lorentz and CPT violation has been shown to occur in Planck scale theories, including:

- string theory
- noncommutative field theory
- quantum loop gravity
- extra dimensions
- etc

However, it is very difficult to build a self-consistent theory with Lorentz violation...

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However, it is very difficult to build a self-consistent theory with Lorentz violation...



Y. Nambu (Nobel prize winner 2008), picture taken from CPT04 at Bloomington, IN

1. Spontaneous Lorentz symmetry breaking (SLSB)

vacuum Lagrangian for fermion $L = i\overline{\Psi}\gamma_{\mu}\partial^{\mu}\Psi$

e.g.) SSB of scalar field in Standard Model (SM) - If the scalar field has Mexican hat potential

$$L = \frac{1}{2} (\partial_{\mu} \varphi)^2 - \frac{1}{2} \mu^2 (\varphi^* \varphi) - \frac{1}{4} \lambda (\varphi^* \varphi)^2$$
$$M(\varphi) = \mu^2 < 0$$





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Particle acquires mass term!

Kostelecký and Samuel PRD39(1989)683

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e.g.) SLSB in string field theory

- There are many Lorentz vector fields

- If any of vector field has Mexican hat potential

$$M(a^{\mu}) = \mu^2 < 0$$





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Kostelecký and Samuel PRD39(1989)683

1. Spontaneous Lorentz symmetry breaking (SLSB)

vacuum Lagrangian for fermion
$$L = i\overline{\Psi}\gamma_{\mu}\partial^{\mu}\Psi - m\overline{\Psi}\Psi + \overline{\Psi}\gamma_{\mu}a^{\mu}\Psi$$

e.g.) SSB of scalar field in Standard Model (SM) - If the scalar field has Mexican hat potential

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$$M(a^{\mu}) = \mu^2 < 0$$



Lorentz symmetry is spontaneously broken!



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1. Spontaneous Lorentz symmetry breaking

Test of Lorentz violation is to find the coupling of these background fields and ordinary fields (electrons, muons, neutrinos etc), then physical quantities may depend on the rotation of the earth.



1. Spontaneous Lorentz symmetry breaking

- 2. What is Lorentz and CPT violation?
- 3. Modern tests of Lorentz and CPT violation
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$\overline{\Psi}(\textbf{x})\gamma_{\mu}\textbf{a}^{\mu}\Psi(\textbf{x})$







Under the particle Lorentz transformation:

 $U \overline{\Psi}(x) \gamma_{\mu} a^{\mu} \Psi(x) U^{-1}$





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Under the particle Lorentz transformation:

$$\begin{split} &\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x}) \rightarrow \mathsf{U}[\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x})]\mathsf{U}^{-1} \\ &\neq \overline{\Psi}(\Lambda \mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\Lambda \mathbf{x}) \end{split}$$

Lorentz violation is observable when a particle is moving in the fixed coordinate space



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Under the observer Lorentz transformation:

$$\overline{\Psi}(x)\gamma_{\mu}a^{\mu}\Psi(x)$$



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Under the observer Lorentz transformation:

$$\overline{\Psi}(\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\mathbf{x}) \xrightarrow{\Lambda^{-1}} \overline{\Psi}(\Lambda^{-1}\mathbf{x})\gamma_{\mu}\mathbf{a}^{\mu}\Psi(\Lambda^{-1}\mathbf{x})$$

Lorentz violation cannot be generated by observers motion (coordinate transformation is unbroken)

all observers agree for all observations





CPT symmetry is the invariance under the CPT transformation

 $L \xrightarrow{CPT} \Theta L \Theta^{-1} = L' = L, \qquad \Theta = CPT$

CPT is the perfect symmetry of the Standard Model, due to CPT theorem

CPT theorem If the relativistic transformation law and the weak microcausality holds in a real neighbourhood of a Jost point, the CPT condition holds everywhere.

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CPT phase = $(-1)^{n}$

/ number of Lorentz indices \rightarrow always even number

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CPT-odd Lorentz violating coefficients (odd number Lorentz indices, e.g., a^{μ} , $g^{\lambda\mu\nu}$) CPT-even Lorentz violating coefficients (even number Lorentz indices, e.g., $c^{\mu\nu}$, $\kappa^{\alpha\beta\mu\nu}$)

2. CPT violation implies Lorentz violation



CPT violation implies Lorentz violation in interactive quantum field theory.



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How to detect Lorentz violation?

Lorentz violation is realized as a coupling of particle fields and the background fields, so the basic strategy is to find the Lorentz violation is:

(1) choose the coordinate system to compare the experimental result

- (2) write down Lagrangian including Lorentz violating terms under the formalism
- (3) write down the observables using this Lagrangian

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The standard choice of the coordinate is Sun-centred coordinates





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Standard Model Extension (SME) is a standard formalism for the general search of Lorentz violation. SME is a minimum extension of QFT with Particle Lorentz violation

$$L_{SME} = L_{SM} + \delta L$$

$$\delta L = \overline{\Psi} \gamma_{\mu} a^{\mu} \Psi + \overline{\Psi} \gamma_{\mu} c^{\mu\nu} \partial_{\nu} \Psi \dots$$

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Standard Model Extension (SME) is a standard formalism for the general search of Lorentz violation. SME is a minimum extension of QFT with Particle Lorentz violation

Various physics is predicted under SME, but among them, the smoking gun of Lorentz violation is the sidereal time dependence of the observables.

Solar time: 24h 00m 00.0s sidereal time: 23h 56m 04.1s

Sidereal time dependent physics is often smeared out in solar time distribution \rightarrow Maybe we have some evidence of Lorentz violation but we just didn't notice?!

3. Modern tests of Lorentz violation

Dedicated group of people formed a meeting since 1998.

http://www.physics.indiana.edu/~kostelec/faq.html



Registration

Program

Proceedings

Travel

Accomm@/dabidms

MEETING ON CPT AND LORENTZ SYMMETRY

November 6 - 8, 1998

Physics Department Indiana University, Bloomington

A meeting on CPT and Lorentz symmetry will be held in the <u>Physics Department</u>, <u>Indiana University</u> in <u>Bloomington</u>, Indiana, U.S.A. on November 6 - 8, 1998. The meeting will focus on recent developments involving tests of these fundamental symmetries, including both experimental and theoretical aspects.

Topics to be covered include:

- · experimental bounds on CPT and Lorentz symmetry from
 - measurements on K, B, and D mesons
 - precision comparisons of particle and antiparticle properties (anomalous moments, charge-to-mass ratios, lifetimes, etc.)
 - spectroscopy of hydrogen and antihydrogen
 - ◇ clock-comparison tests
 - properties of light
 other tests

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theoretical descriptions of and constraints on possible violations

3. Modern tests of Lorentz violation

http://www.physics.indiana.edu/~kostelec/faq.html



The second 3. Modern tests of Lorentz violation meeting was in 2001. http://www.physics.indiana.edu/~kostelec/fag.html Second Meeting on に現在地である **CPT** and Lorentz Symmetry August 15-18, 2001 Meeting home Indiana University, Bloomington Registration Program Meeting home A meeting on CPT and Lorentz symmetry will be held in the Physics Department, Indiana University in Bloom U.S.A. on August 15-18, 2001. The meeting will focus on experimental tests of these fundamental symmetric issues, including scenarios for possible violations. **Proceedings** Registration Subjects to be covered include: Travel Program · experimental constraints on CPT and Lorentz symmetry from oscillations and decays of K, B, D mesons and other particles Accommodations comparisons of particle and antiparticle properties Proceedings spectroscopy of hydrogen and antihydrogen

The third 3. Modern tests of Lorentz violation meeting was in 2004. http://www.physics.indiana.edu/~kostelec/fag.html ٥ い思想が行きたが Third Meeting CPT and Lorentz S Meeting home August 4-7, 20 Registration Indiana University, Bloo Program Meeting home Meeting home **Proceedings** Registration The Third Meeting on CPT and Lorentz Symmetry will be held in the Physics Departm Registration August 4-7, 2004. The meeting will focus on experimental tests of these fundamental sy possible violations. Travel Program Program Subjects to be covered include: Accommodations Teppei Katori, MIT Proceedings Proceedings experimental searches for CPT and Lorentz violations involving arrity and interferometric hehavior of nhot

3. Modern tests of Lorentz violation

The fourth meeting was in 2007.


3. Modern tests of Lorentz violation

The latest meeting was in summer 2010. (next meeting will be June 2013)

http://www.physics.indiana.edu/~kostelec/faq.html



Fifth Meeting on CPT AND LORENTZ SYMMETRY

June 28-July 2, 2010

Indiana University, Bloomington

The *Fifth Meeting on CPT and Lorentz Symmetry* will be held in the <u>Physics Department, Indiana</u> <u>University</u> in <u>Bloomington</u>, Indiana, U.S.A. on June 28-July 2, 2010. The meeting will focus on tests of these fundamental symmetries and on related theoretical issues, including scenarios for possible violations.

Topics include:

- · searches for CPT and Lorentz violations involving
 - o birefringence and dispersion from cosmological sources
 - clock-comparison measurements
 - · CMB polarization
 - collider experiments
 - electromagnetic resonant cavities
 - equivalence principle
 - gauge and Higgs particles
 - high-energy astrophysical observations
 - · laboratory and gravimetric tests of gravity

3. Modern tests of Lorentz violation

http://www.physics.indiana.edu/~kostelec/faq.html





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4. Neutrinos

Neutrinos in the standard model

The standard model describes 6 quarks and 6 leptons and 3 types of force carriers.



Neutrinos are special because,

1. they only interact with weak nuclear force.



2. interaction eigenstate is not Hamiltonian eigenstate (propagation eigenstate). Thus propagation of neutrinos changes their species, called neutrino oscillation.

Neutrino oscillation is an interference experiment (cf. double slit experiment)



For double slit experiment, if path v_1 and path v_2 have different length, they have different phase rotations and it causes interference.

Neutrino oscillation is an interference experiment (cf. double slit experiment)



If 2 neutrino Hamiltonian eigenstates, v_1 and v_2 , have different phase rotation, they cause quantum interference.

Neutrino oscillation is an interference experiment (cf. double slit experiment)



If 2 neutrino Hamiltonian eigenstates, v_1 and v_2 , have different phase rotation, they cause quantum interference.

If v_1 and v_2 , have different mass, they have different velocity, so thus different phase rotation.

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The detection may be different flavor (neutrino oscillations).

4. Lorentz violation with neutrino oscillation

Neutrino oscillation is an interference experiment (cf. double slit experiment)



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If 2 neutrino Hamiltonian eigenstates, v_1 and v_2 , have different phase rotation, they cause quantum interference.

If v_1 and v_2 , have different coupling with Lorentz violating field, interference fringe (oscillation pattern) depend on the sidereal motion.

4. Lorentz violation with neutrino oscillation

Neutrino oscillation is an interference experiment (cf. double slit experiment)



If 2 neutrino Hamiltonian eigenstates, v_1 and v_2 , have different phase rotation, they cause quantum interference.

If v_1 and v_2 , have different coupling with Lorentz violating field, interference fringe (oscillation pattern) depend on the sidereal motion.

The measured scale of neutrino eigenvalue difference is comparable the target scale of Lorentz violation (<10⁻¹⁹GeV).

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5. MiniBooNE experiment

MiniBooNE neutrino oscillation experiment at Fermilab is looking for ν_{μ} to ν_{e} oscillation



Signature of v_e event is the single isolated electron like events

Booster Neutrino Beamline (BNB) creates ~800(600)MeV neutrino(anti-neutrino) by pion decay-in-flight from 8GeV Booster protons on Be-target in the magnetic focusing horn.



MiniBooNE collaboration, NIM.A599(2009)28

5. MiniBooNE experiment

MiniBooNE detector is the spherical Cherenkov detector

- v-baseline is ~520m

Booster

- filled with 800t mineral oil
- -1280 of 8" PMT in inner detector
- 240 veto PMT in outer region



FNAL Booster target and horn



primary beam secondary beam tertiary beam (protons) (mesons) (neutrinos)

5. MiniBooNE experiment •*Muons*

-Sharp, clear rings

Long, straight tracks

•Electrons

-Scattered rings

Multiple scattering

Radiative processes





MiniBooNE collaboration, NIM.A599(2009)28

5. MiniBooNE experiment

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MiniBooNE collaboration, NIM.A599(2009)28

5. MiniBooNE oscillation analysis result

Neutrino mode low energy excess MiniBooNE see the excess at low energy region. Antineutrino mode excess MiniBooNE see the excess at combined region.



These excesses are not predicted by neutrino Standard Model (vSM). Oscillation candidate events may have sidereal time dependence.

All backgrounds are measured in other data sample and their errors are constrained

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8. Conclusion

Test for Lorentz violation in MiniBooNE data;

(1) fix the coordinate system

- (2) write down Lagrangian including Lorentz violating terms under the formalism
- (3) write down the observables using this Lagrangian

- Booster neutrino beamline is described in Sun-centred coordinates









 $\Phi = \omega_{e} T_{e}$

 $T_{\oplus}=0$



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- Booster neutrino beamline is described in Sun-centred coordinates
- Standard Model Extension (SME)

Modified Dirac Equation (MDE) of neutrinos

$$i(\Gamma_{AB}^{\nu}\partial_{\nu} - M_{AB})\nu_{B} = 0$$

SME parameters

$$\begin{split} \Gamma_{AB}^{\nu} &= \gamma^{\nu} \delta_{AB} + c_{AB}^{\mu\nu} \gamma_{\mu} + d_{AB}^{\mu\nu} \gamma_{\mu} \gamma_{5} + e_{AB}^{\nu} + i f_{AB}^{\nu} \gamma_{5} + \frac{1}{2} g_{AB}^{\lambda\mu\nu} \sigma_{\lambda\mu} \\ M_{AB} &= m_{AB} + i m_{5AB} \gamma_{5} + a_{AB}^{\mu} \gamma_{\mu} + b_{AB}^{\mu} \gamma_{5} \gamma_{\mu} + \frac{1}{2} H_{AB}^{\mu\nu} \sigma_{\mu\nu} \end{split}$$

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4

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$$M_{AB} = m_{AB} + i m_{5AB} \gamma_{5} + a_{AB}^{\mu} \gamma_{\mu} + b_{AB}^{\mu} \gamma_{5} \gamma_{\mu} + \frac{1}{2} H_{AB}^{\mu\nu} \sigma_{\mu\nu}$$
$$CPT \text{ even}$$

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CPT odd

Test for Lorentz violation in MiniBooNE data;

(1) fix the coordinate system

(2) write down Lagrangian including Lorentz violating terms under the formalism

(3) write down the observables using this Lagrangian

- Booster neutrino beamline is described in Sun-centred coordinates
- Standard Model Extension (SME)
- Sidereal time dependent oscillation probability



$$P_{\nu_{\mu} \rightarrow \nu_{e}} \sim \frac{|(h_{eff})_{e\mu}|^{2} L^{2}}{(\hbar c)^{2}}$$

= $\left(\frac{L}{\hbar c}\right)^{2} |(C)_{e\mu} + (A_{s})_{e\mu} \sin w_{\oplus} T_{\oplus} + (A_{c})_{e\mu} \cos w_{\oplus} T_{\oplus}$
+ $(B_{s})_{e\mu} \sin 2w_{\oplus} T_{\oplus} + (B_{c})_{e\mu} \cos 2w_{\oplus} T_{\oplus} |^{2}$

sidereal frequency
$$W_{\oplus} = \frac{2\pi}{23h56m4.1s}$$

sidereal time T_{\oplus}

Sidereal variation analysis for MiniBooNE is 5 parameter fitting problem

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6. Lorentz violation with MiniBooNE neutrino data

Unbinned extended maximum likelihood fit

- It has the maximum statistic power
- Best fit parameters are extracted

5 parameter fit

sidereal frequency	₩⊕	=	$\frac{2\pi}{23h56m4.1s}$
sidereal time	T⊕		

$$\mathsf{P}_{v_{e} \to v_{\mu}} = \left(\frac{\mathsf{L}}{\hbar \mathsf{c}}\right)^{2} \left| (\mathsf{C})_{e\mu} + (\mathsf{A}_{s})_{e\mu} \sin w_{\oplus} \mathsf{T}_{\oplus} + (\mathsf{A}_{c})_{e\mu} \cos w_{\oplus} \mathsf{T}_{\oplus} + (\mathsf{B}_{s})_{e\mu} \sin 2w_{\oplus} \mathsf{T}_{\oplus} + (\mathsf{B}_{c})_{e\mu} \cos 2w_{\oplus} \mathsf{T}_{\oplus} \right|^{2}$$

- Due to high correlation of parameters, we focus on 3 parameter fit for error evaluation

- Contours are evaluated from fake data study

3 parameter fit

fit

$$\mathsf{P}_{\mathsf{v}_{e} \to \mathsf{v}_{\mu}} = \left(\frac{\mathsf{L}}{\hbar \mathsf{c}}\right)^{2} \left| (\mathsf{C})_{\mathsf{e}\mu} + (\mathsf{A}_{\mathsf{s}})_{\mathsf{e}\mu} \sin \mathsf{w}_{\oplus} \mathsf{T}_{\oplus} + (\mathsf{A}_{\mathsf{c}})_{\mathsf{e}\mu} \cos \mathsf{w}_{\oplus} \mathsf{T}_{\oplus} \right|^{2}$$

MiniBooNE collaboration, ArXiv:1109.3480

6. Lorentz violation with MiniBooNE neutrino data

Neutrino mode result, low energy region

Only C-parameter is nonzero, but this is sidereal independent parameter.

26.9% C.L. with flat hypothesis by fake data $\Delta \chi^2$ study

The neutrino mode low energy excess is consistent with no sidereal variation.



MiniBooNE collaboration, ArXiv:1109.3480

6. Lorentz violation with MiniBooNE anti-neutrino data

Anti-neutrino mode result, combined energy region

As and Ac-parameters are nonzero, which are sidereal dependent parameters.

3.0% C.L. with flat hypothesis by fake data $\Delta \chi^2$ study

The anti-neutrino mode combined energy region excess prefer sidereal time dependent solution, but not statistically significant level.



6. Summary of results

Neutrino result summary

- The low energy excess data fit prefer sidereal time independent solution.
- 26.9% C.L. with flat hypothesis

Anti-neutrino result summary

- The fit for combined region excess data prefers sidereal time dependent solution.

- 3.0% C.L. flat hypothesis

SME coefficients

- The combinations of SME coefficients are extracted
- 2σ limits are set
- First time constrained time independent SME coefficients for e- $\!\mu$ sector

	$\nu\mathrm{-mode}\;\mathrm{BF}$	2σ limit	$\bar{\nu}\mathrm{-mode}\;\mathrm{BF}$	2σ limit	SME coefficients combination (unit 10^{-20} GeV)
$ (\mathcal{C})_{e\mu} $	$3.1\pm0.6\pm0.9$	< 4.2	$0.1\pm0.8\pm0.1$	< 2.6	$\pm \left[(a_L)_{e\mu}^T + 0.75(a_L)_{e\mu}^Z \right] - \langle E \rangle \left[1.22(c_L)_{e\mu}^{TT} + 1.50(c_L)_{e\mu}^{TZ} + 0.34(c_L)_{e\mu}^{ZZ} \right]$
$ (\mathcal{A}_s)_{e\mu} $	$0.6\pm0.9\pm0.3$	< 3.3	$2.4\pm1.3\pm0.5$	< 3.9	$\pm [0.66(a_L)_{e\mu}^Y] - \langle E \rangle [1.33(c_L)_{e\mu}^{TY} + 0.99(c_L)_{e\mu}^{YZ}]$
$ (\mathcal{A}_c)_{e\mu} $	$0.4\pm0.9\pm0.4$	< 4.0	$2.1\pm1.2\pm0.4$	< 3.7	$\pm [0.66(a_L)_{e\mu}^{\dot{X}}] - \langle E \rangle [1.33(c_L)_{e\mu}^{\dot{T}X} + 0.99(c_L)_{e\mu}^{\dot{X}Z}]$

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Double Chooz collaboration PRL108(2012)131801

7. Double Chooz experiment

Reactor electron antineutrino disappearance experiment

- The first result shows small anti- v_e disappearance!



Reactor electron antineutrino disappearance experiment

- The first result shows small anti- v_e disappearance!
- The second result reaches 3.1σ signal
- DayaBay and RENO experiments saw disappearance signals, too

Double Chooz collaboration PRL108(2012)131801 arXiv:1207.6632 DayaBay collaboration PRL108(2012)171803 RENO collaboration PRL108(2012)191802

Double Chooz reactor neutrino candidate





Teppei Katori, MIT

Reactor electron antineutrino disappearance experiment

- The first result shows small anti- v_e disappearance!
- The second result reaches 3.1σ signal
- DayaBay and RENO experiments saw disappearance signals, too



Double Chooz reactor neutrino candidate





Teppei Katori, MIT

Reactor electron antineutrino disappearance experiment

- The first result shows small anti- v_e disappearance!
- The second result reaches 3.1σ signal
- DayaBay and RENO experiments saw disappearance signals, too
- This small disappearance may have sidereal time dependence





So far, we have set limits on 1. $v_e \leftrightarrow v_\mu$ channel: LSND, MiniBooNE, MINOS (<10⁻²⁰ GeV) 2. $v_\mu \leftrightarrow v_\tau$ channel: MINOS, IceCube (<10⁻²³ GeV) The last untested channel is $v_e \leftrightarrow v_\tau$



It is possible to limit $v_e \leftrightarrow v_\tau$ channel from reactor v_e disappearance experiment

 $\mathsf{P}(v_e {\leftrightarrow} v_e) = 1 - \mathsf{P}(v_e {\leftrightarrow} v_{\mu}) - \mathsf{P}(v_e {\leftrightarrow} v_{\tau}) \sim 1 - \mathsf{P}(v_e {\leftrightarrow} v_{\tau})$



Leonard: What do you think about the latest Double Chooz result?09/24/12 Sheldon: I think this is Lorentz violation..., check sidereal time dependence

So far, we have set limits on 1. $v_e \leftrightarrow v_\mu$ channel: LSND, MiniBooNE, MINOS (<10⁻²⁰ GeV) 2. $v_{\mu} \leftrightarrow v_{\tau}$ channel: MINOS, IceCube (<10⁻²³ GeV) The last untested channel is $v_e \leftrightarrow v_{\tau}$

It is possible to limit $v_{\rho} \leftrightarrow v_{\tau}$ channel from reactor v_{ρ} disappearance experiment



09/24/12

(flat)

7. Superluminal neutrinos



7. Superluminal neutrinos

OPERA

v(neutrino) = c + $(2.37\pm0.32) \times 10^{-5}$ c = c + $(16\pm2) \times 10^{3}$ mph

CERN to Gran Sasso Neutrino Beam


7. Superluminal neutrinos	News US World Sports Comment Culture Business Environ News Science Particle physics	/ ////
OPERA v(neutrino) = c + $(2.37\pm0.32) \times 10^{-5}$ c = c + $(16\pm2) \times 10^{3}$ mph	Neutrinos still faster than light in latest version of experimentFinding that contradicts Einstein's theory of special relativity is repeated with fine-tuned procedures and equipment	World Business Investigatio
The Washington Post HOME PAGE TODAY'S PAPER VIDEO MOST POPULAR TIMES TOPICS Posted at 08:25 AM ET, 09/23/2011 Neutrinos may have tra CERN to Griv Posted at 08:25 AM ET, 09/23/2011 Neutrinos may have tra Colspan="2">Image to Day's Paper VIDEO MOST POPULAR TIMES TOPICS Posted at 08:25 AM ET, 09/23/2011 Neutrinos may have tra Ight By Elizabeth Flock Scientists Report Second Sighting of Faster-Than-Light Scientists Report Second Sighting of Faster-Than-Light		faster than the speed of hysics lab, announced a enough to make Albert Einstein s, called <u>neutrinos</u> , have been ed of light.
Neutrinos By DENNIS OVERBYE Published: November 18, 2011 Few scientists are betting against Einstein yet, but the phantom neutrinos of Opera are still eluding explanation. Two months after scientists reported that they had clocked subatomic particles known as neutrinos going	Monday 06 February 2012 The Speed of light appears to have been broken again af	CULTURE TRAVEL LIFESTYLE ence Defence Health News Roy ution Steve Jones Science Picture ests test neutrino ests test neutrino ter scientists carried out a new vize the laws of physics to be

7. Superluminal neutrinos

OPERA

v(neutrino) = c + $(2.37\pm0.32) \times 10^{-5}$ c = c + $(16\pm2) \times 10^{3}$ mph

It is fascinating result, but...

- time of flight is kinematic test (less sensitive than neutrino oscillations)
- no indication of Lorentz violation from any neutrino oscillation experiments
- superluminal neutrino is unstable (vacuum Cherenkov radiation) ArXiv:1109.6562
- pion phase space is limited to create such neutrinos ArXiv:1109.6630
- SN1987A neutrinos provide severe limit to superluminal neutrinos PRL58(1987)1490
 etc...

It is very difficult to interpret superluminal neutrinos at OPERA by Lorentz violation with field theory approach.

OPERA collaboration, ArXiv:1109.4897

7. Superluminal neutrinos



The Washington Post | Make us your start page OPINIONS Posted at 01:23 PM ET. 02/23/2012

Faster-than-light neutrinos aren't?

By Alexandra Petri



You can return to your homes. There is nothing more to see.

It turns out those faster-than-light neutrinos at Europe's CERN lab

theguardian

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News > Science > Cern

Faster-than-light neutrinos: was a faulty connection to blame?

A dodgy optical fibre connection may have skewed results that appeared to show neutrinos travelling faster than light



Faster-than-light neutrinos would breach Einstein's theory of special relativity.

By Jason Palmer Science and technology reporter, BBC News

What might have been the biggest physics story of the past century may instead be down to a faulty connection.

In September 2011, the Opera experiment reported it had seen particles called neutrinos evidently travelling faster than the speed of light.

The team has now found two problems that may have affected their test in opposing ways: one in its timing gear and one in an optical fibre



The neutrinos are fired deep under the Italian

It is hard to topple the giant...

Teppei Katori, MIT

Conclusion

- Lorentz and CPT violation has been shown to occur in Planck scale physics.
- There are world wide effort for the test of Lorentz violation using various type of state-of-art technologies.
- LSND and MiniBooNE data suggest Lorentz violation is an interesting solution of neutrino oscillation.
- MiniBooNE neutrino mode data prefer sidereal time independent solution. On the other hand, anti-neutrino mode data prefer sidereal time dependent solution, although statistical significance is not high enough.
- Constraints from LSND, MiniBooNE, MINOS, IceCube, and Double Chooz set stringent limits on Lorentz violation in neutrino sector in terrestrial level.

Thank you for your attention!