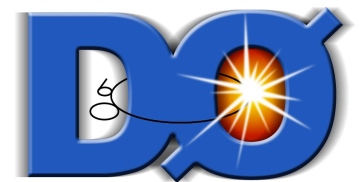


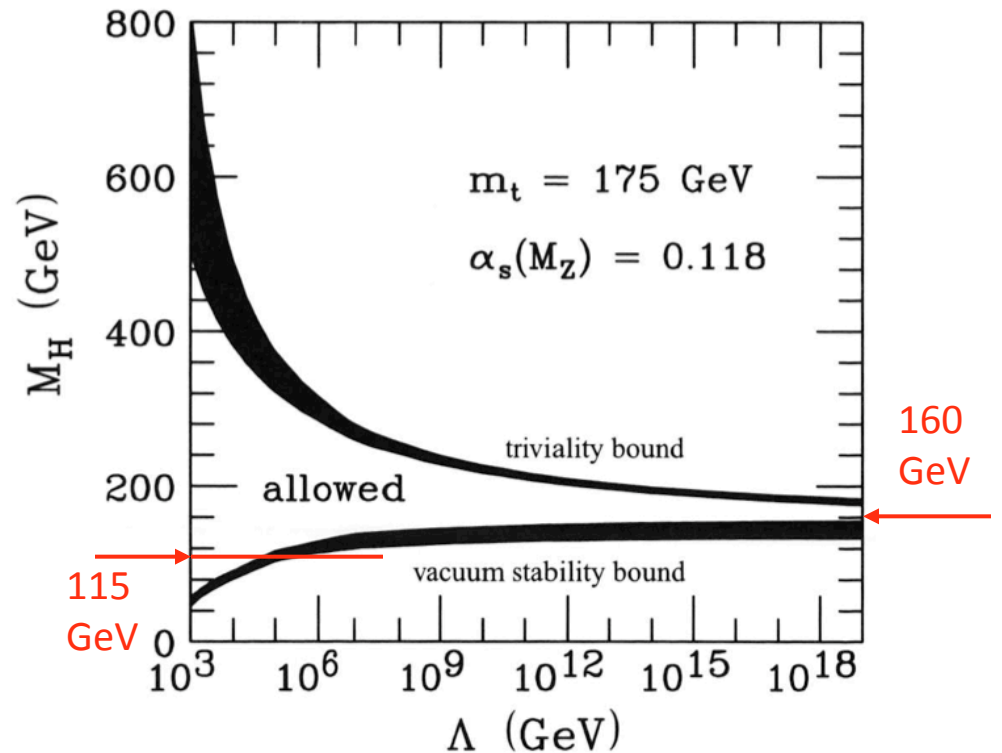
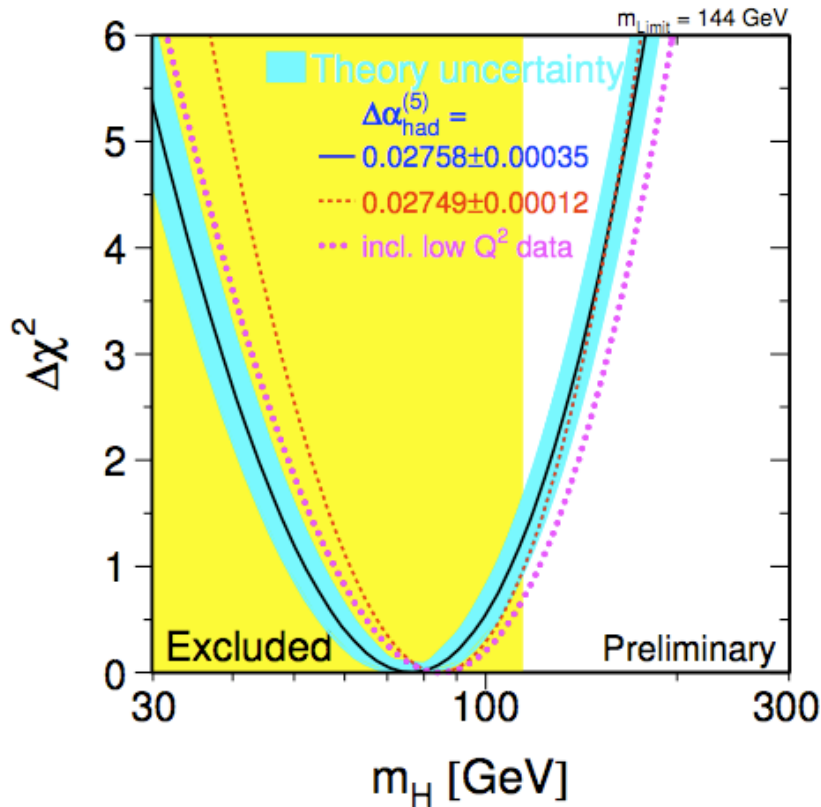
Observation of $Z\gamma \rightarrow \nu\nu\gamma$ at DØ

Yurii Maravin (KSU)



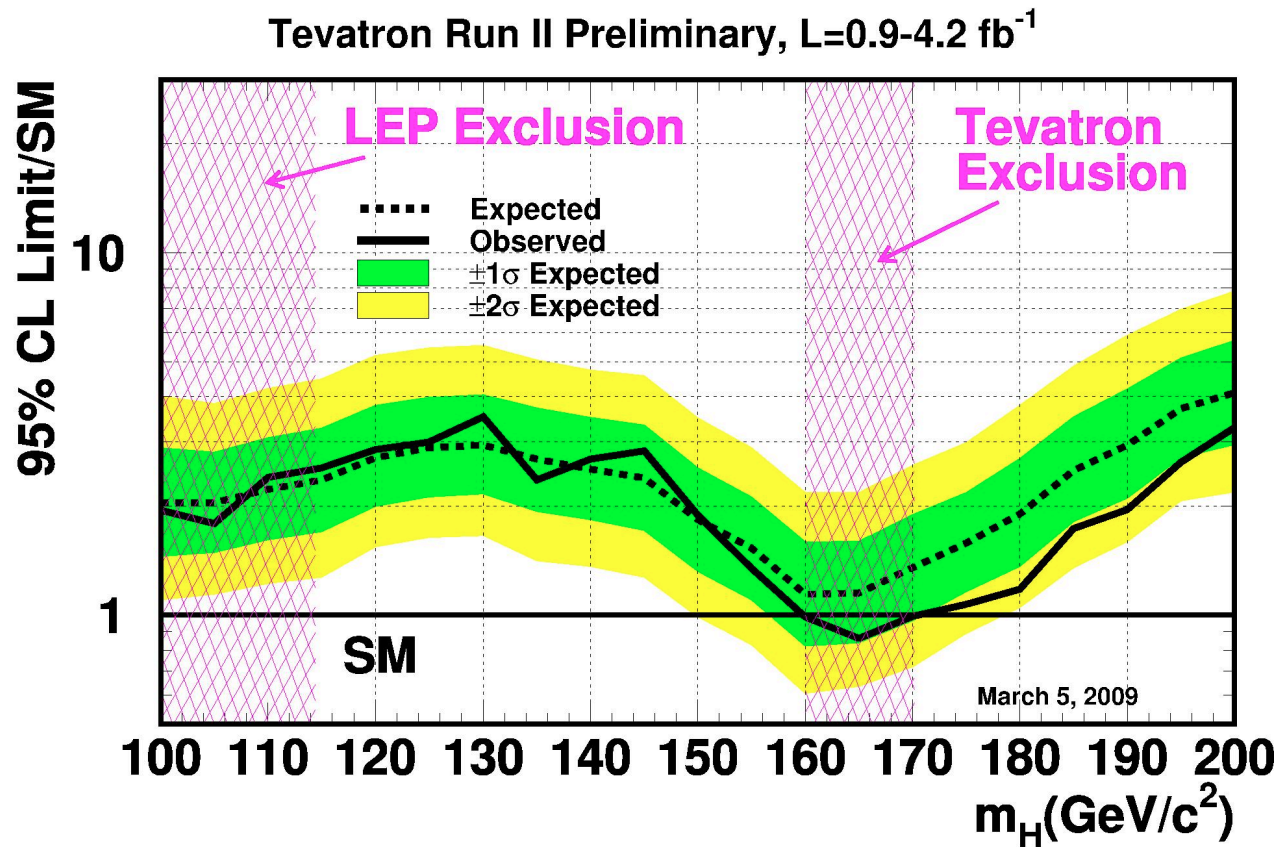
Where is Higgs?

- Artificial add-on to the standard model (is it even there?)
- There is a mounting tension between direct and indirect limits
- SM fits prefer light Higgs that leads to vacuum instabilities at high renormalization scales



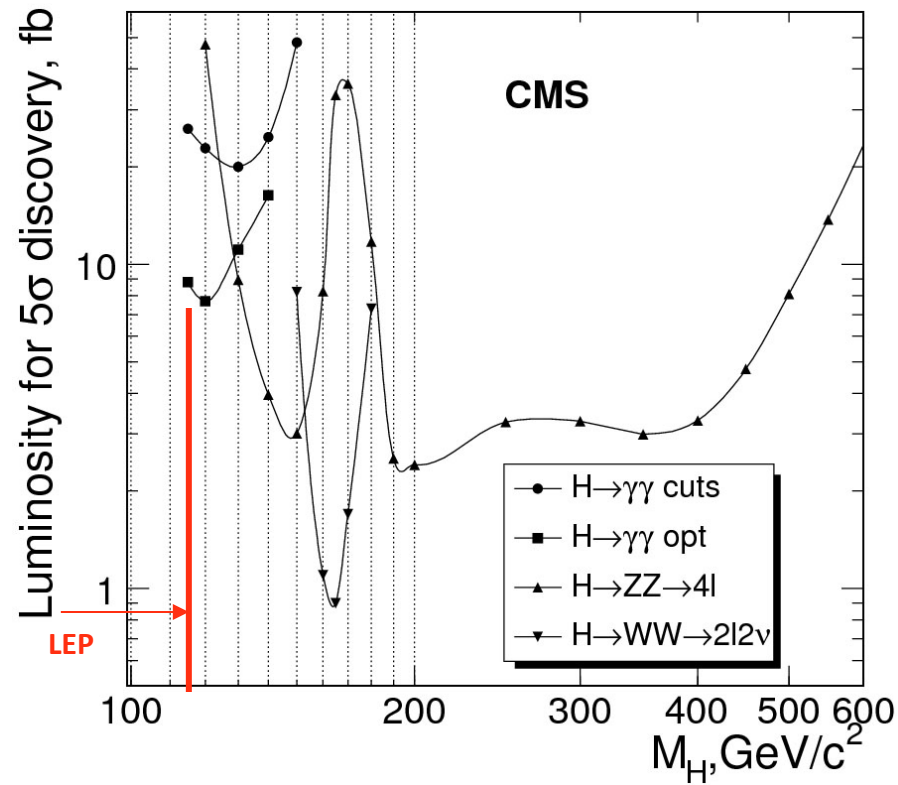
There is no high mass Higgs

- Recent results from Tevatron exclude 160-170 GeV region



Search for Higgs at LHC

- Low mass Higgs discovery channel is $H \rightarrow \gamma\gamma$



Photons?



Yes I see the light!



Physics with photons

- Higgs!
 - Low mass $\gamma\gamma$ resonances – *but it will take time to get there*
- Randall-Sundrum gravitons
 - High mass $\gamma\gamma$ resonances
- Large extra dimensions
 - Excess of high mass $\gamma\gamma$ pairs (virtual gravitons)
 - Mono-photons (photon recoils against a graviton that escapes into extra dimension)
- GMSB SUSY or UED
 - $\gamma\gamma$ + missing ET
- Hidden Valley, GMSB SUSY, 4th generation
 - Long lived particles decaying into photons or electrons
- $Z\gamma/W\gamma$ production
 - Measure gauge boson couplings, search for resonances

In this talk: photons @ DØ

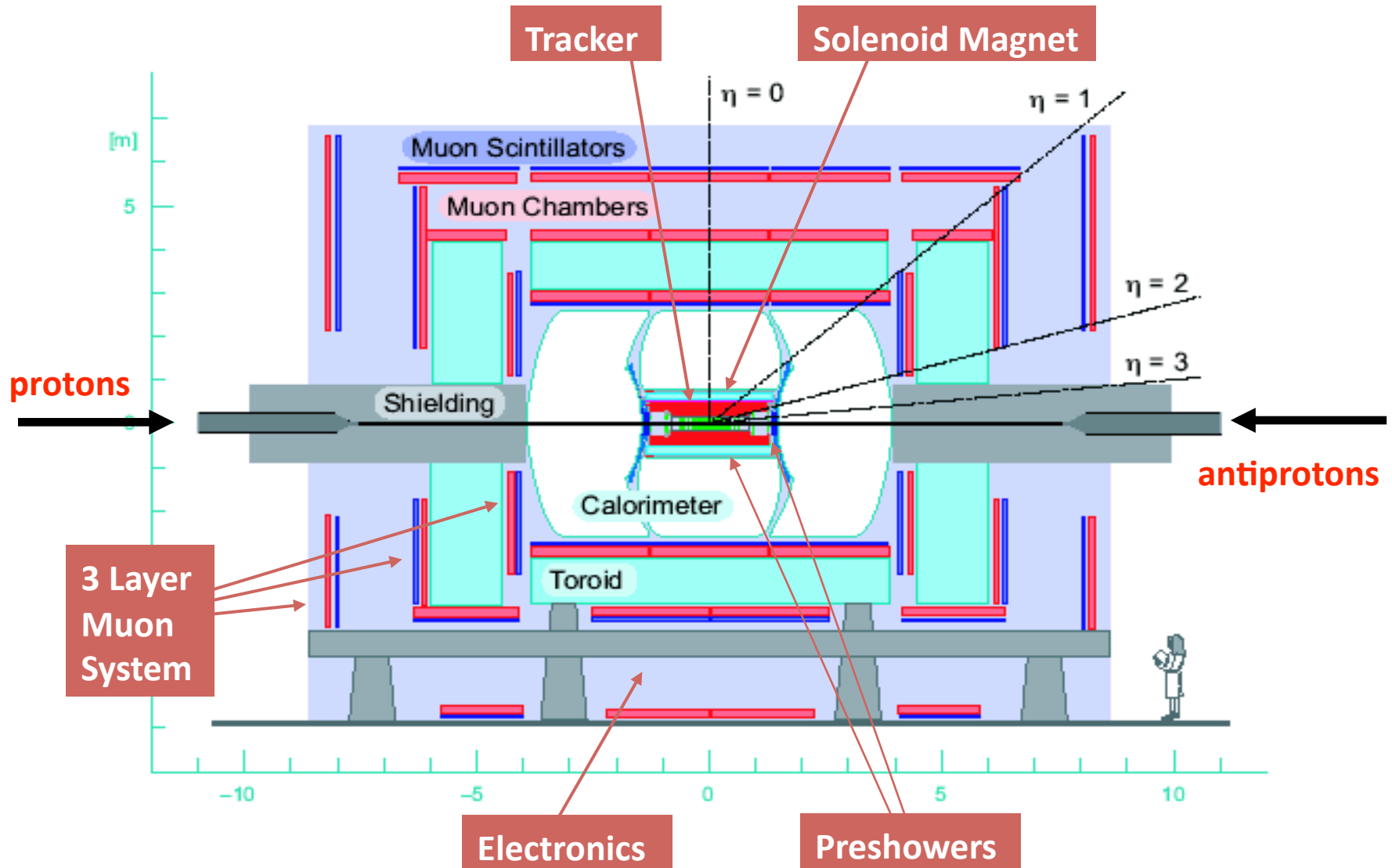
- Example of a photon physics program at DØ
 - Photon identification criteria
 - Synergy with CMS
 - Analysis of mono-photon events with 3.6 /fb
 - Limits on $ZZ\gamma$ and $Z\gamma\gamma$ anomalous couplings
 - Observation of $Z\gamma \rightarrow \nu\nu\gamma$

DØ Collaboration

- 18 countries
- 82 institutions
- 500 authors



DØ Detector

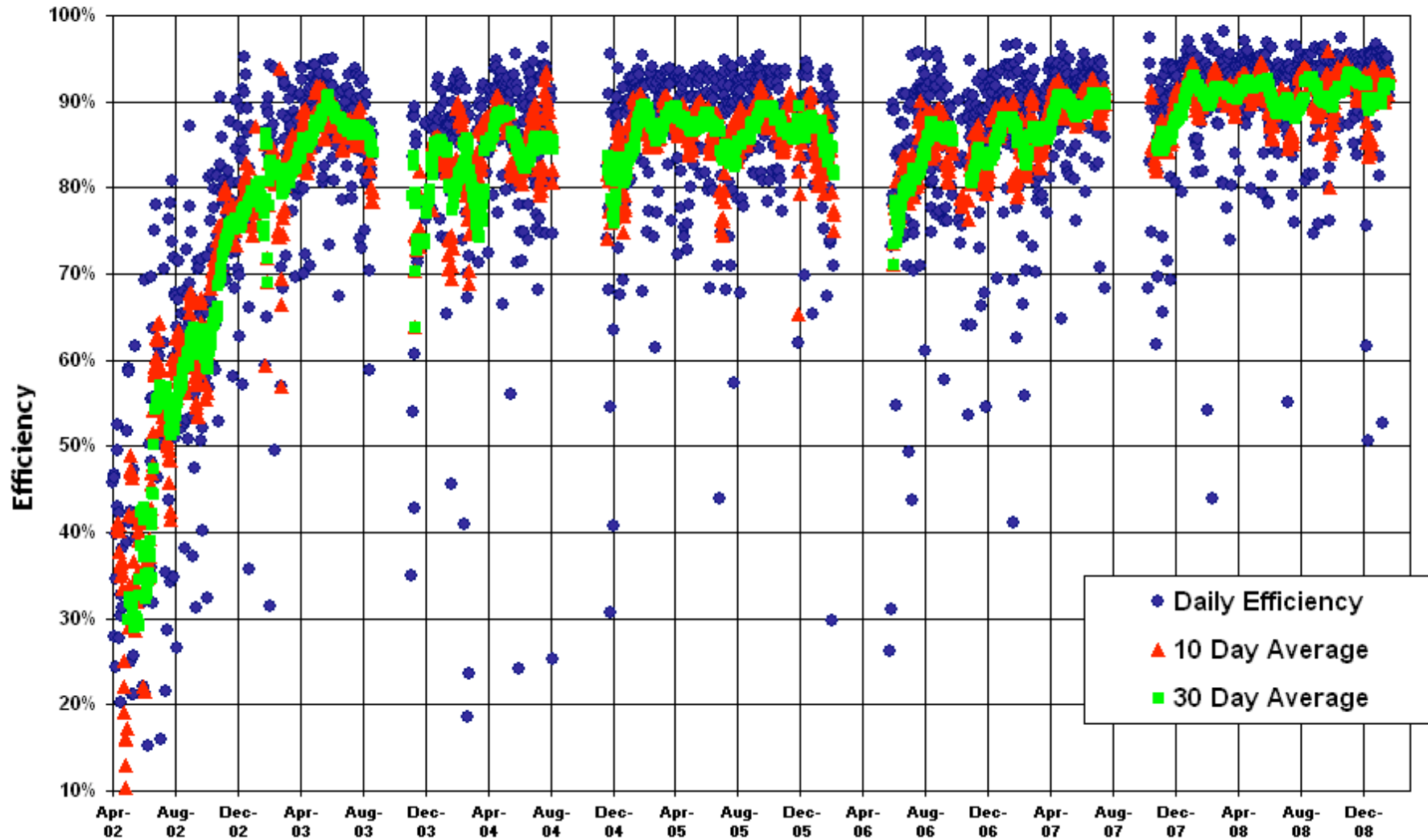


Data taking efficiency

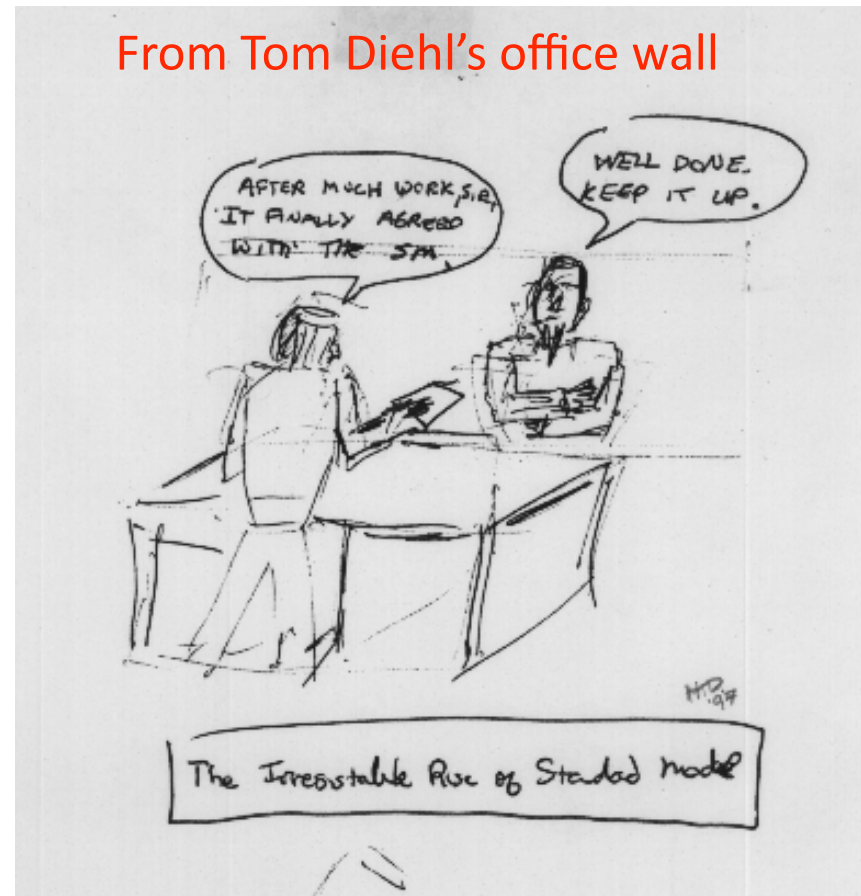


Daily Data Taking Efficiency

19 April 2002 - 8 February 2009



The irresistible rise of the SM

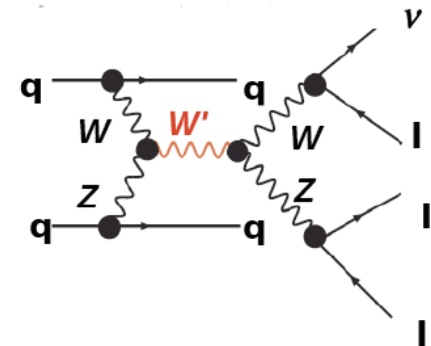
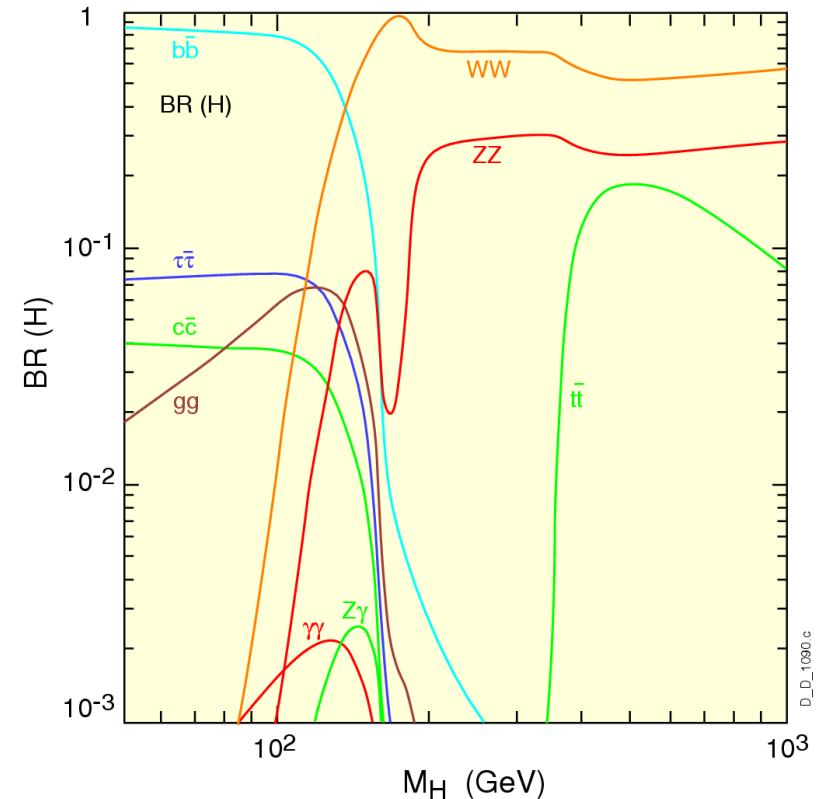


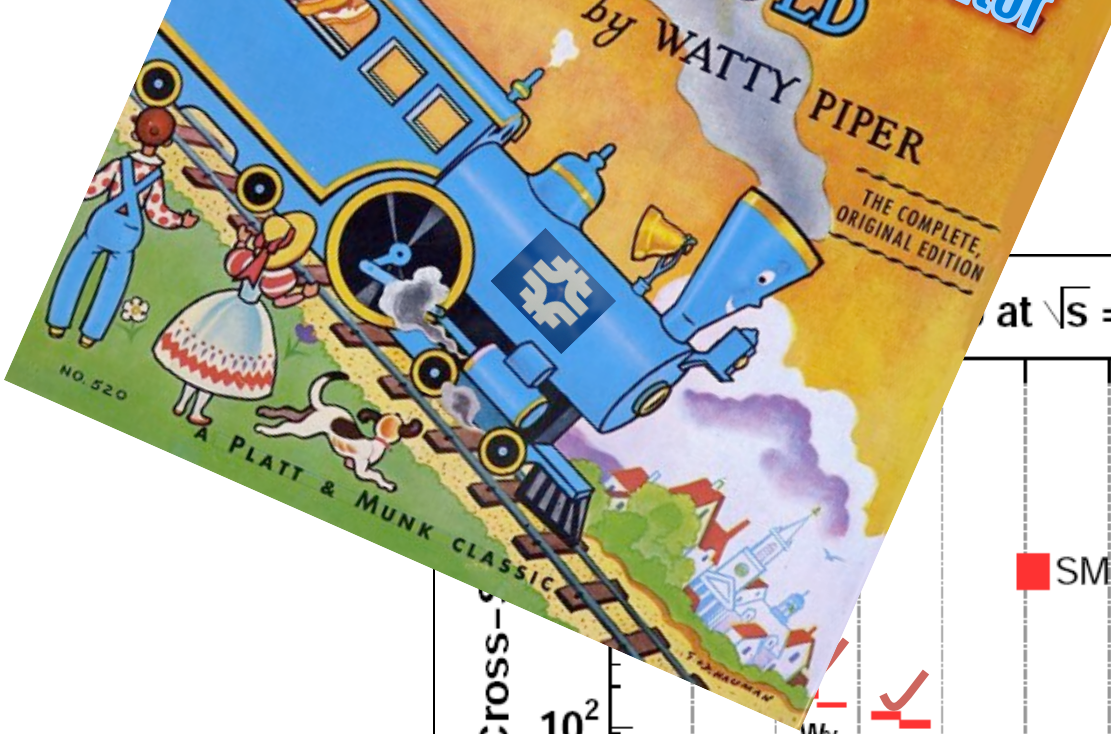
- Wonderful agreement with experimental results
 - What lies ahead?

Diboson physics

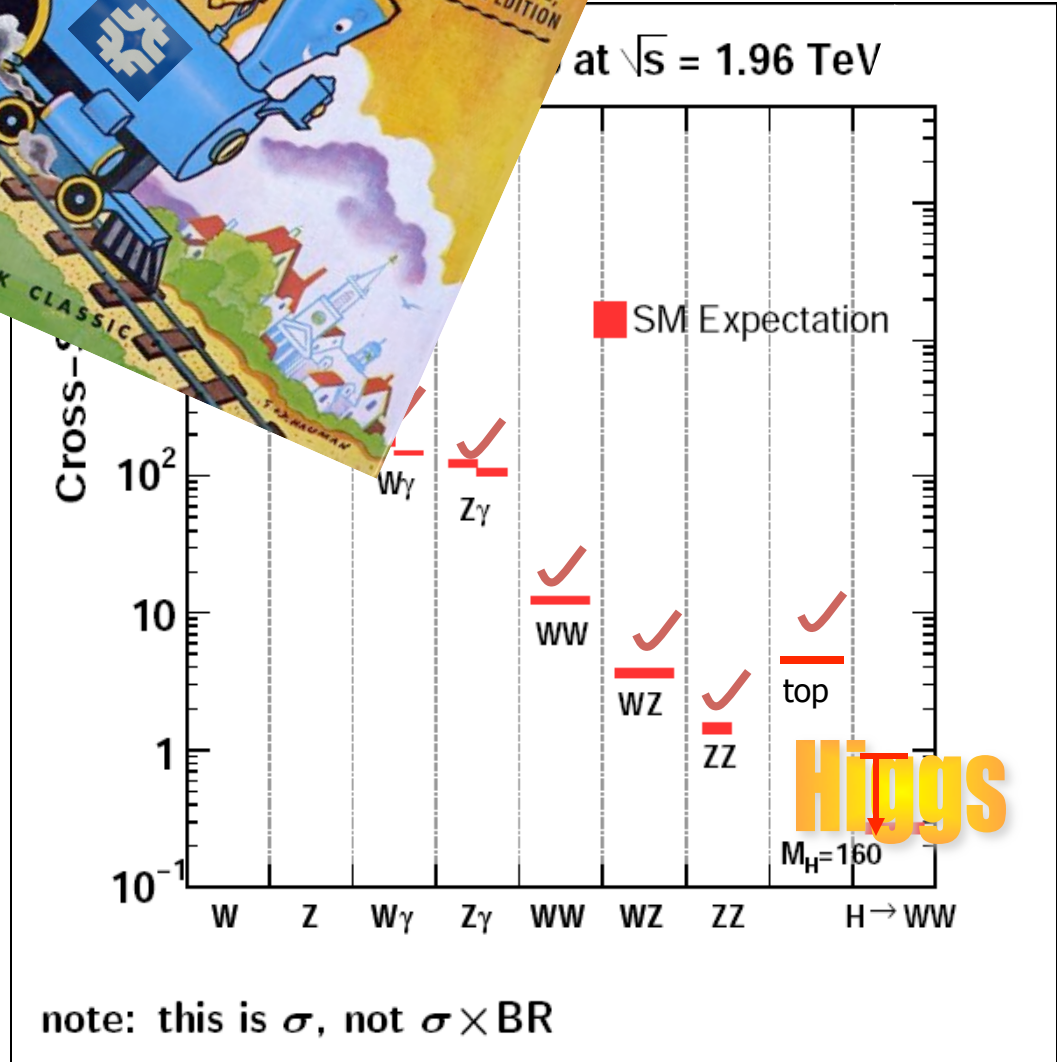
A. Djouadi, J. Kalinowski, M. Spira

- Physics with multiple bosons in the final state
 - Such as WW , WZ , $Z\gamma$, $\gamma\gamma$, ...
- A number of important measurements and searches
 - Cross section
 - Search for resonant production
 - Such as Higgs, or fermiophobic higgs, or whatever...
 - Self-interaction boson couplings are the least well known parameters of the EW sector of the standard model





at $\sqrt{s} = 1.96 \text{ TeV}$

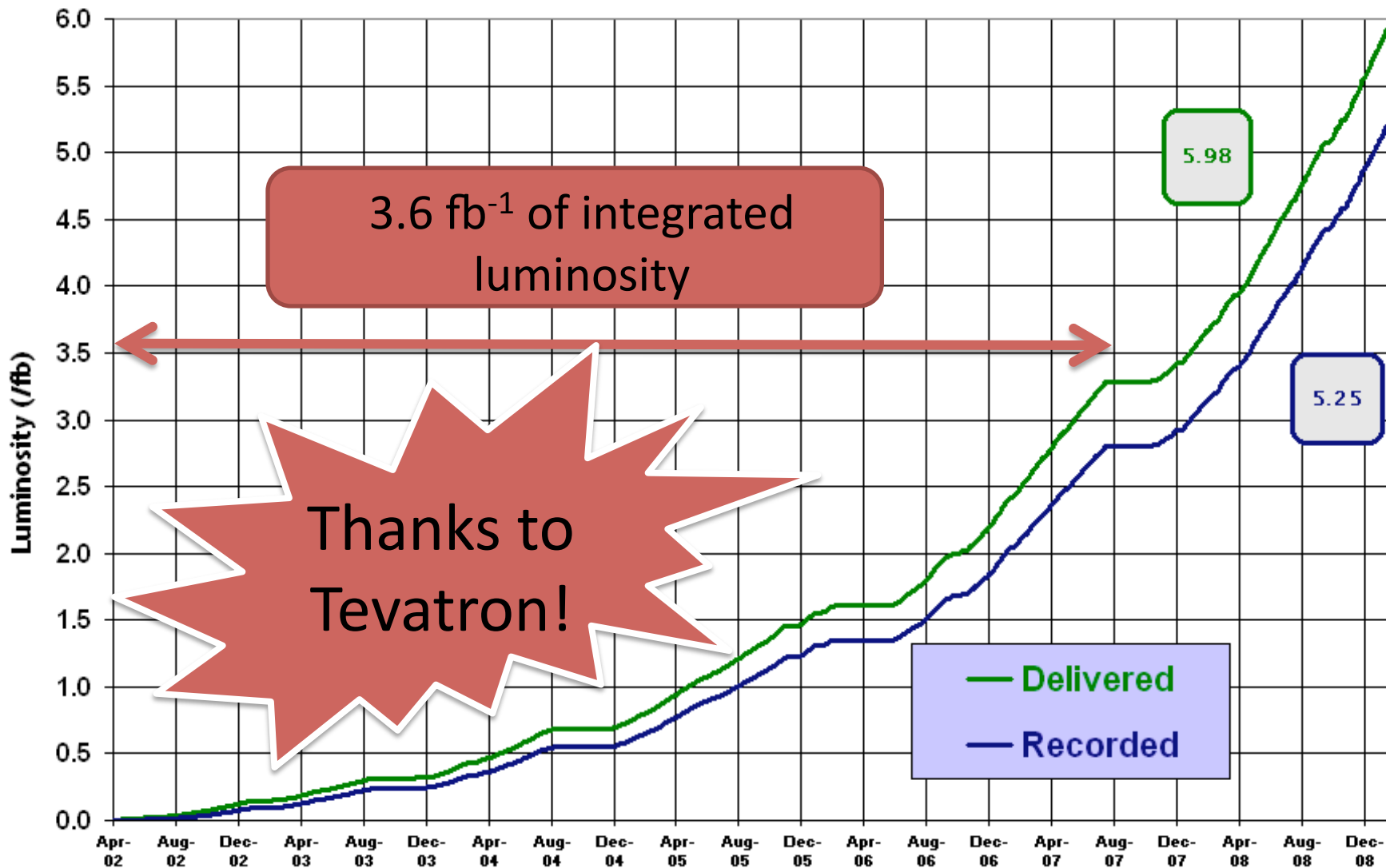


Integrated luminosity



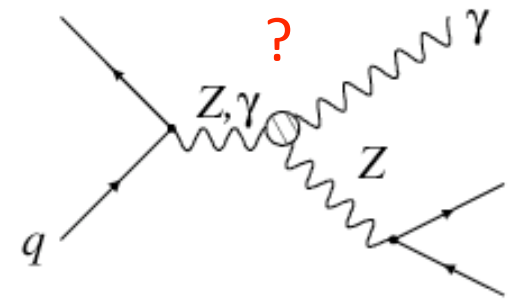
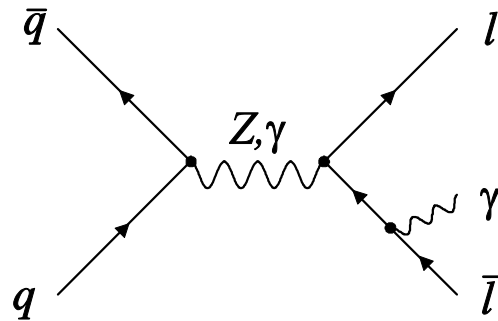
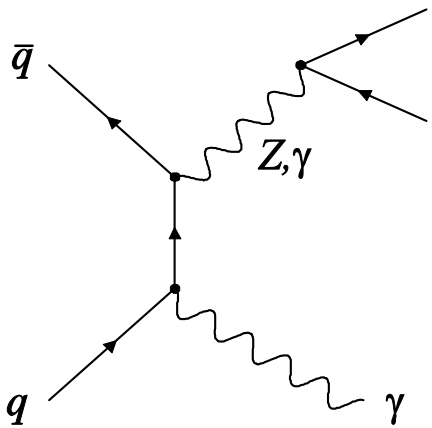
Run II Integrated Luminosity

19 April 2002 - 8 February 2009



Z γ production

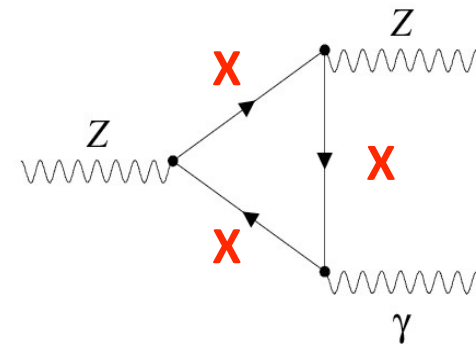
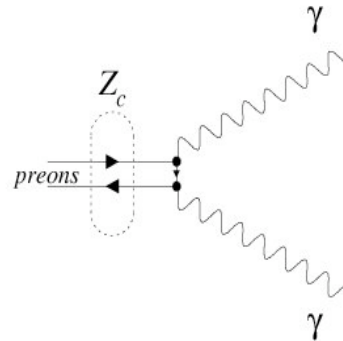
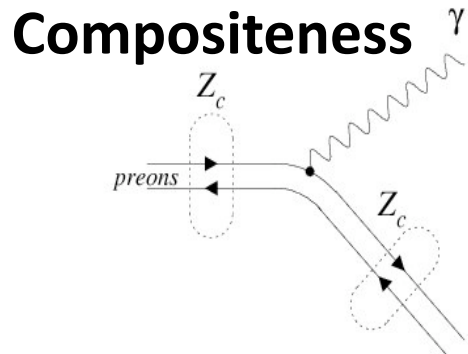
- SM predicts only two tree-level diagrams of Z γ via initial and final state radiation
 - No final state radiation in $\nu\nu\gamma$ final state



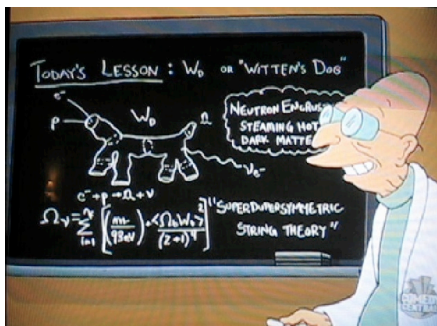
- ZZ γ and Z $\gamma\gamma$ couplings are almost zero
 - QED corrections are at the 10^{-4} level

New phenomena in $Z\gamma$

- Numerous possible extensions of the standard model result in non-zero $ZZ\gamma$ and $Z\gamma\gamma$ couplings



Something else?



- Follow effective Lagrangian approach
 - Parameterize the $ZZ\gamma/Z\gamma\gamma$ vertex in the most general way

Most general parameterization

Baur & Berger, 1993

- $ZV\gamma$ vertex can be parameterized by 8 complex couplings h_i^Z and h_i^γ where i is 1-4
 - h_1, h_2 are CP-odd,
 h_3, h_4 are CP-even
 - Unitarity is violated at high \hat{s} , use form-factor ansatz to enforce good energy behavior

$$d_{Z\tau} = -\frac{e}{\sqrt{2}} \frac{k^2}{M_Z^3} (h_{30}^Z - h_{40}^Z)$$

$$\mu_{Z\tau} = -\frac{e}{\sqrt{2}} \frac{k^2}{M_Z^3} (h_{10}^Z - h_{20}^Z)$$

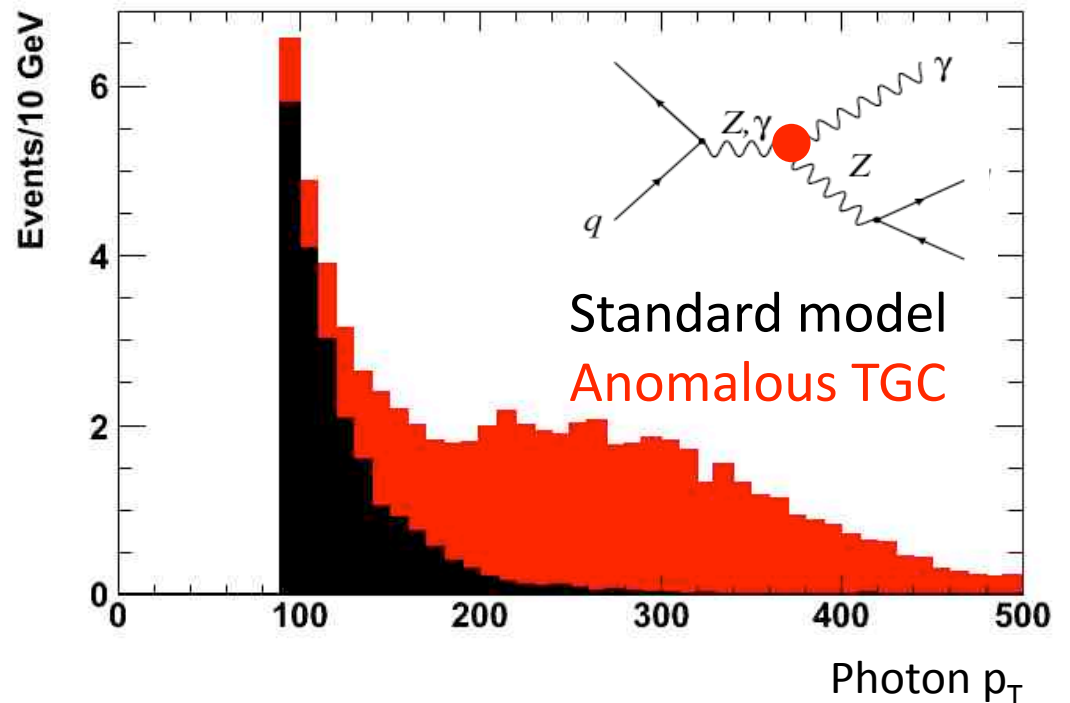
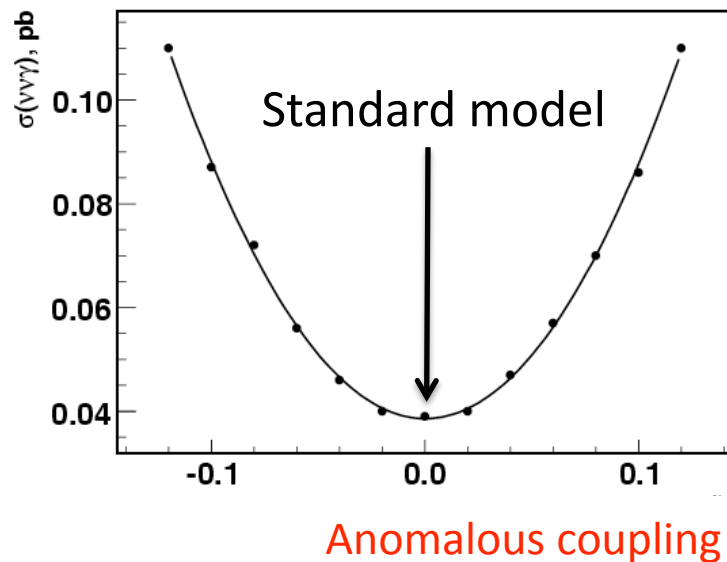
$$h_i^V = \frac{h_{i0}^V}{(1 + \hat{s}/\Lambda^2)^n}$$

Low energy approximation

- Here, Λ is a new physics scale that is responsible for conserving unitarity at high \hat{s}
 - Customary, $n = 3$ for $h_{1,3}^V$ and, and 4 for $h_{2,4}^V$

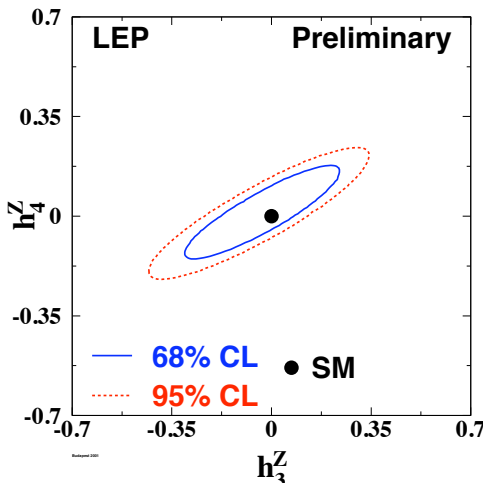
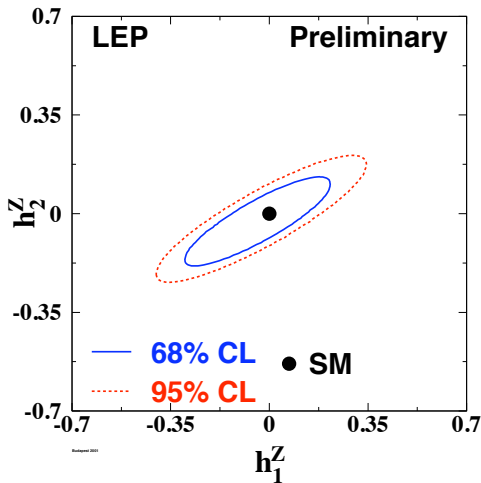
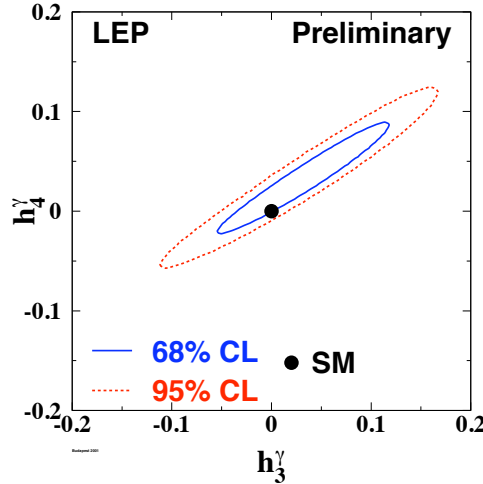
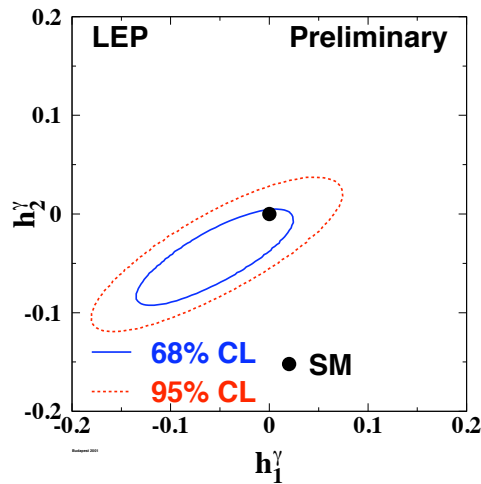
Effect of anomalous coupling

- Any non-zero coupling result in increase of the cross section and harder p_T spectrum of the photon and Z
 - Produced from Baur MC 4-vector output (LO)



Previous results on $Z\gamma$: LEP

LEP EWWK 2003, Preliminary



$$-0.056 < h_1^\gamma < 0.055$$

$$-0.045 < h_2^\gamma < 0.025$$

$$-0.049 < h_3^\gamma < -0.008$$

$$-0.002 < h_4^\gamma < 0.034$$

$$-0.13 < h_1^Z < 0.13$$

$$-0.078 < h_2^Z < 0.071$$

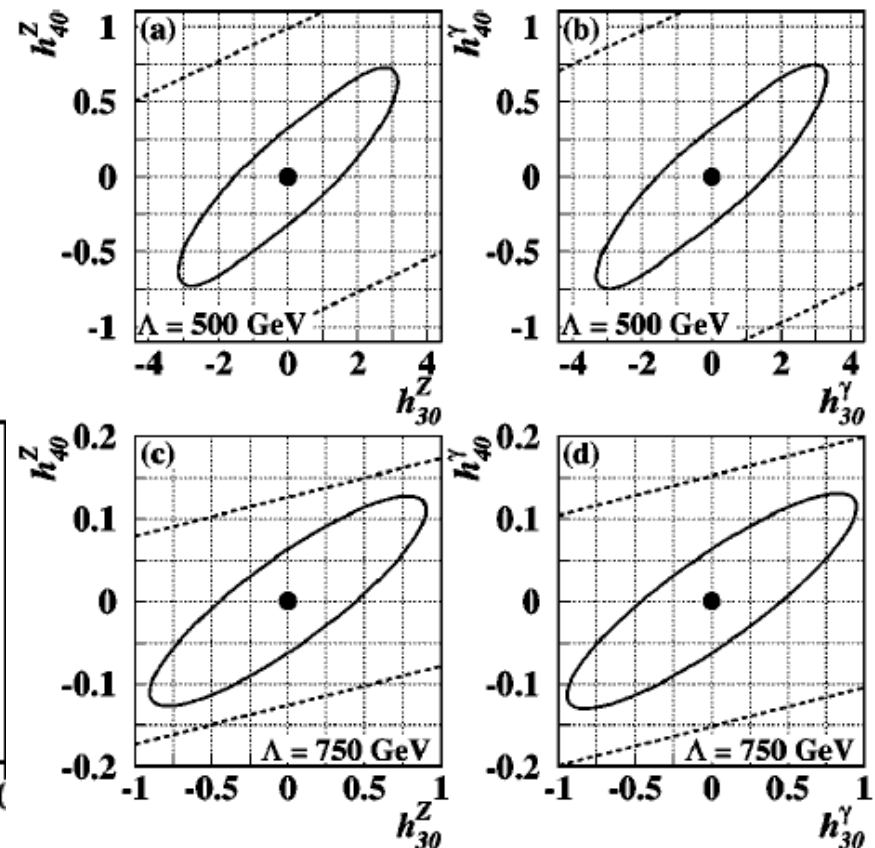
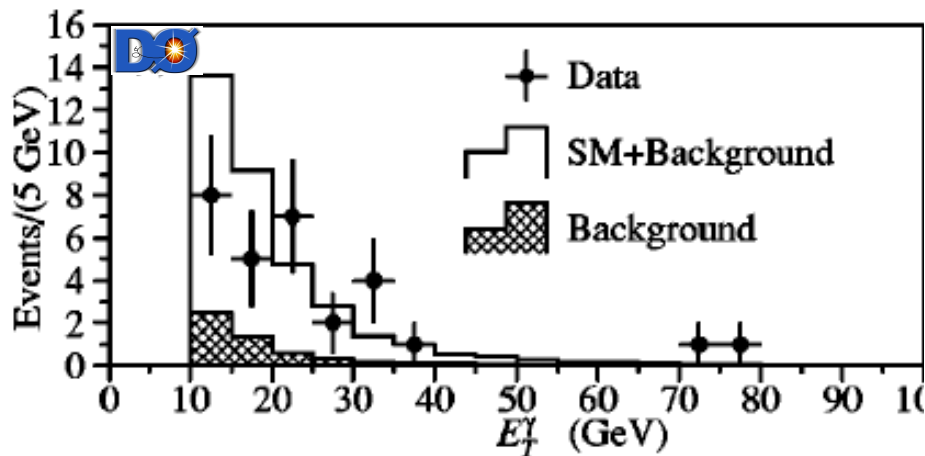
$$-0.20 < h_3^Z < 0.07$$

$$-0.05 < h_4^Z < 0.12$$

- Measured $ZZ\gamma$ and $Z\gamma\gamma$ couplings agree with SM at $10^{-1} - 10^{-2}$ level

Previous Tevatron results $Z\gamma \rightarrow ll\gamma$

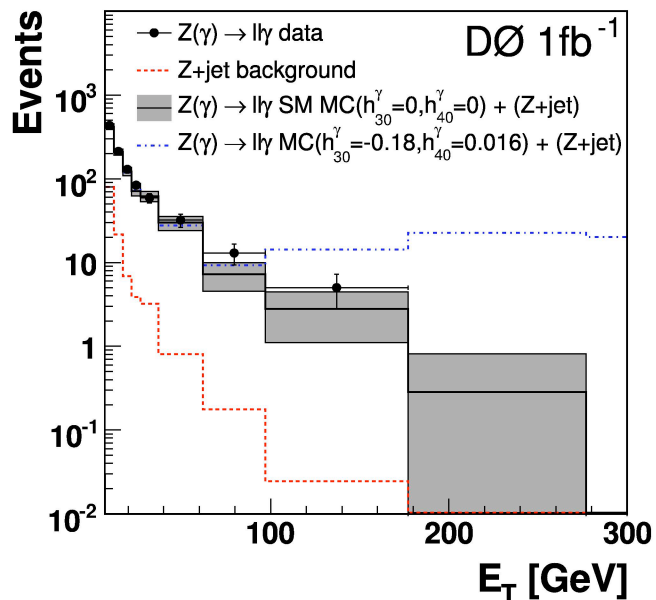
- $D\bar{D}$ set limits $Z\gamma \rightarrow ll\gamma$ in Run I
 - Observed 29 events
 - The p_T^γ spectrum agree with standard model prediction



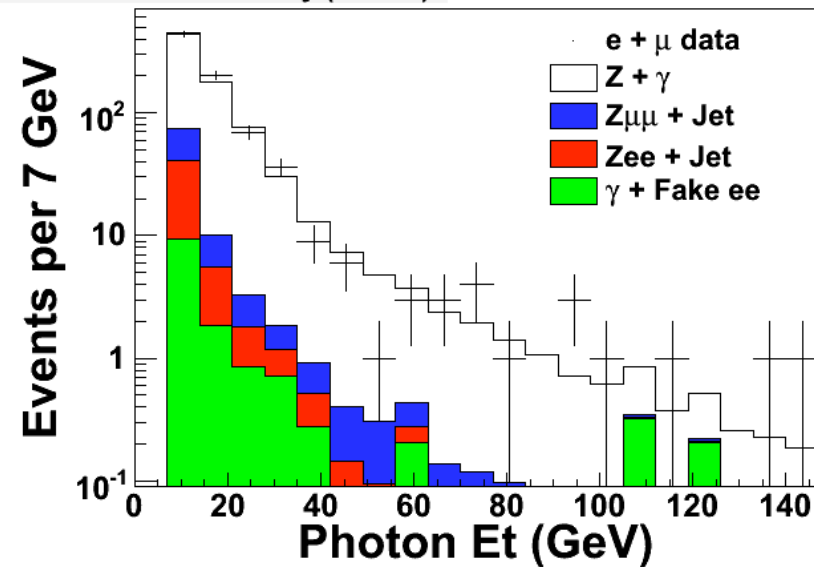
Agreement with SM at $10^{-1} - 10^{-2}$ level

Previous Run II results on $Z\gamma \rightarrow ll\gamma$

- Both CDF and DØ performed extensive studies of the $Z\gamma$ production in $Z\gamma \rightarrow ll\gamma$
 - Both cross section and p_T^γ spectrum agree with standard model prediction
 - Limits are $|h_3^V| < 0.085$ $|h_4^V| < 0.0047$



CDF Run II Preliminary (2.0 fb⁻¹)



Agreement with SM at $10^{-1} - 10^{-3}$ level

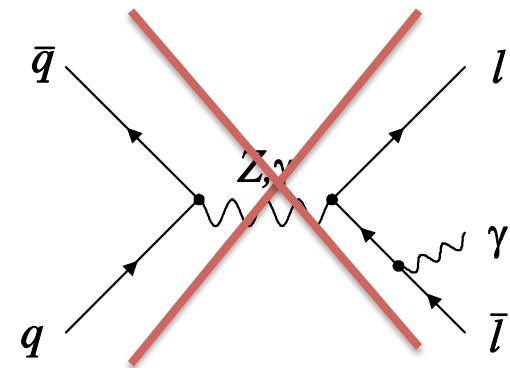
Can we do better?



- Precision is still dominated by statistics
 - Sensitivity is in the tail of the p_T^γ distribution
- Major limiting factors:
 - Three particle final state
 - Low $Z \rightarrow ll$ branching fraction

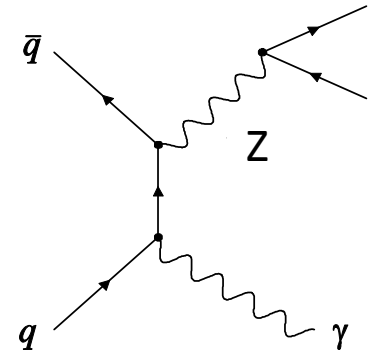
- Challenging alternative: $Z\gamma \rightarrow \nu\nu\gamma$

- Much higher acceptance
- No FSR processes!
- Neutrino branching fraction is three times the $Z \rightarrow ll$



Precision should double for the $\nu\nu\gamma$ channel alone!

$$Z\gamma \rightarrow \nu\nu\gamma$$



- Very challenging
 - Has not been seen at Tevatron!
- Final state is a single photon and a missing transverse energy (MET) consistent with $Z \rightarrow \nu\nu$ production
 - Backgrounds:
 - QCD processes and W production ($e \rightarrow \gamma$)
 - Beam-halo, bremsstrahlung cosmic muons
- A crucial ingredient to this analysis is identification of photons

Photon identification

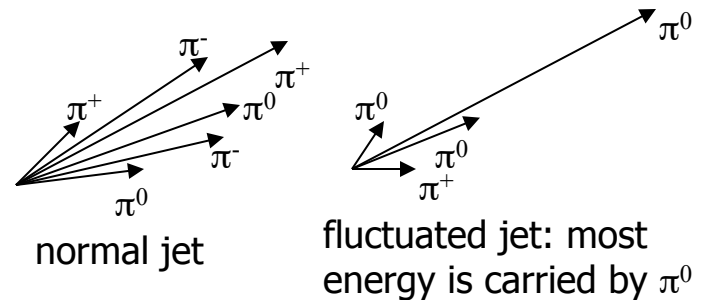
- Unconverted photons do not have much redundancy: just a shower in the calorimeter

– Handles to suppress backgrounds:

- Isolation in tracker and hadron calorimeter
- Shower profile should be consistent with that of a photon

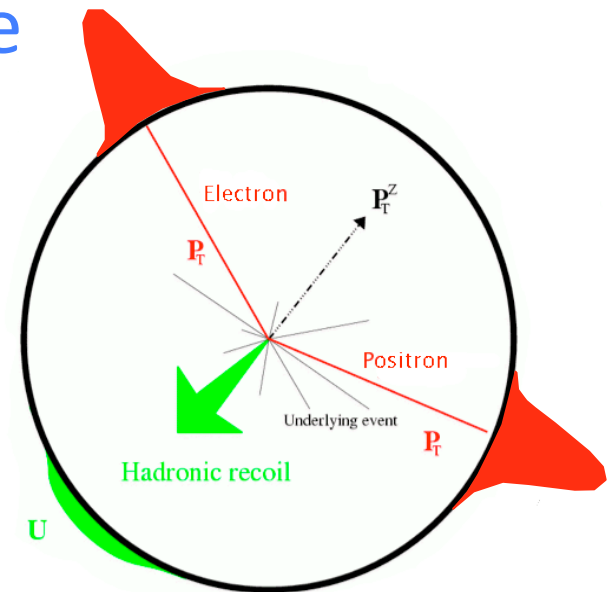
- No track pointing to the photon candidate

– No additional hits in the vicinity of a photon candidate is consistent with not reconstructed track



Calibrating photons

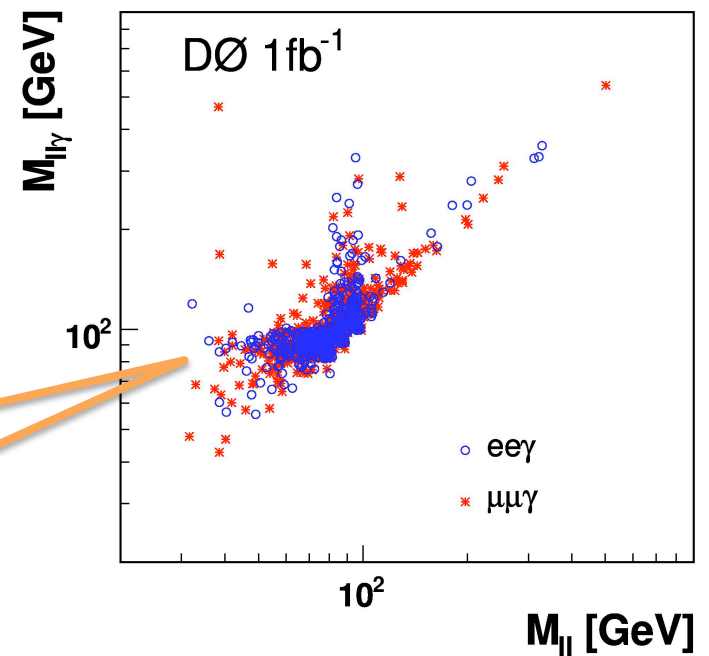
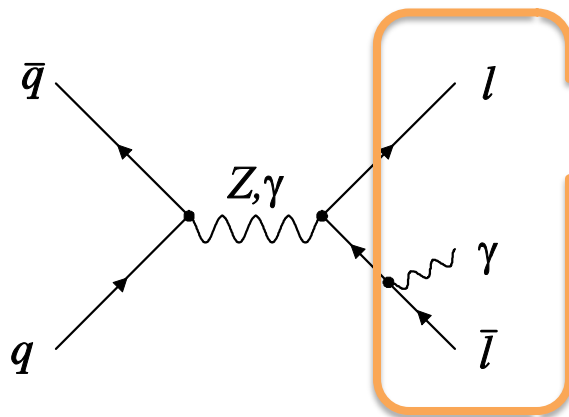
- We must find Higgs to produce a photon calibration signal!
- Use data for calibration: $Z \rightarrow ee$
 - Use Monte Carlo to describe the difference between photon and electron shower well
 - Tune Monte Carlo so that it describes electrons well
 - Cross checked in data with FSR $Z\gamma$ events



$Z \rightarrow ll\gamma$ as a standard candle

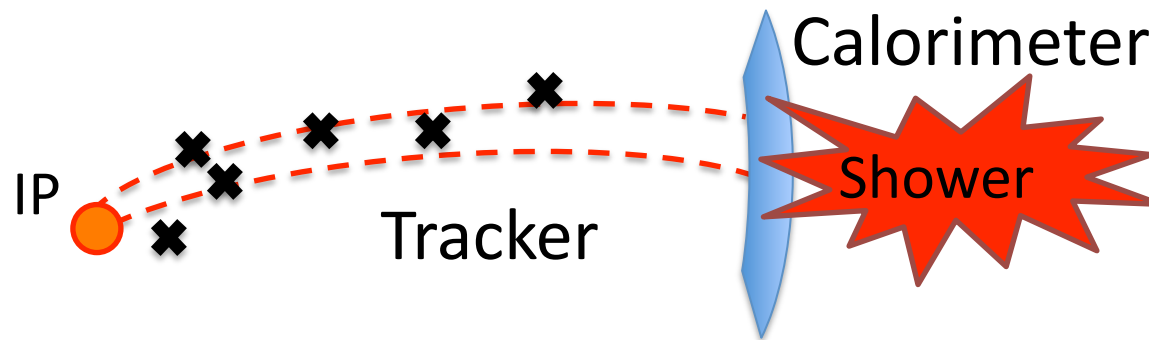
- FSR $Z\gamma$ is the cleanest source of photons

– One can use FSR production to make photon sample very clean and to infer photon energy scale!



Suppression of electrons

- In addition to the standard matching algorithm in ϕ and η space, require tracker hits density along the EM trajectory to be consistent with noise
 - Hit density and resolution is determined in data



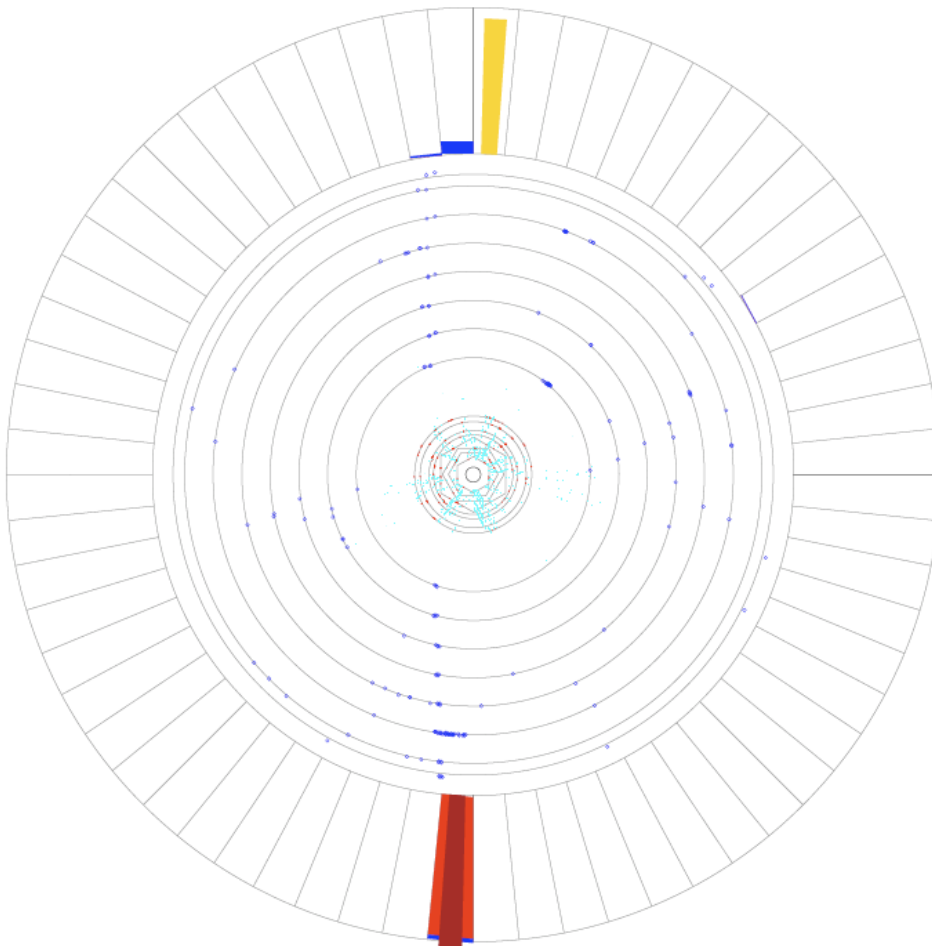
- Improves electron track matching efficiency and decreases the $e \rightarrow \gamma$ misidentification rate by a factor of four!

Non-pointing background

- One of the major backgrounds to the $\nu\nu\gamma$ final state is cosmic muon that radiated a photon in the calorimeter

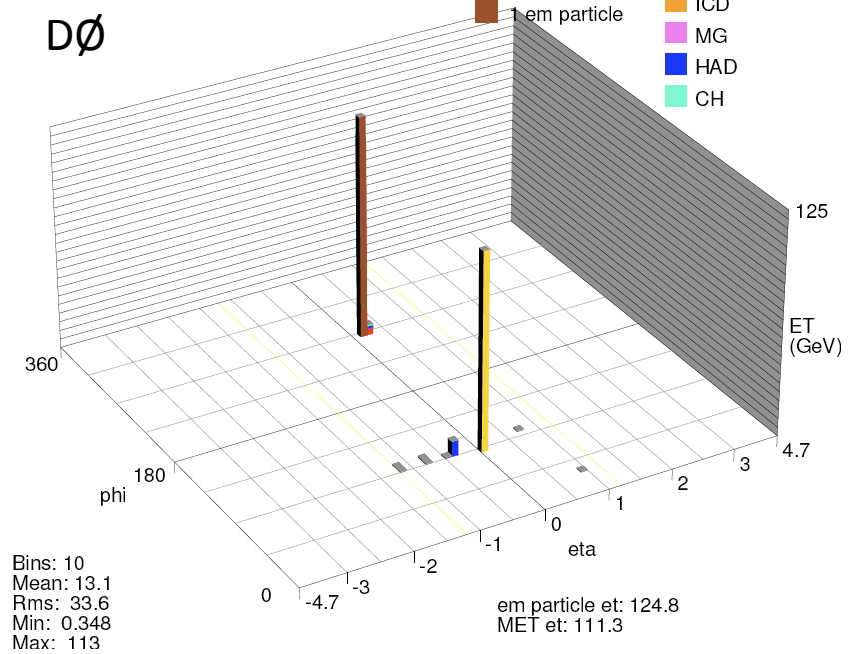
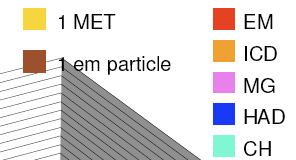
ET scale: 120 GeV

DØ



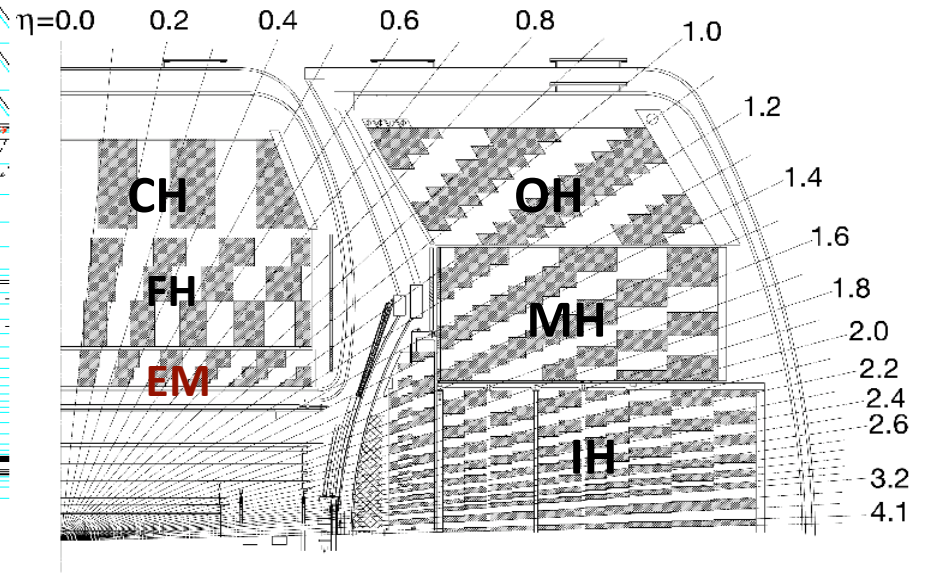
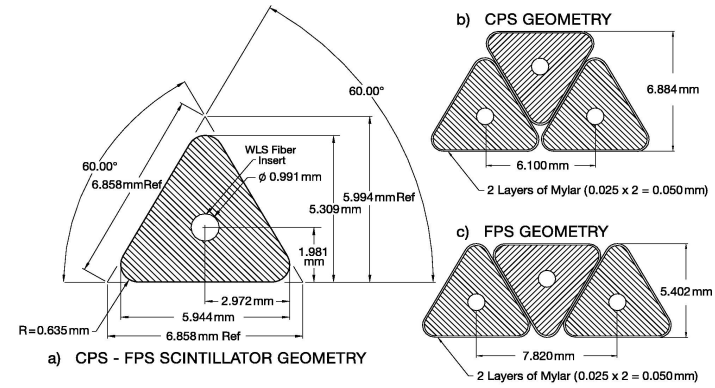
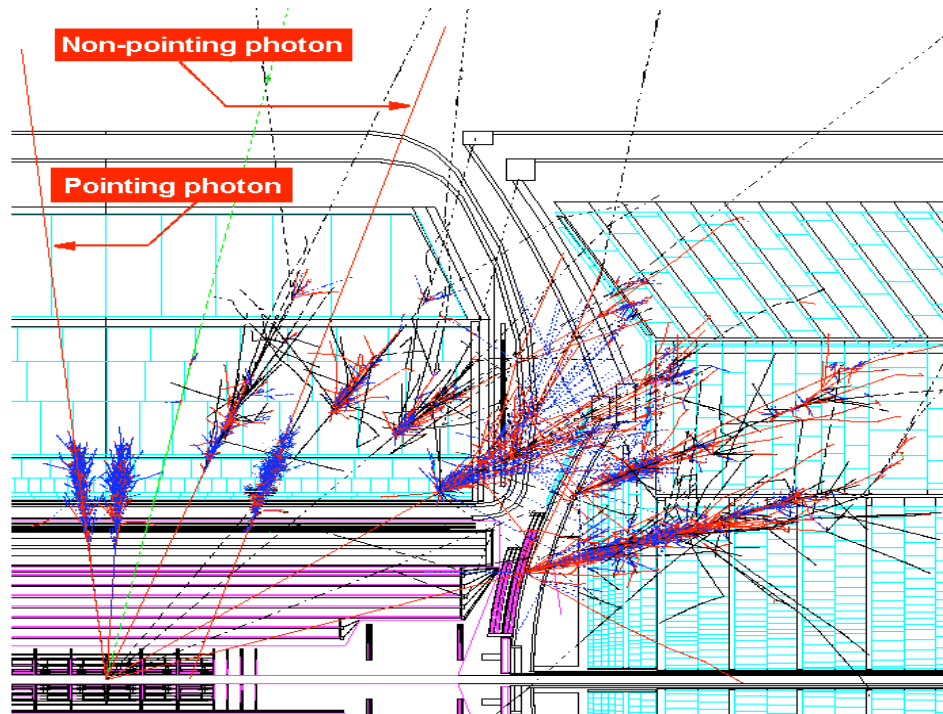
Triggers:

DØ



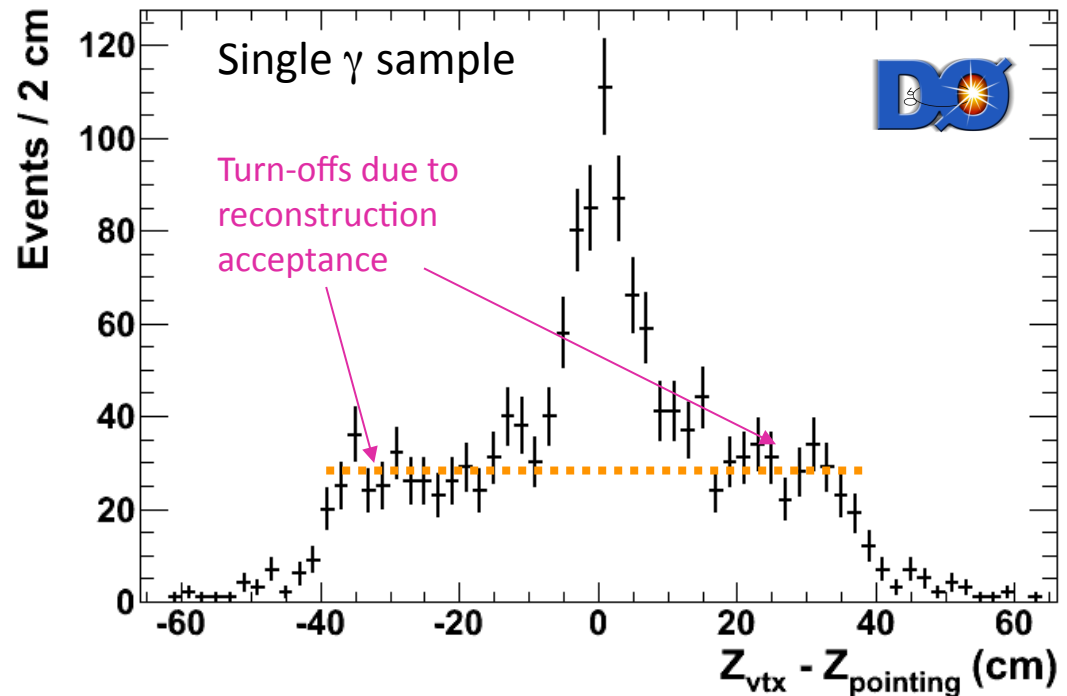
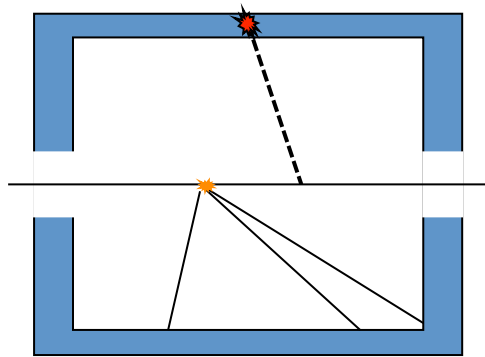
DØ Calorimetry

- DØ calorimeter is highly segmented
 - Use it to pinpoint the shower direction in 3D!

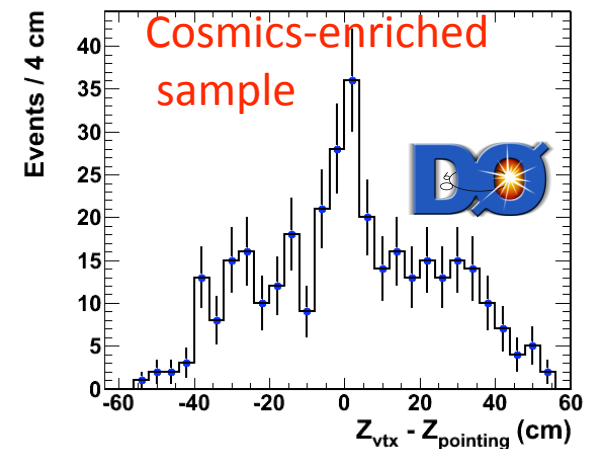
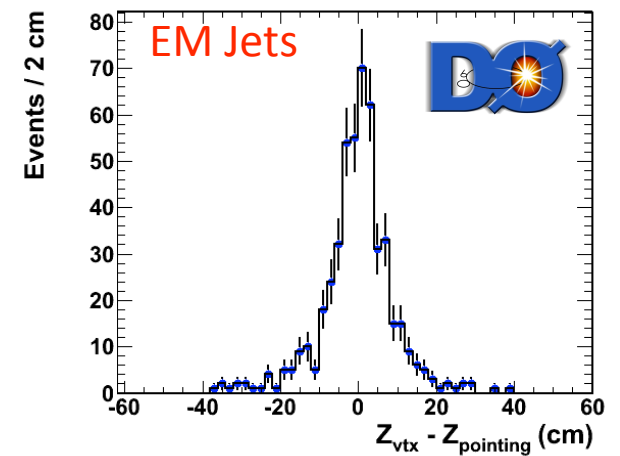
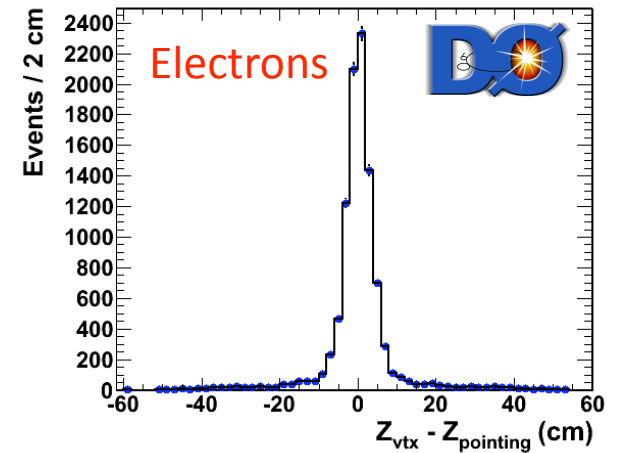


Pointing in z-coordinate

- Cosmics and beam-halo photons would not point to the primary vertex of the event
 - Exploit this to identify and reject non-collision backgrounds

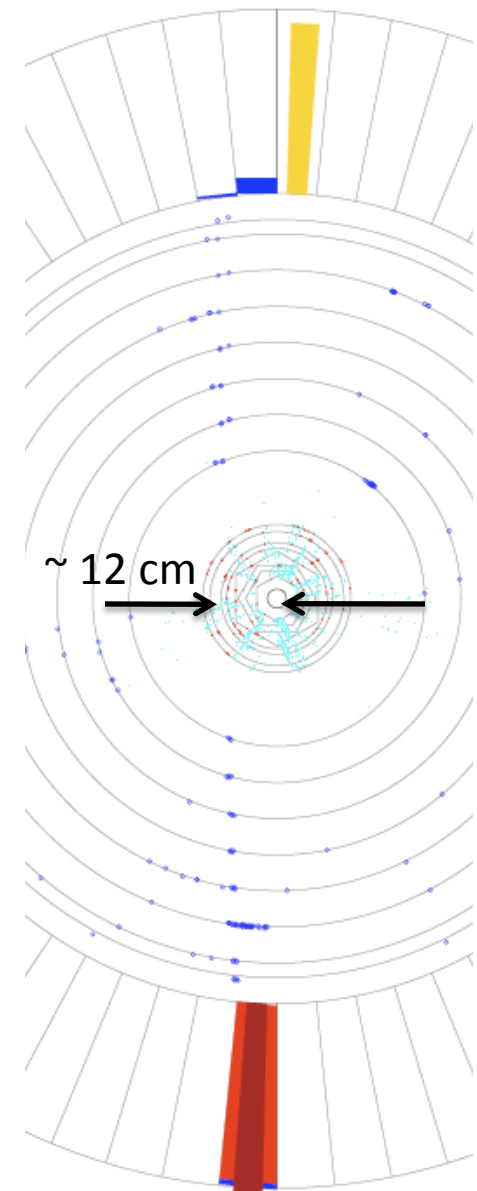
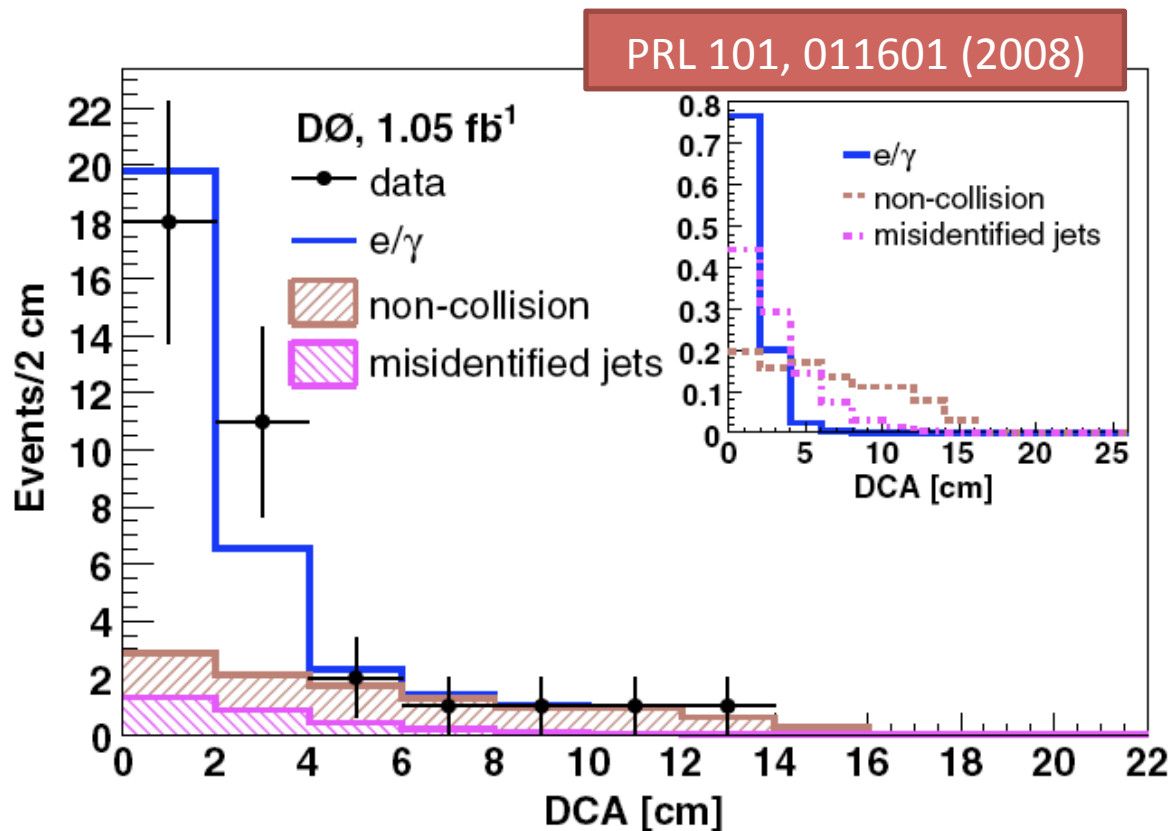


- Δz distributions from data
 - Extremely important for the MET measurement
 - Electrons and photons have narrow showers and thus small Δz resolution
 - Use $Z \rightarrow ee$ sample
 - Misidentified jets have wider shower profile and thus larger pointing resolution
 - Use “bad” EM sample and γ +jet
 - Non-collision is pretty flat
 - Use cosmics-enriched data



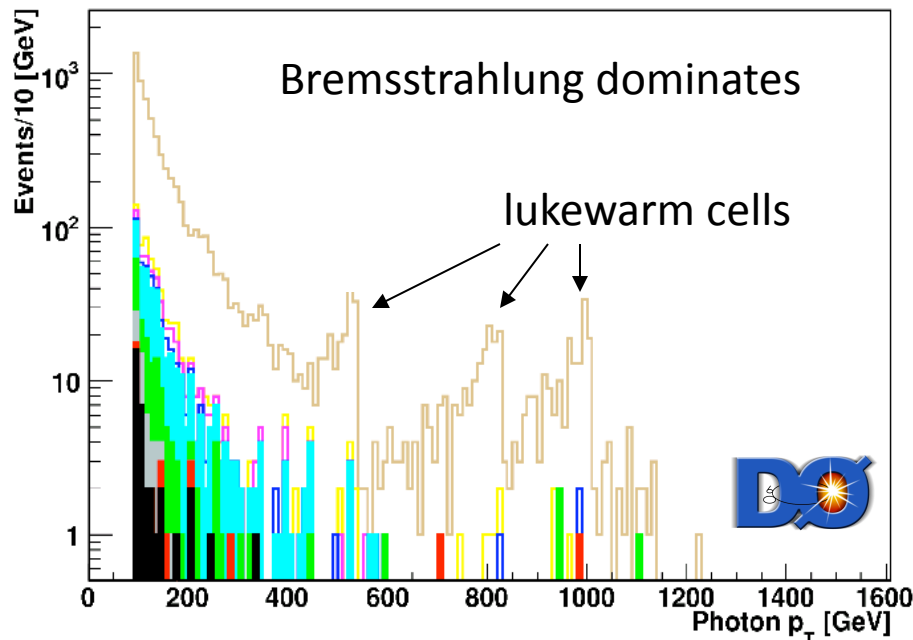
Identification of non-pointing γ



- Determine DCA distributions from data



Applying the pointing algorithm

- Select data sample with photon $p_T > 90$ GeV
 - Sample is dominated by muon Bremsstrahlung
 - Applying pointing requirements reduces cosmics and beam-halo considerably!



-  Kinematics requirements only
-  Applying CPS information

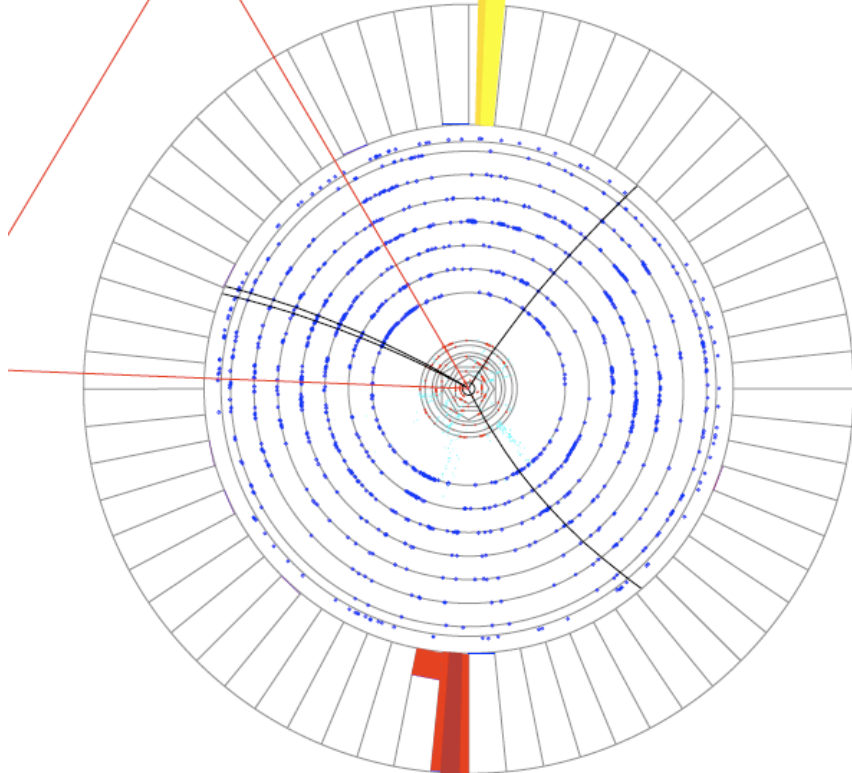
Selecting $Z\gamma \rightarrow \nu\nu\gamma$

- Select events with single EM triggers
 - Fully efficient at 40 GeV
- Require event to have missing $E_T > 70$ GeV
- Require a clean event
 - No jets with $p_T > 15$ GeV, isolated tracks, cosmic rays, muons...
- Photon candidate has $p_T > 90$ GeV, $|\eta| < 1.1$, isolated, and have shower profile consistent with that of a photon

$\nu\nu\gamma$ candidate event

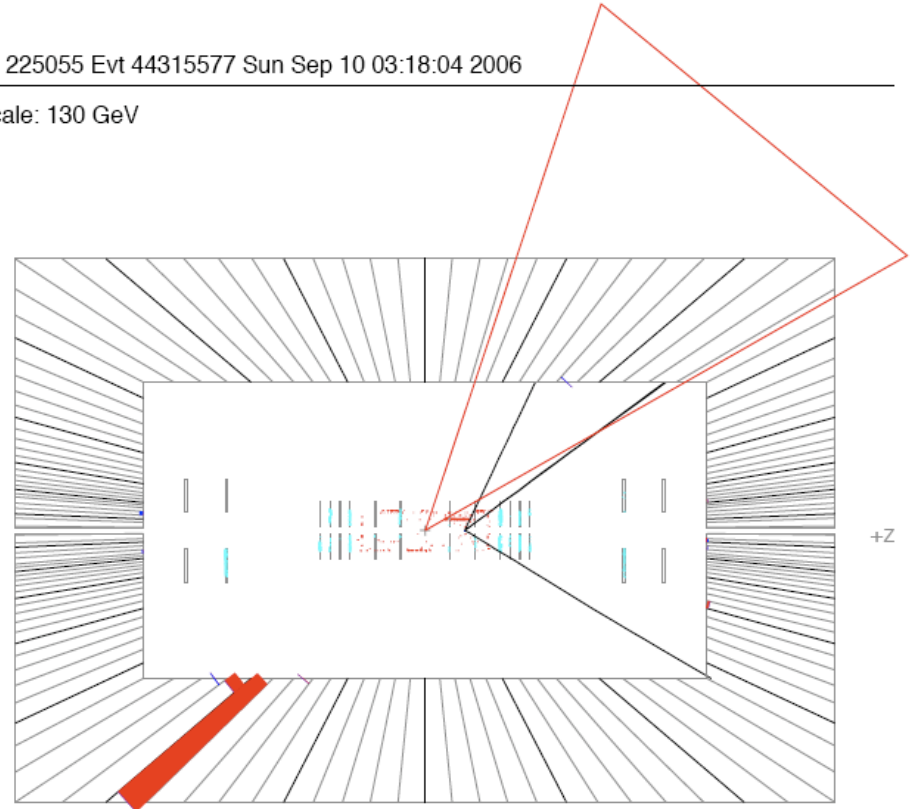
Run 225055 Evt 44315577 Sun Sep 10 03:18:04 2006

ET scale: 114 GeV

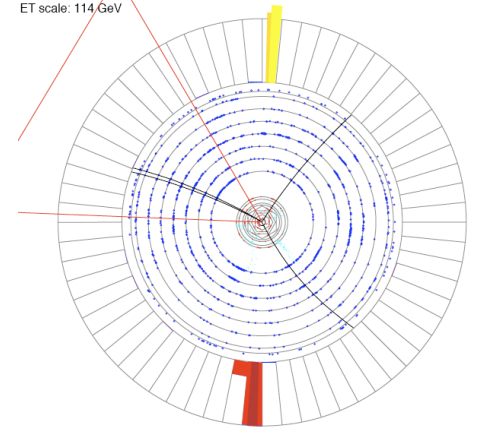


Run 225055 Evt 44315577 Sun Sep 10 03:18:04 2006

E scale: 130 GeV



Backgrounds



- $W \rightarrow e\nu$: electron is misidentified as a photon
 - Estimated from W data sample
- **Non-collision**: cosmic or halo muon Bremsstrahlung
 - Estimated from DCA template method
- W or Z + jet: jet is misidentified as a photon
 - Estimated from the DCA template method
- $W + \gamma \rightarrow l\nu + \gamma$: lepton is lost
 - Small, estimated from Monte Carlo simulation

| | |
|----------------------|--|
| $W \rightarrow e\nu$ | 9.67 ± 0.30 (stat.) ± 0.48 (syst.) |
| <i>non-collision</i> | 5.33 ± 0.39 (stat.) ± 1.91 (syst.) |
| $W/Z + jet$ | 1.37 ± 0.26 (stat.) ± 0.91 (syst.) |
| $W\gamma$ | 0.90 ± 0.07 (stat.) ± 0.12 (syst.) |

Simulation

- There is a number of Monte Carlo generators on the market: use Baur generator
 - The generator of choice for CDF, DØ, CMS,...
- Use both NLO and LO generators to simulate the process kinematics and acceptance and calculate theoretical cross section
 - NLO generator is used to calculate NLO k -factor
 - Detector simulation is done by using Parameterized Monte Carlo Simulation (PMCS)
 - Very fast and reliable!

Cross section measurement

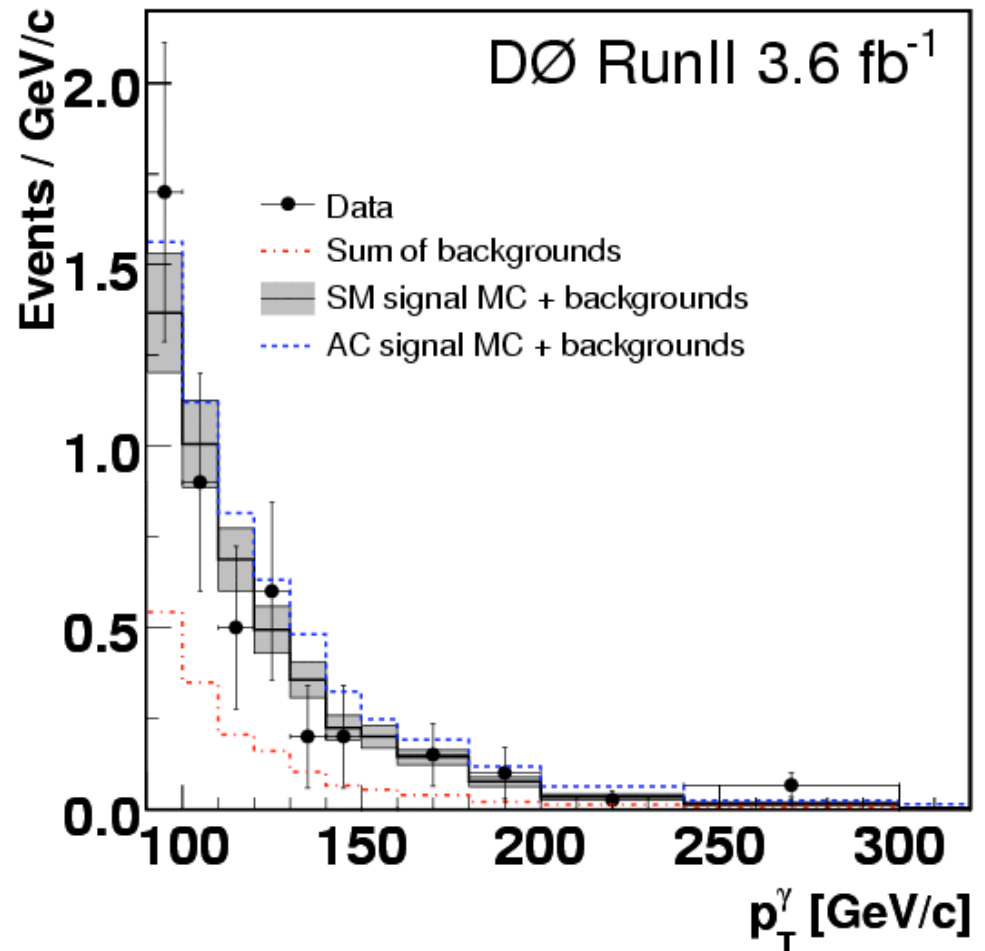
- Using 3.64 fb^{-1} of data we observe 51 $Z\gamma \rightarrow \nu\nu\gamma$ candidate events with an estimated $17.3 \pm 0.6 \text{ (stat.)} \pm 2.3 \text{ (syst.)}$ background events

$$\sigma \cdot \text{Br}(Z \rightarrow \nu\nu) = 32 \pm 9 \text{ (stat+syst)} \pm 2 \text{ (lumi)} \text{ fb}$$

- Theory predicts $39 \pm 4 \text{ fb}$ (NLO)
- Perform 10^8 pseudo-experiments with background-only hypothesis to find out that the probability for for background to fluctuate up is 3.1×10^{-7} which corresponds to 5.1σ
 - First observation of $Z\gamma \rightarrow \nu\nu\gamma$ at the Tevatron!

Measuring $ZZ\gamma/Z\gamma\gamma$ couplings

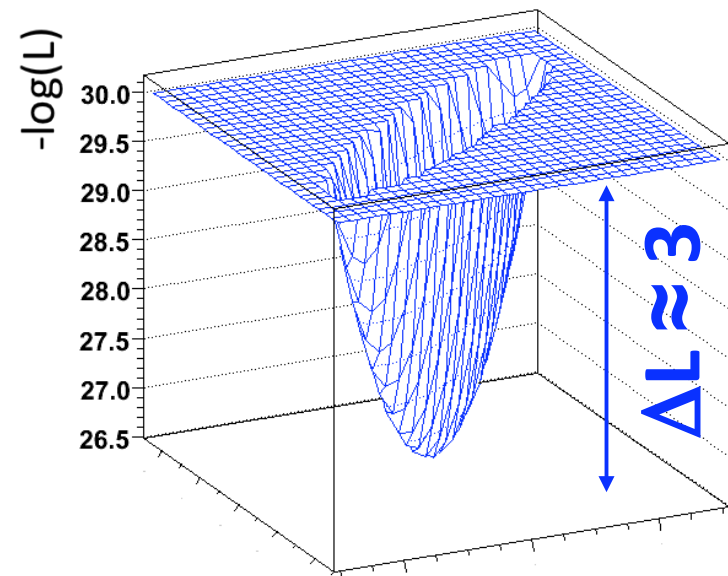
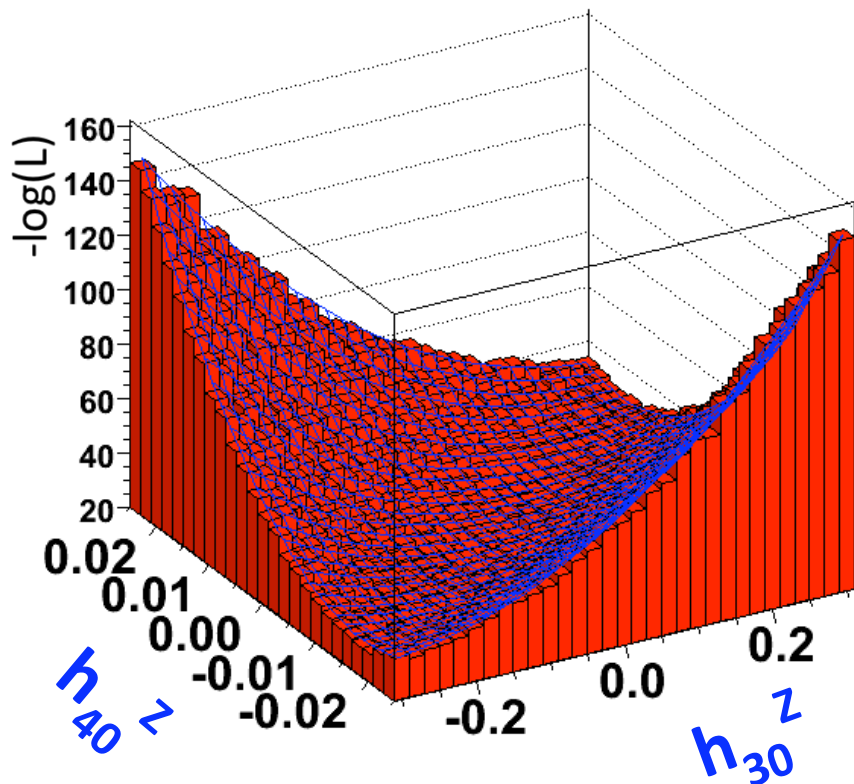
- Data are consistent with standard model production
 - Proceed with setting limits on anomalous couplings



Setting limits on $ZZ\gamma$ and $Z\gamma\gamma$

- The observable (photon p_T) is sensitive to the strength of the coupling
 - We present results for CP-even couplings: sensitivity to h_1 is similar to h_3 , and similarly h_2 is similar to h_4
- Generate a 2D grid of simulation with different values of couplings (h_{30} and h_{40})
 - Set CP-odd couplings to zero

- Assume Poisson statistics for signal and Gaussian statistics for systematic uncertainties and background, calculate the likelihood of data to be described by aTGC simulation and background
 - Repeat for every point of the generated grid



Limits on anomalous couplings

Submitted to PRL

- Set 1D limits by setting all the other aTGCs to zero

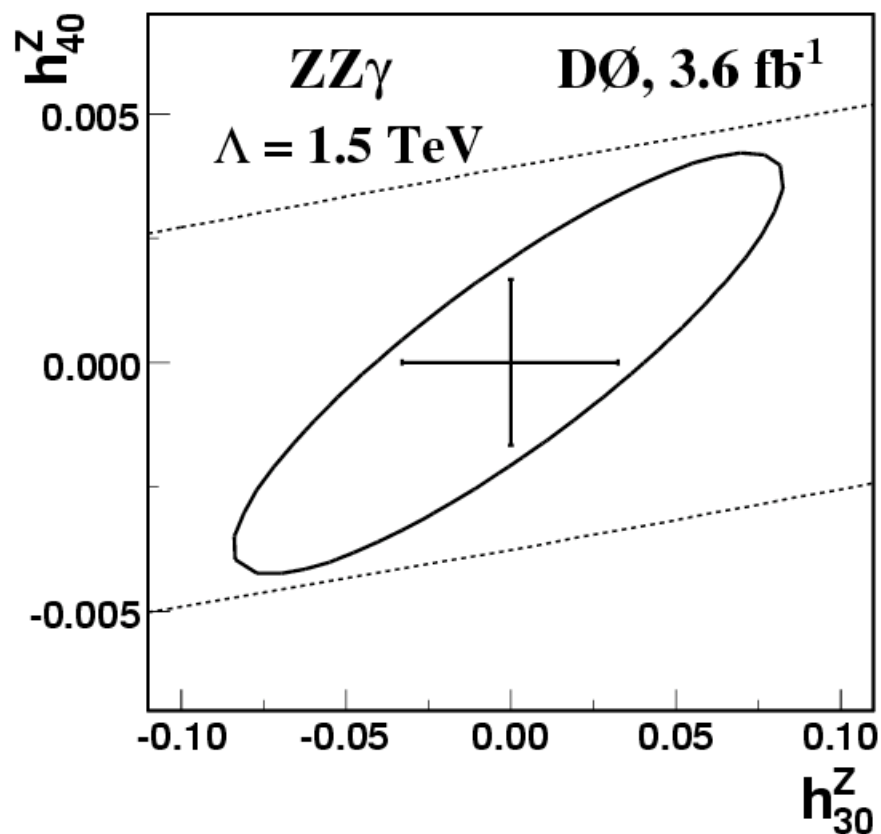
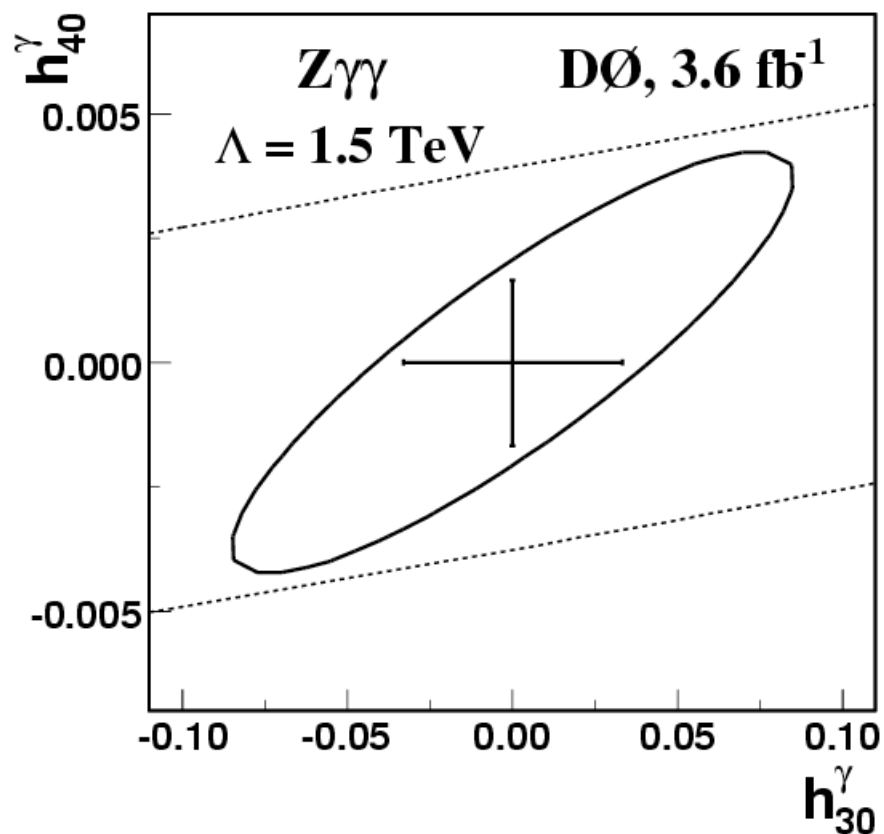
| Old DØ result! | $\Lambda = 1.2 \text{ TeV}$ | $\Lambda = 1.5 \text{ TeV}$ |
|--|--|--|
| $Z\gamma \rightarrow ll\gamma \ 1 \text{ fb}^{-1}$ | $ h_{30}^\gamma < 0.085 \ h_{40}^\gamma < 0.0054$ | $ h_{30}^\gamma < 0.079 \ h_{40}^\gamma < 0.0036$ |
| $Z\gamma \rightarrow ll\gamma \ 1 \text{ fb}^{-1}$ | $ h_{30}^Z < 0.083 \ h_{40}^Z < 0.0054$ | $ h_{30}^Z < 0.075 \ h_{40}^Z < 0.0037$ |
| $Z\gamma \rightarrow \nu\nu\gamma \ 3.6 \text{ fb}^{-1}$ | $ h_{30}^\gamma < 0.042 \ h_{40}^\gamma < 0.0029$ | $ h_{30}^\gamma < 0.037 \ h_{40}^\gamma < 0.0020$ |
| $Z\gamma \rightarrow \nu\nu\gamma \ 3.6 \text{ fb}^{-1}$ | $ h_{30}^Z < 0.041 \ h_{40}^Z < 0.0029$ | $ h_{30}^Z < 0.036 \ h_{40}^Z < 0.0020$ |
| Zγ combination | $ h_{30}^\gamma < 0.038 \ h_{40}^\gamma < 0.0025$ | $h_{30}^\gamma < 0.033 \ h_{40}^\gamma < 0.0017$ |
| Zγ combination | $ h_{30}^Z < 0.037 \ h_{40}^Z < 0.0025$ | $h_{30}^Z < 0.033 \ h_{40}^Z < 0.0017$ |

This result!

- Best limits from Tevatron!

Limits on anomalous couplings

- The most probable values of the $ZZ\gamma$ and $Z\gamma\gamma$ couplings is at the standard model predictions



Comparison with LEP

- These results: $|h_{30}^V| < 0.033$, $|h_{40}^V| < 0.0017$
 - Similar results for CP-odd couplings

- LEP results

$$-0.056 < h_1^\gamma < 0.055$$

$$-0.13 < h_1^Z < 0.13$$

$$-0.045 < h_2^\gamma < 0.025$$

$$-0.078 < h_2^Z < 0.071$$

$$-0.049 < h_3^\gamma < -0.008$$

$$-0.20 < h_3^Z < 0.07$$

$$-0.002 < h_4^\gamma < 0.034$$

$$-0.05 < h_4^Z < 0.12$$

$$h_i^V = \frac{h_{i0}^V}{(1 + \hat{s}/\Lambda^2)^n}$$

- LEP does not scale couplings with the form-factor, which makes direct comparison more complex
 - Additional $e^{i\pi/2}$ factor from Baur MC

Summary of these results

- We observed $Z\gamma \rightarrow \nu\nu\gamma$ for the first time at the Tevatron and measured the cross section to be in excellent agreement with the standard model
- We set the tightest limits on anomalous $ZZ\gamma$ and $Z\gamma\gamma$ couplings at the Tevatron

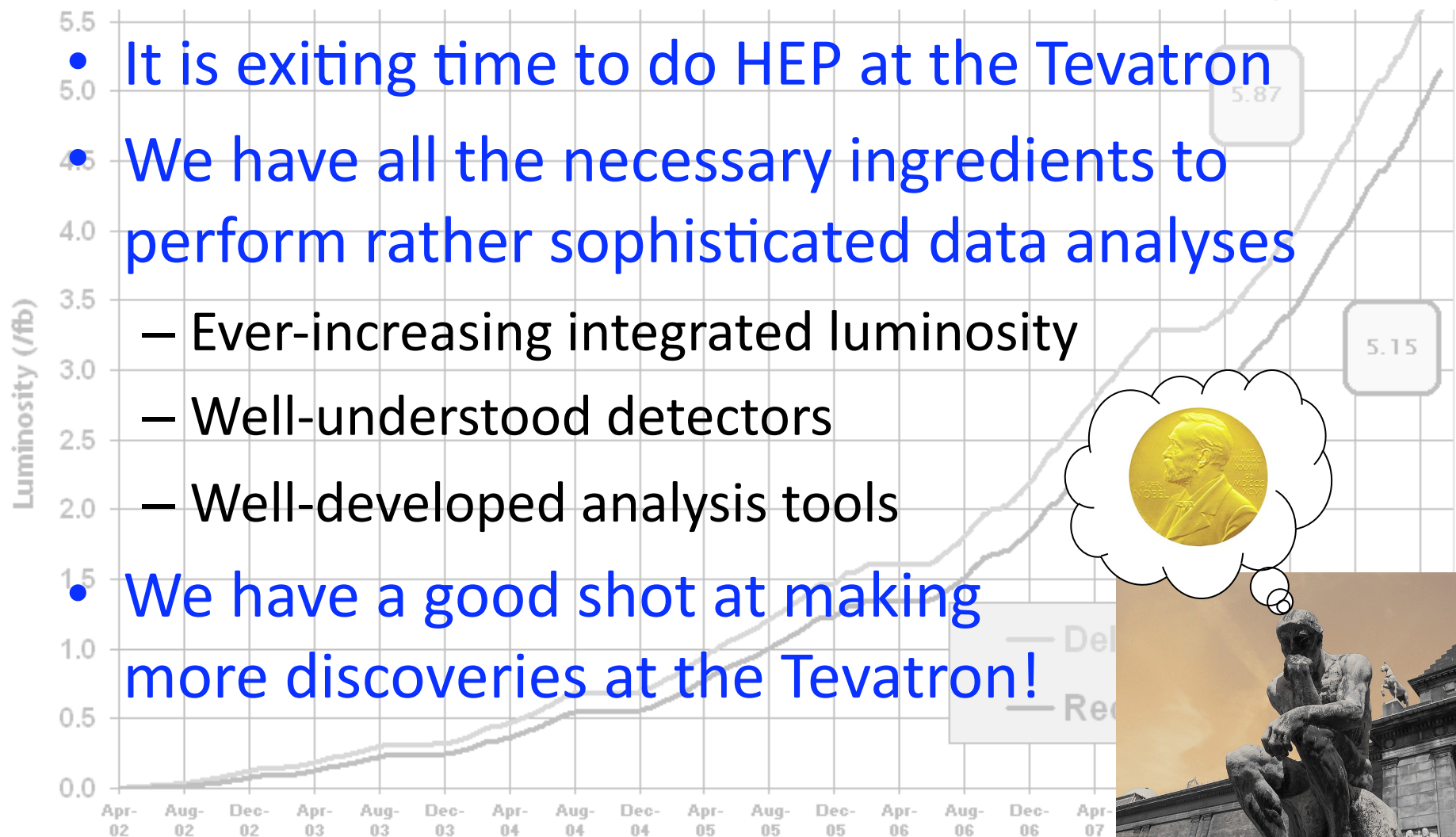
$$|h_{30}^\gamma| < 0.033 \quad |h_{40}^\gamma| < 0.0017$$

$$|h_{30}^Z| < 0.033 \quad |h_{40}^Z| < 0.0017$$

What comes next?

Prediction is very difficult, especially if it is about the future

Mark Twain, Niels Bohr, Yogi Berra



- It is exiting time to do HEP at the Tevatron
- We have all the necessary ingredients to perform rather sophisticated data analyses
 - Ever-increasing integrated luminosity
 - Well-understood detectors
 - Well-developed analysis tools
- We have a good shot at making more discoveries at the Tevatron!