



# NOvA

## Thomas Coan

- Intro
- Detector technology
- Physics reach
- Conclude

**SMU, April 2014**





# Clarification

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## Neutrinos, NOT Nutria !!





# Clarification



## Neutrinos, NOT Nutria !!



Fred: Can you check at Peggy Sue's ??



# Neutrinos Mix



Solar vs



Homestake, Gallex, SAGE, Super-K, SNO, Borexino

Reactor vs



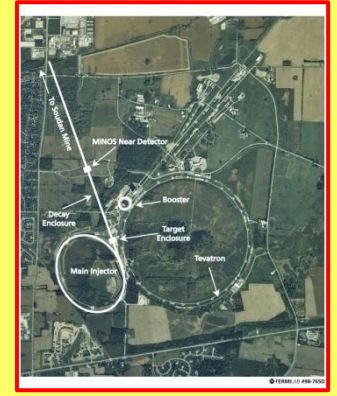
KamLAND, CHOOZ, ...

Atmospheric vs



Kamiokande, Super-K

Accelerator vs



K2K, MINOS, MiniBOONE

Neutrinos change flavor

⇒ Neutrinos have mass and mix

flavor eigenstates

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

mass eigenstates



# PMNS Leptonic Mixing Matrix



$$|\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle$$

$U_{\alpha i}$  can be written w/ 3 mixing angles and 1 phase  
and w/  $c_{ij} \equiv \cos \theta_{ij}$  and  $s_{ij} = \sin \theta_{ij}$

atmospheric

cross-mixing

solar

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$= \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

- Mixing angle  $\theta_{13}$  large ( $\sin^2 2\theta_{13} = 0.090 \pm 0.009$ ).
- CP violating phase  $\delta$  unknown.



# Vacuum Oscillations



Neutrinos propagate as mass eigenstates  
Interact as different (i.e., weak) eigenstates

➤ Oscillations result

In absence of matter (key caveat) & to lowest order:

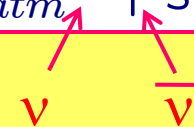
$$P_{vac}(\nu_\mu \rightarrow \nu_e) = \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \Delta_{atm}$$

$$\Delta_{atm} \simeq 1.27 \left( \frac{\Delta m_{32}^2 (\text{eV}^2) L (\text{km})}{E (\text{GeV})} \right)$$

Additional sub-dominant terms, sensitive to CPV phase  $\delta$ :

$$\Delta P_\delta(\nu_\mu \rightarrow \nu_e) = J \sin \Delta_{sol} \sin \Delta_{atm} (\cos \delta \cos \Delta_{atm} - \sin \delta \sin \Delta_{atm})$$

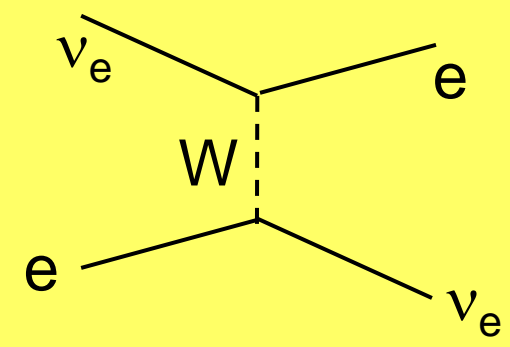
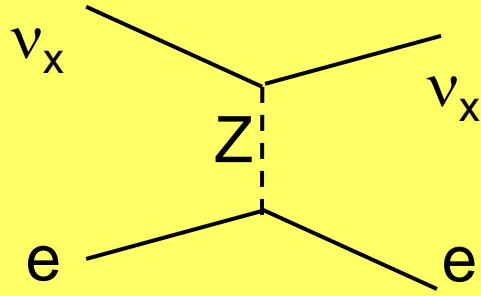
$$J = \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13}$$



... but NOvA's neutrinos travel through Earth ...



# “Matter Effects”



$\nu_e - e$  scattering Hamiltonian modified  
Effective mass eigenstates & mixing angles altered

$$P_{mat}(\nu_\mu \rightarrow \nu_e) \simeq (1 \pm 2 \frac{E}{E_R}) P_{vac}(\nu_\mu \rightarrow \nu_e)$$

$E_R \simeq 12 \text{ GeV}$  (earth's mantle)

$E(\nu) \simeq 2 \text{ GeV} \Rightarrow 30\%$  enhancement/suppression

NOvA's long baseline key



# Key NO $\nu$ A Physics Goals



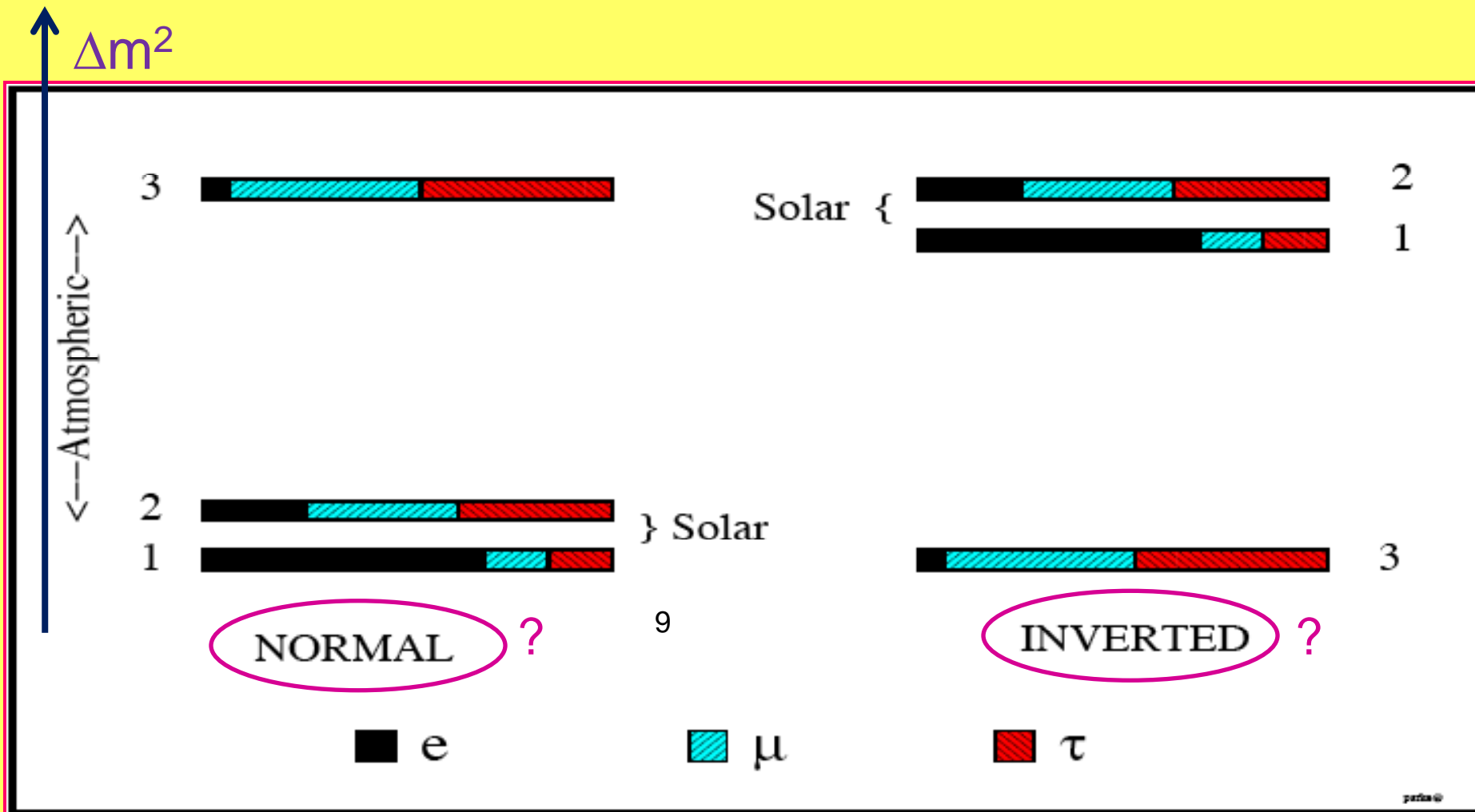
L/E  $\sim$  400 km/GeV regime:

- Measure  $\nu_{\mu} \rightarrow \nu_e$  &  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$  transitions
- Measure  $\nu_{\mu} \rightarrow \nu_{\mu}$  &  $\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{\mu}$  survival probability
  
- Search for CP violation in neutrino sector  
i.e., measure/constrain  $\delta_{CP}$
- Resolve neutrino mass hierarchy.
- Measure  $\sin^2(2\theta_{23})$  with high precision.
- Determine octant of  $\theta_{23}$
- Measure PMNS mixing angle  $\theta_{13}$
- Other (monopoles, supernovae, NSI...)





# Neutrino Ignorance...





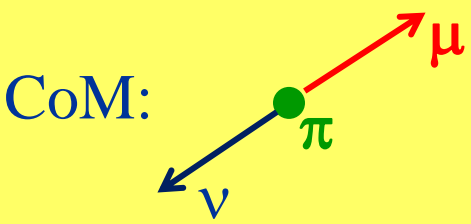
# 2-Detector Configuration



- Near detector measures un-osc  $\nu$  flux & bkg rates.
- Far detector measures osc  $\nu$  flux.



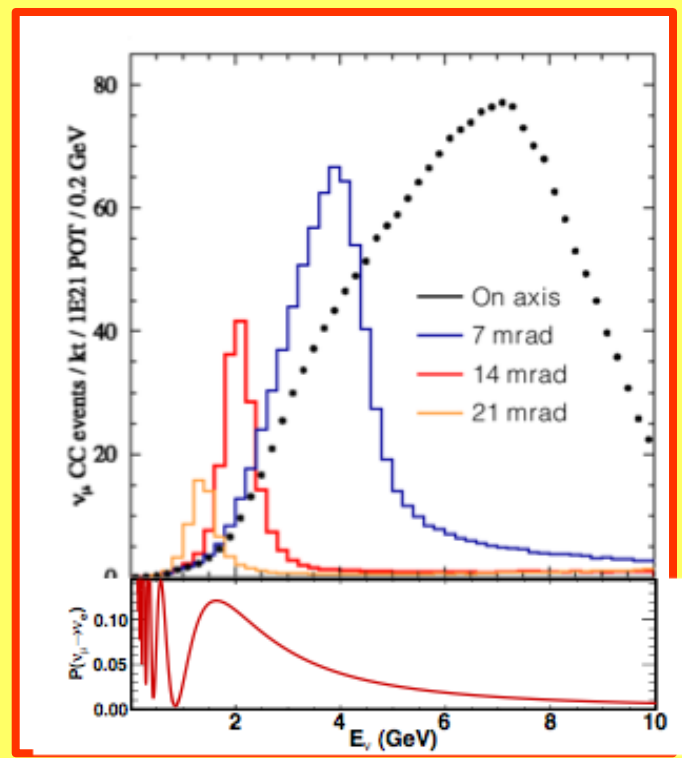
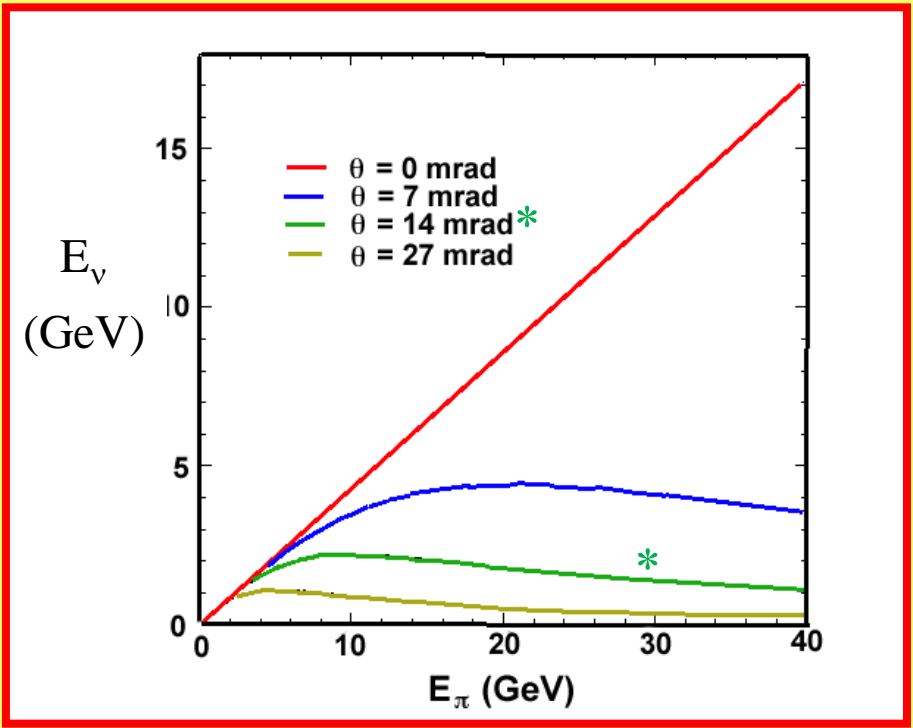
# Off-Axis Beam Technique



Lab Frame:

$$E_\nu \propto \frac{E_\pi}{1 + (E_\pi/m_\pi)^2 \theta^2}$$

$\therefore E(\nu) \sim$ independent of parent hadron E





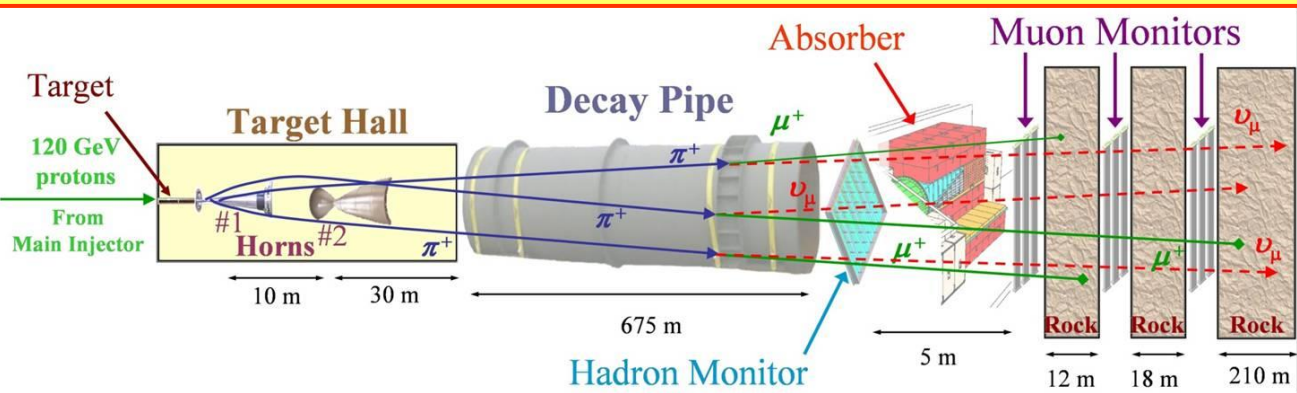


# NuMI Beam Upgrade



Booster → Recycler → MI → target  
p: 8 GeV → 120 GeV

- Recycler:  $\bar{p} \rightarrow p$  storage ring
  - slip-stack batches in Recycler
  - single turn extraction into MI
  - 53 MHz RF added
- 10  $\mu$ sec spill every 1.3 sec
- $4.9 \times 10^{13}$  p/pulse @120 GeV/p
- 700 kW beam power
- $6 \times 10^{20}$  POT/yr







# NuMI Beam Composition

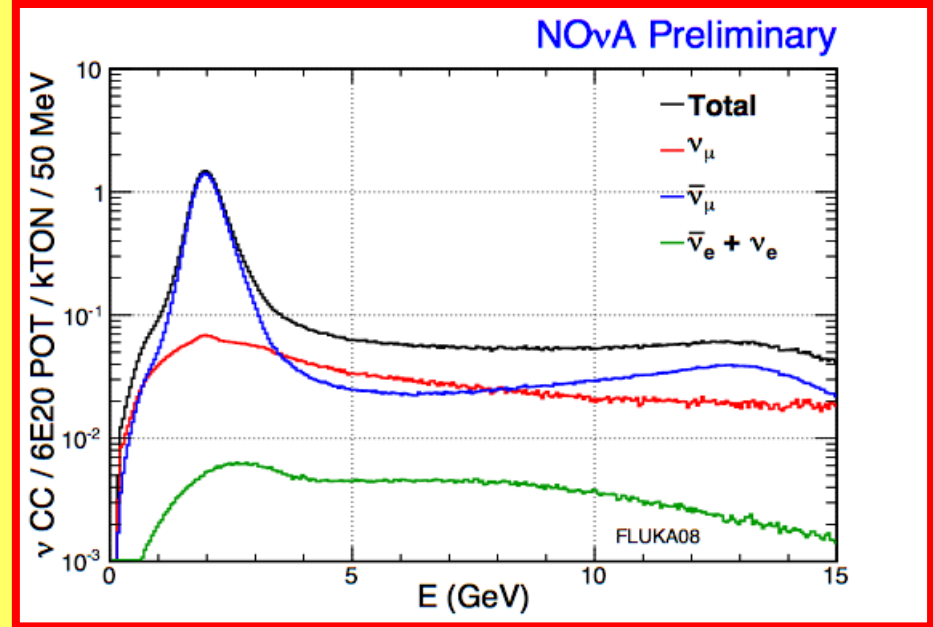
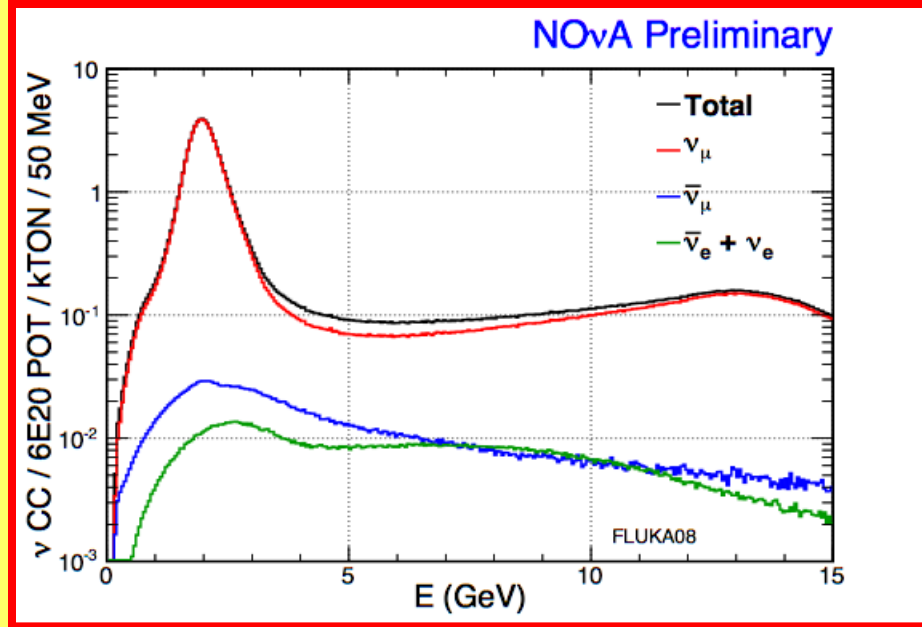


Horn focuses positive hadrons ( $\nu$  runs)

Horn focuses negative hadrons ( $\bar{\nu}$  runs)

NOvA Preliminary

NOvA Preliminary

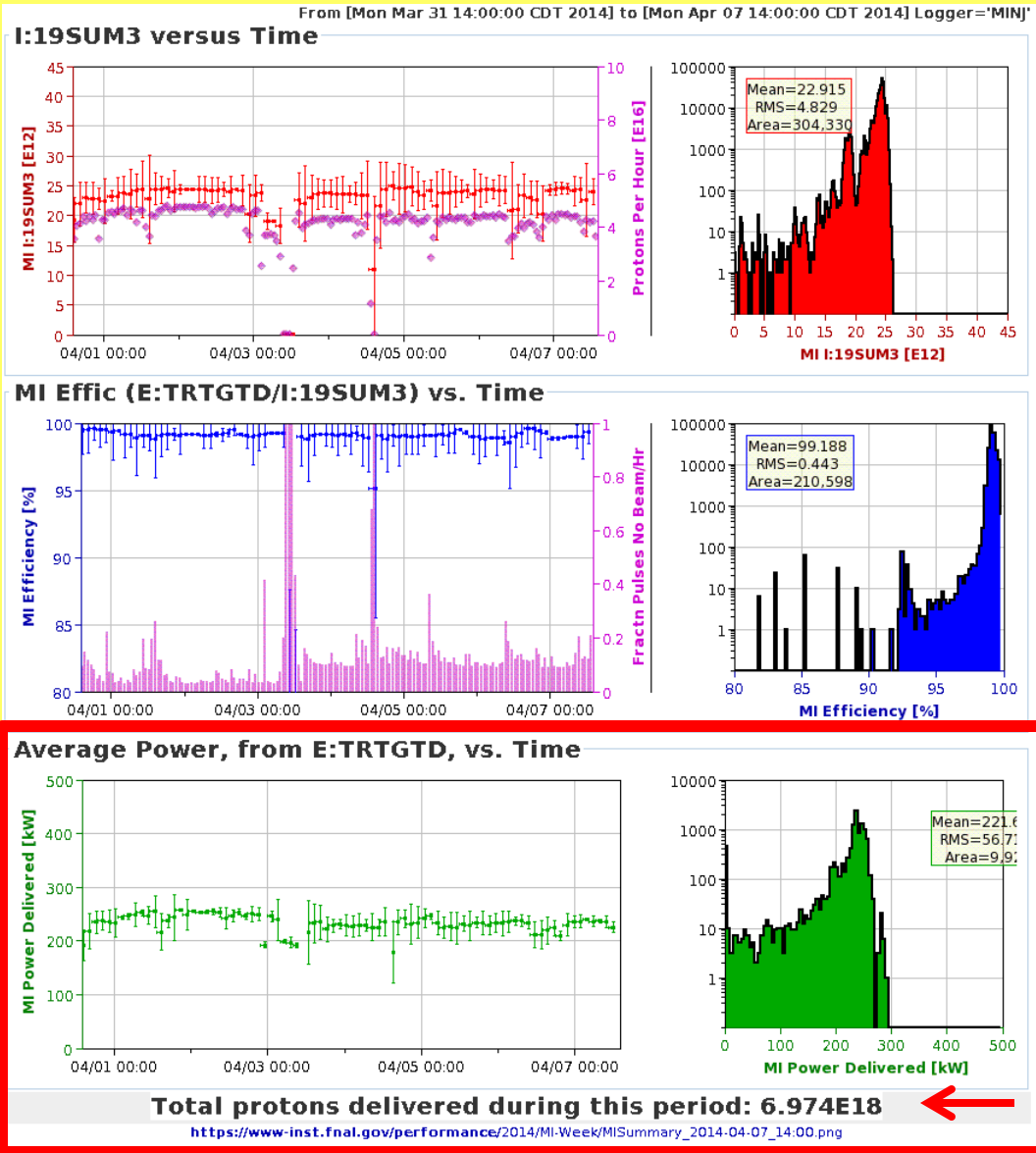




# NuMI Beam Performance



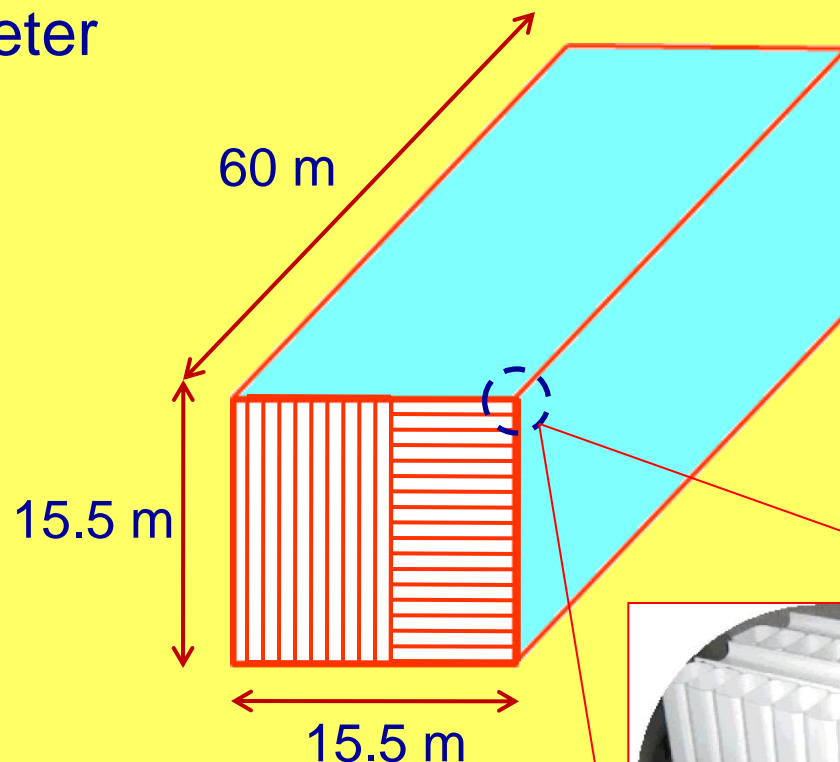
- Beam Power < 500 kW pending booster upgrade



# NO $\nu$ A Far Detector Overview



- Low-Z tracking calorimeter  
65% active
- Surface location
- 14 kT total mass
- 896 Detector planes  
Alternate x-y layers  
0.15  $X_0$ /layer  
 $R_M = 9.8$  cm (2.5 cells)
- Liquid scintillator cells  
32 PEs from far end
- 1-sided readout/plane  
via avalanche photodiodes (APDs)

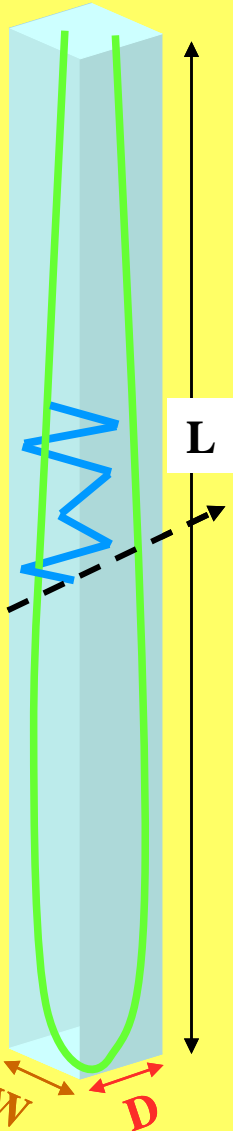




# NO<sub>v</sub>A Detector “Atom”



To 1 APD pixel



typical  
charged  
particle  
path

## Liquid Scintillator

Mineral oil solvent: 94.6 % (BW)

Primary scintillator: 5.2% (BW) pseudocumene

Waveshifters: PPO + bis-MSB

## Hollow PVC cells provide granularity

15% (BW) TiO<sub>2</sub>: high reflectivity walls

Each cell: 3.6 cm x 5.7 cm x 15.5 m long

## Looped Wavelength Shifting Fiber

Maximizes light collection: no mirrors

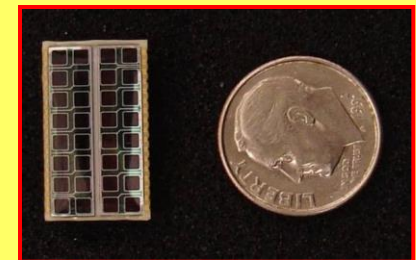
Diameter = 0.7mm, K-27 dye @ 300ppm

## Avalanche Photodiode

QE = 85%

Gain = 100

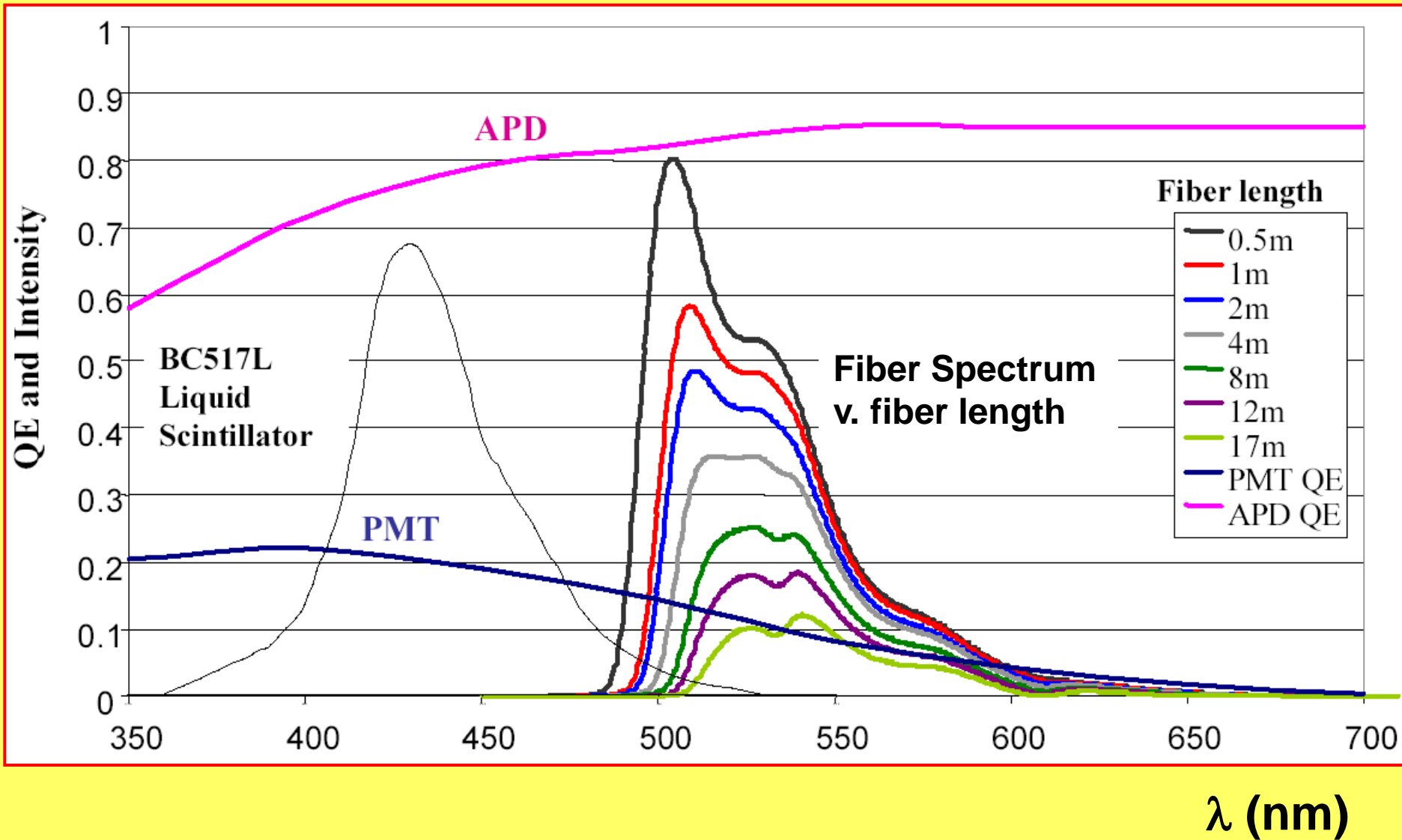
T<sub>run</sub> = -15 C







# Why APDs

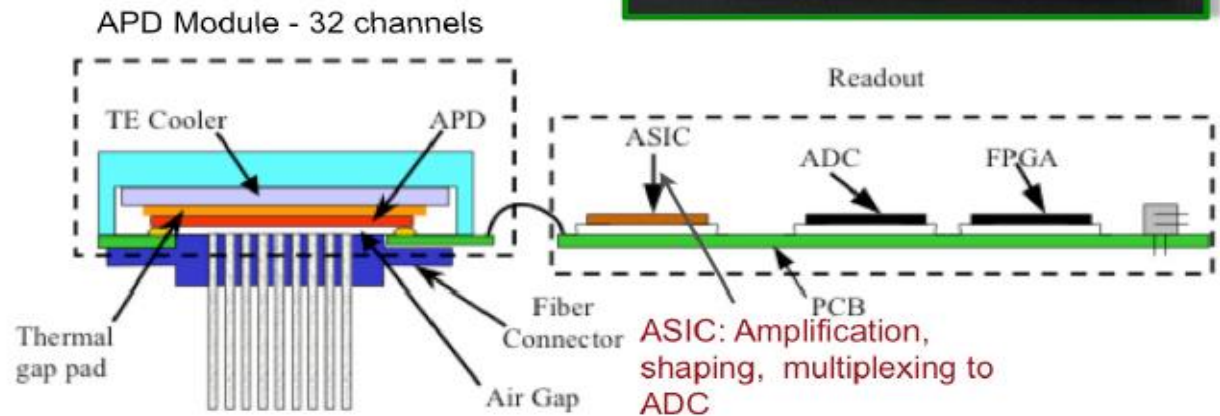




# Front-end Electronics



APD



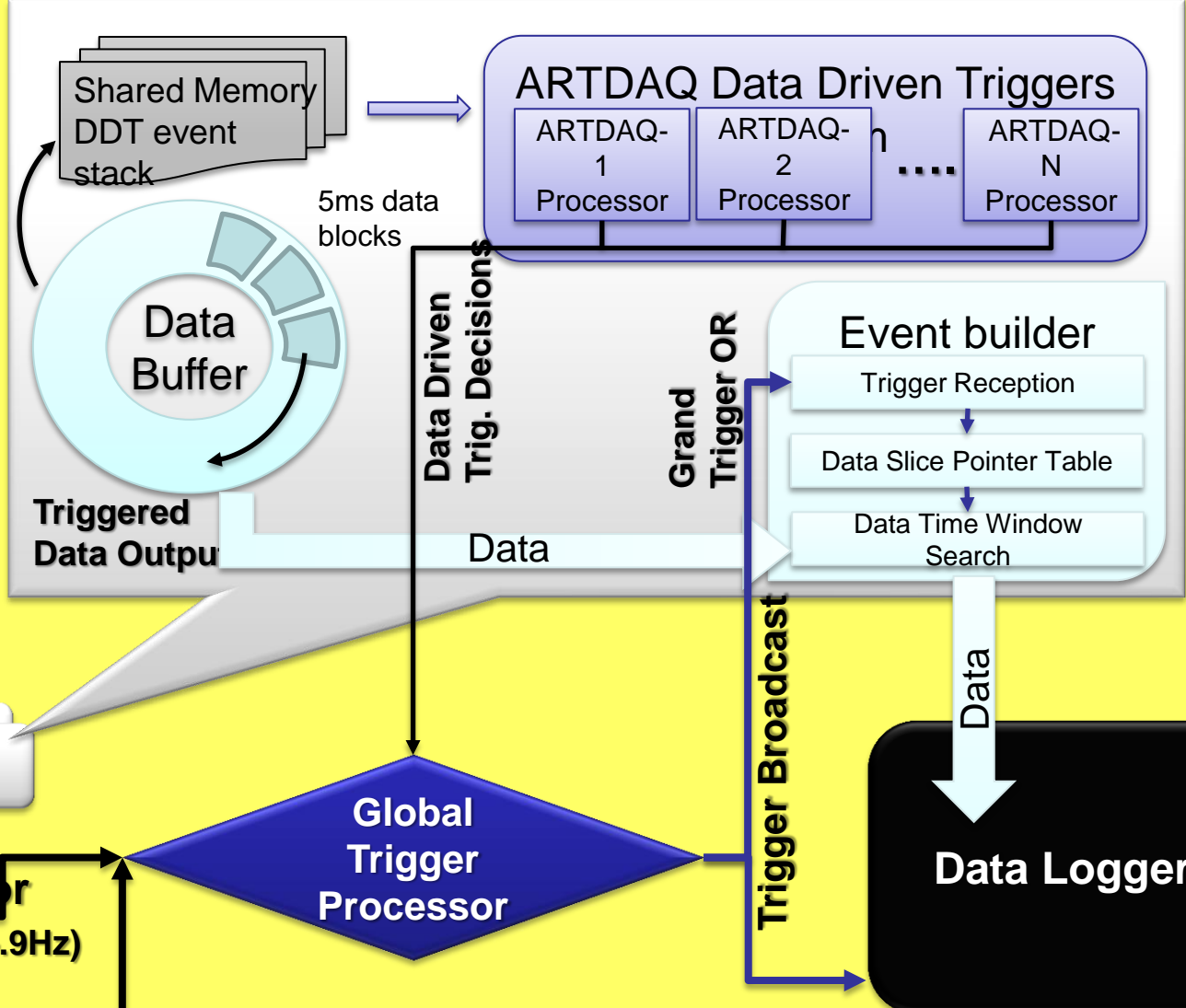
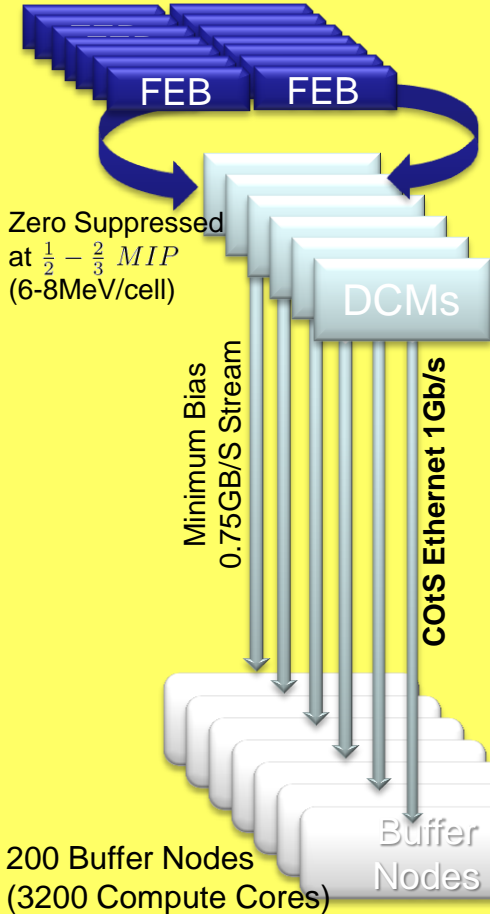


# DAQ Architceture



## DAQ: Triggerless, continuous, no dead-time

10 752 FEBs  
(344 064 det. channels)



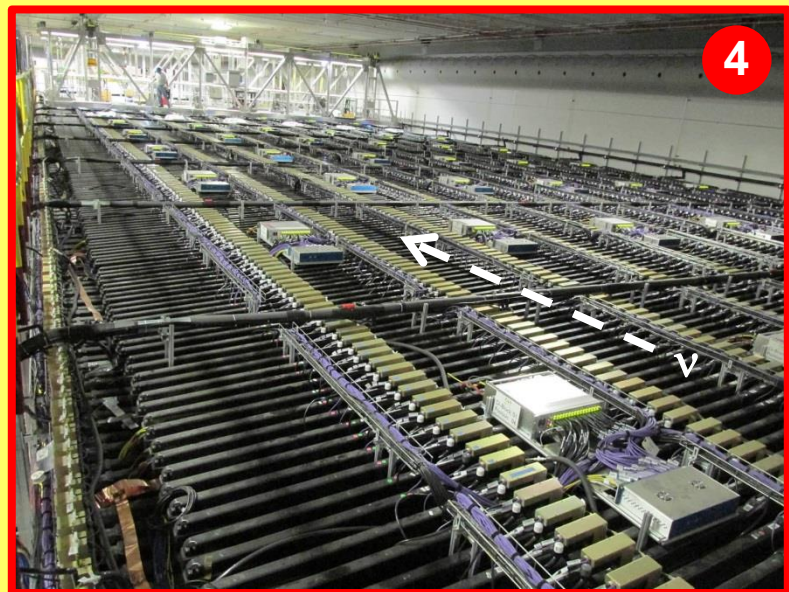
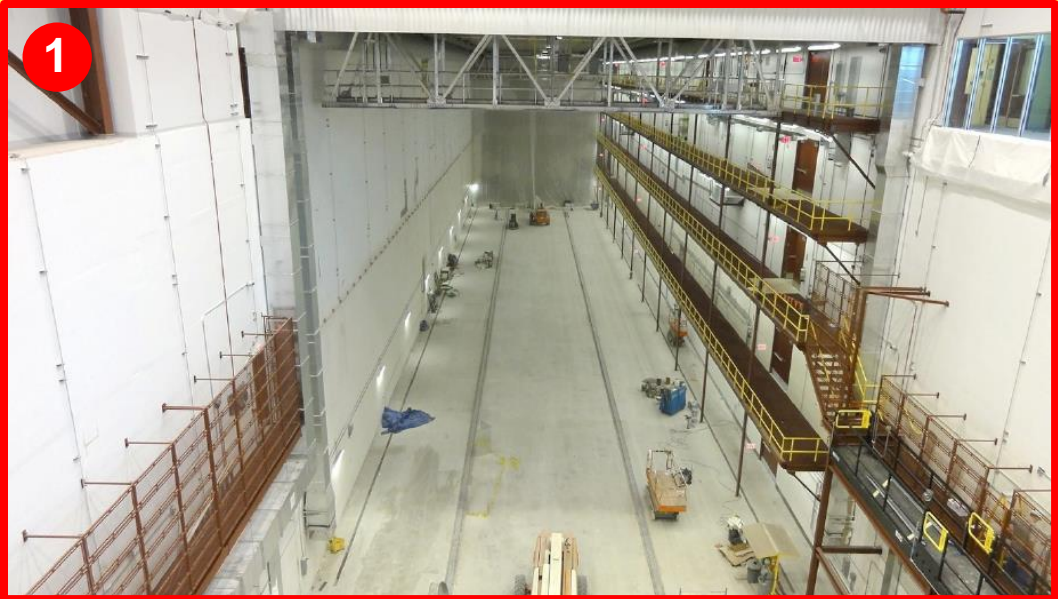
**Beam Spill Indicator**  
(Async from FNAL @ .5-.9Hz)

**Calib. Pulser ( 50-91Hz)**





# Far Detector Construction





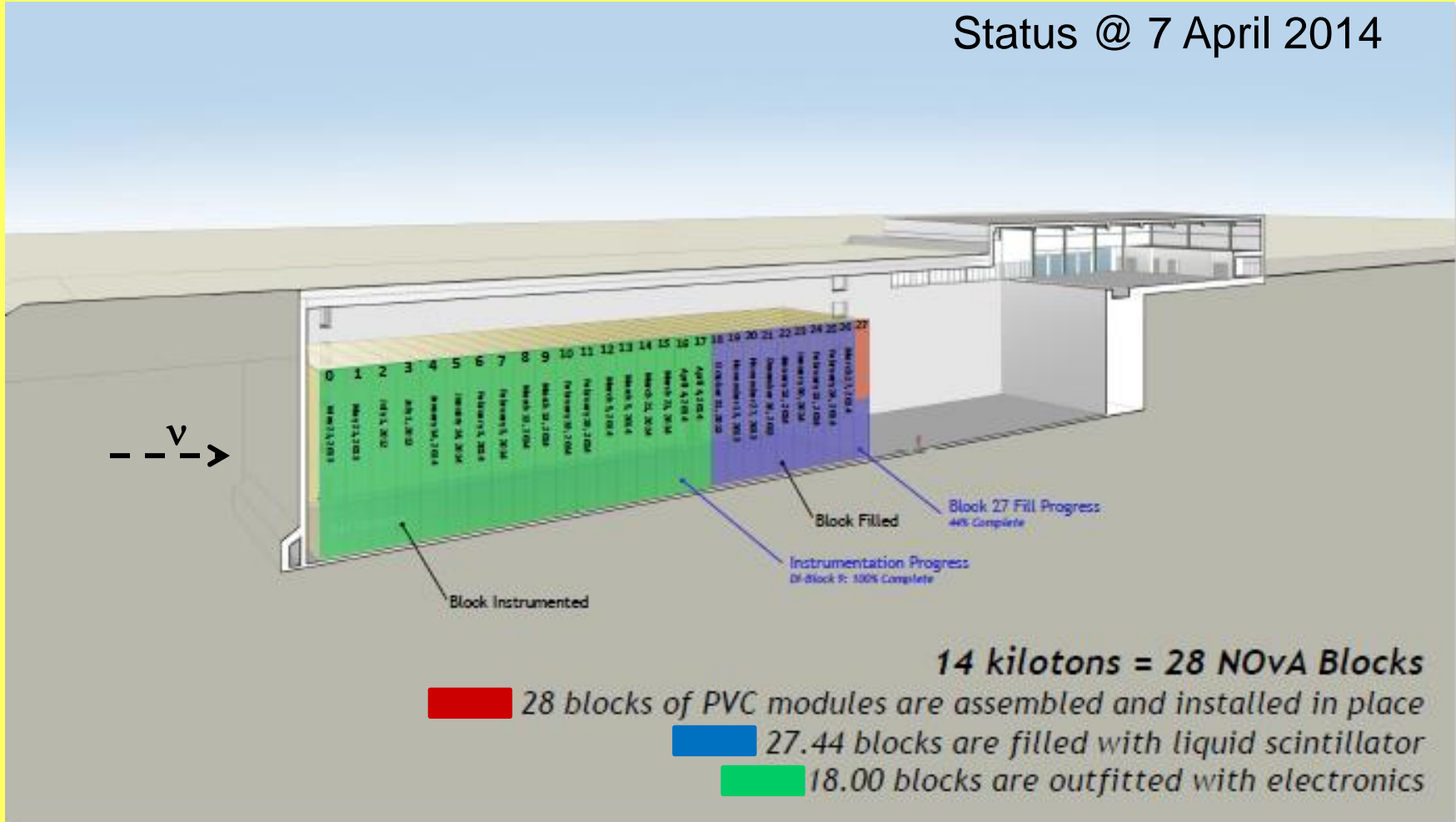


# Far Detector Status



Completion date - July 2014

Status @ 7 April 2014

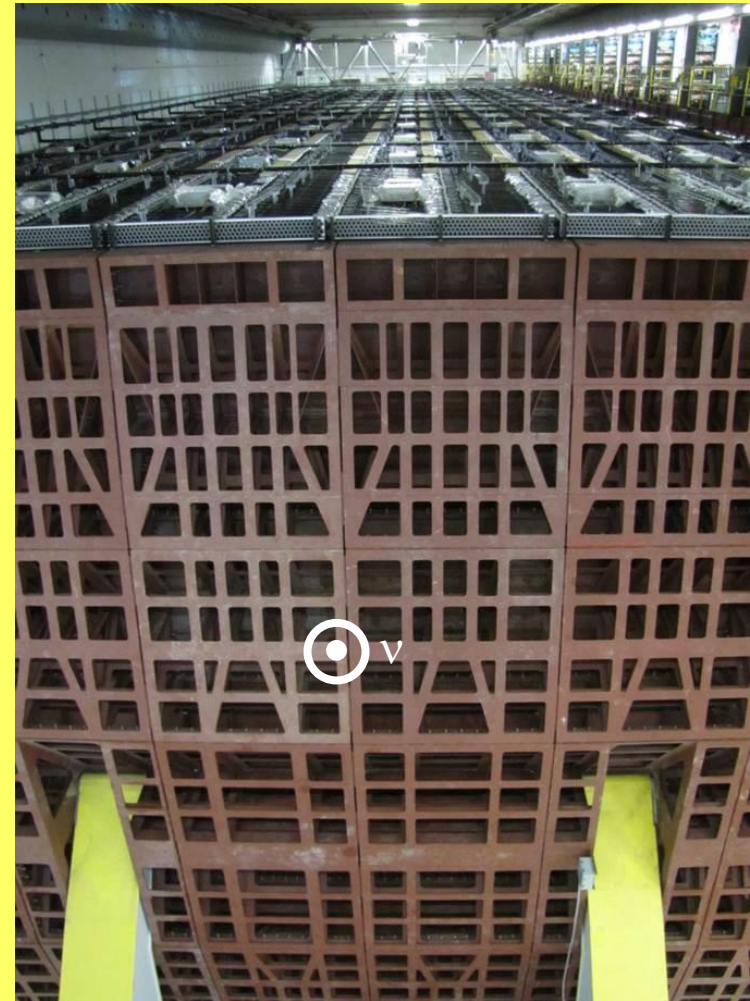




# Far Detector Status (2)



Last block installed 25 Feb 2014

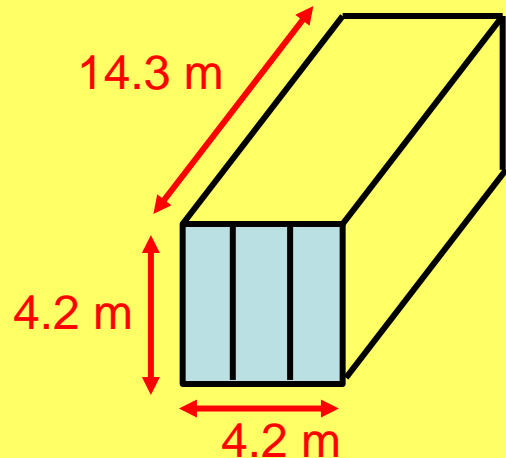




# Near Detector Construction



- 0.3 kT mass
- 20k channels
- 1 km from target & 100 m underground
- Cell structure similar to far detector
- Front end & DAQ identical to far detector
- Completion ~May 2014



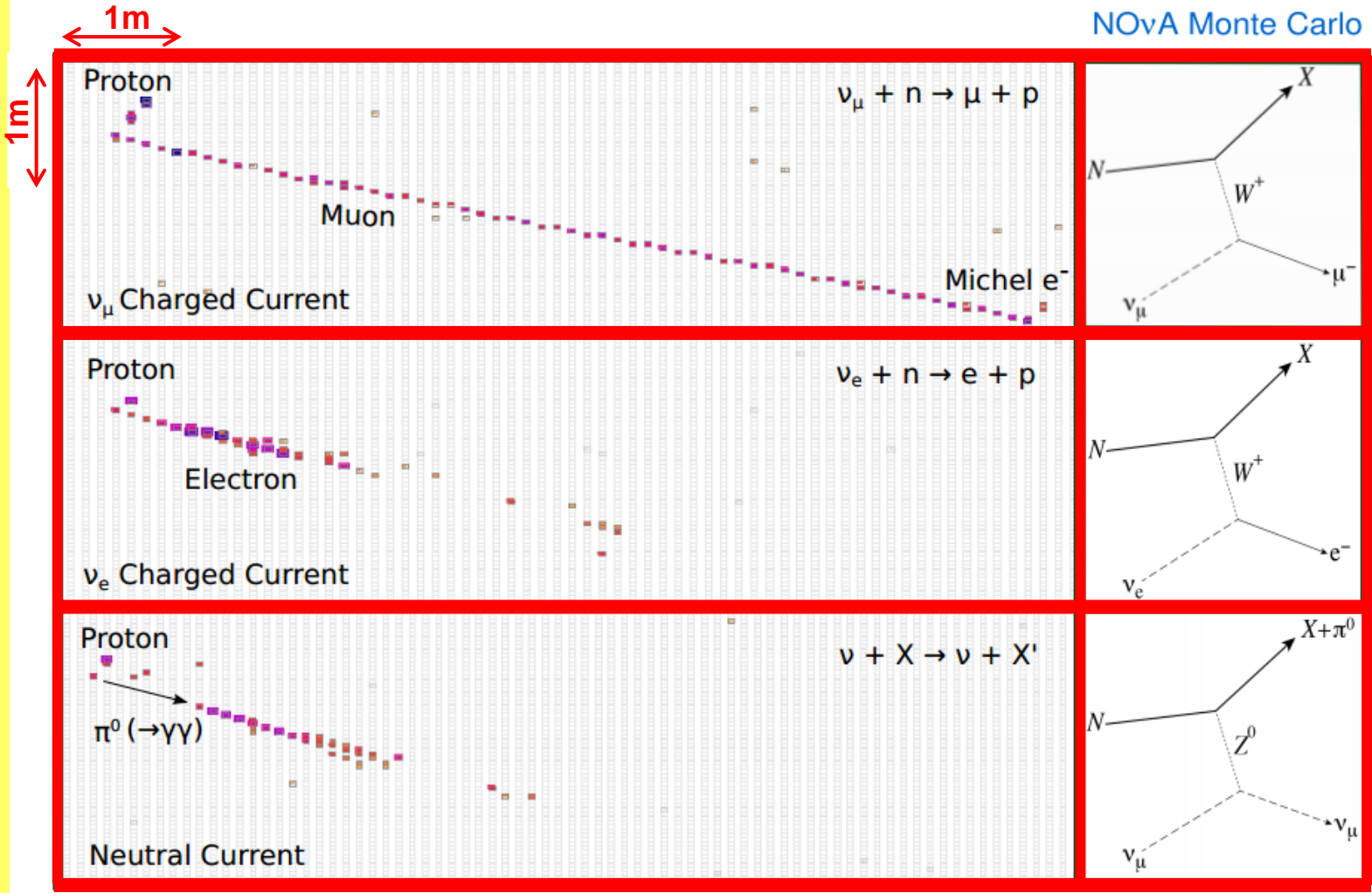




# Final State Topologies



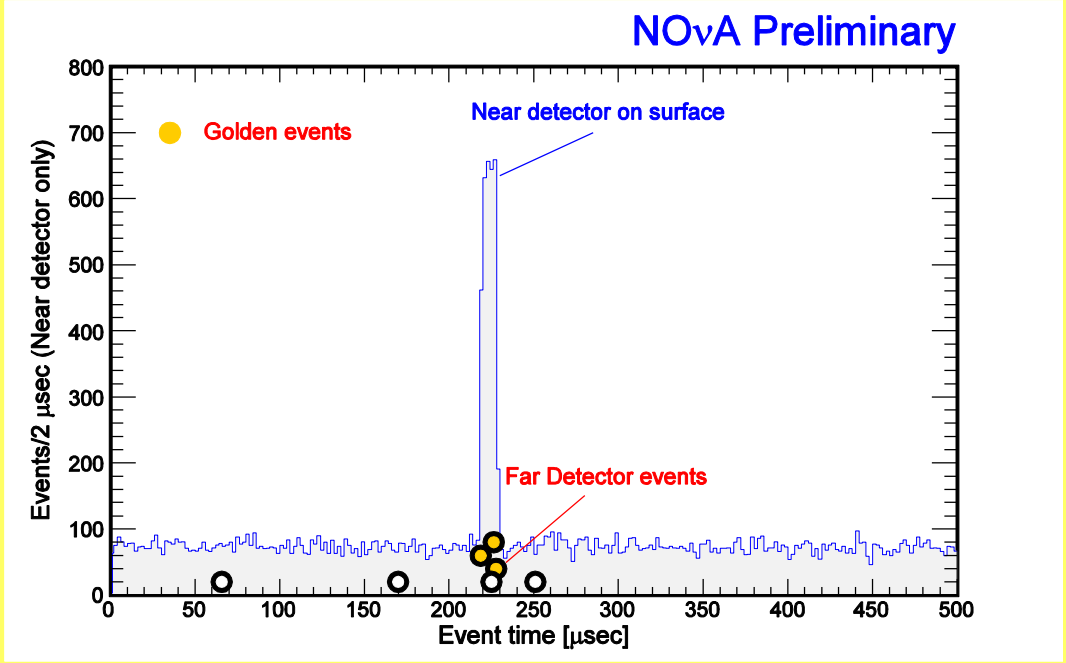
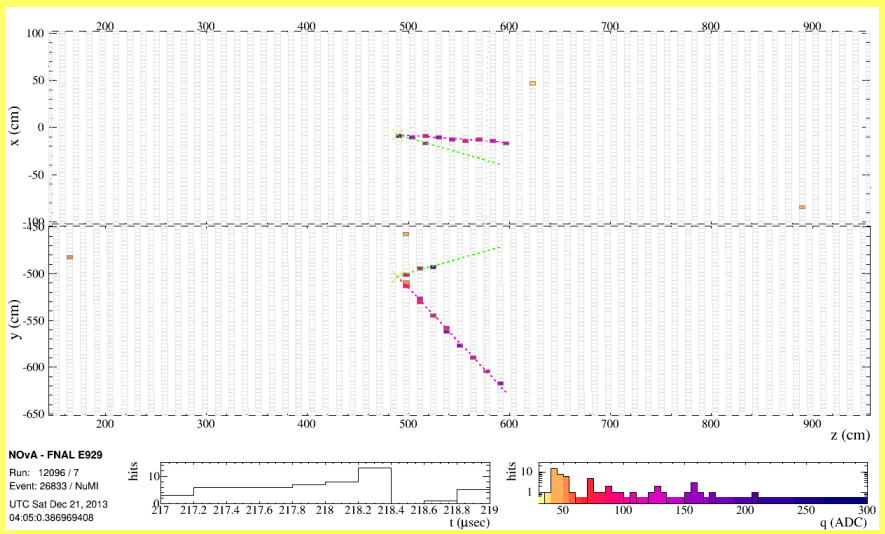
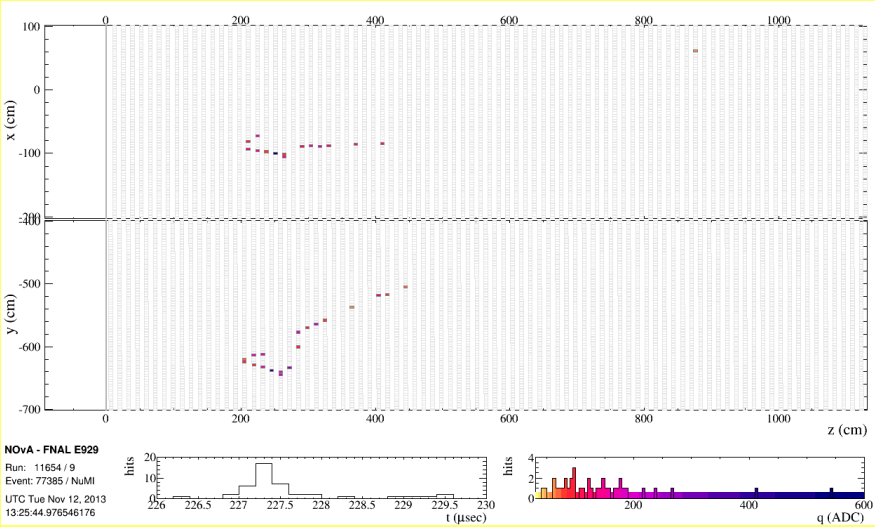
NOvA Monte Carlo







# Neutrinos !!

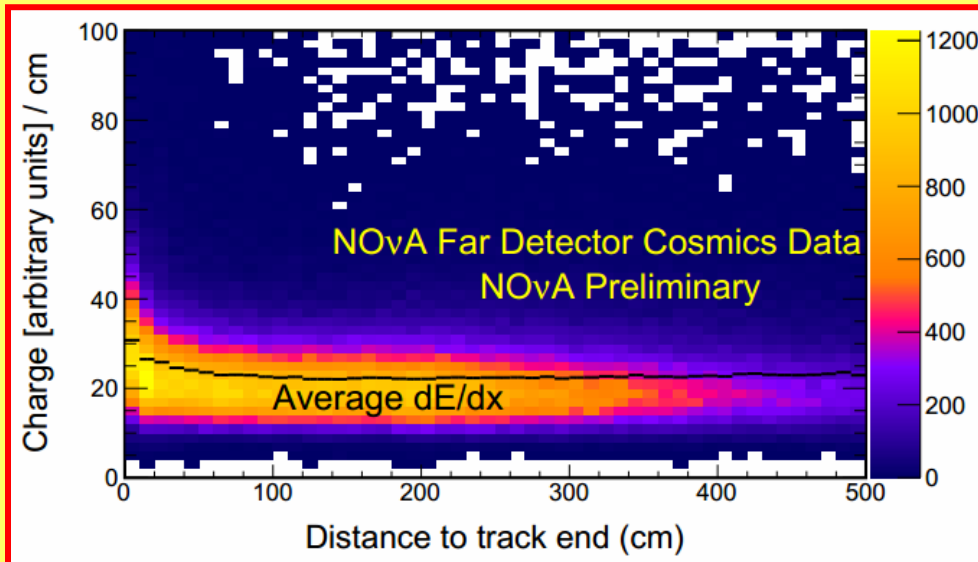
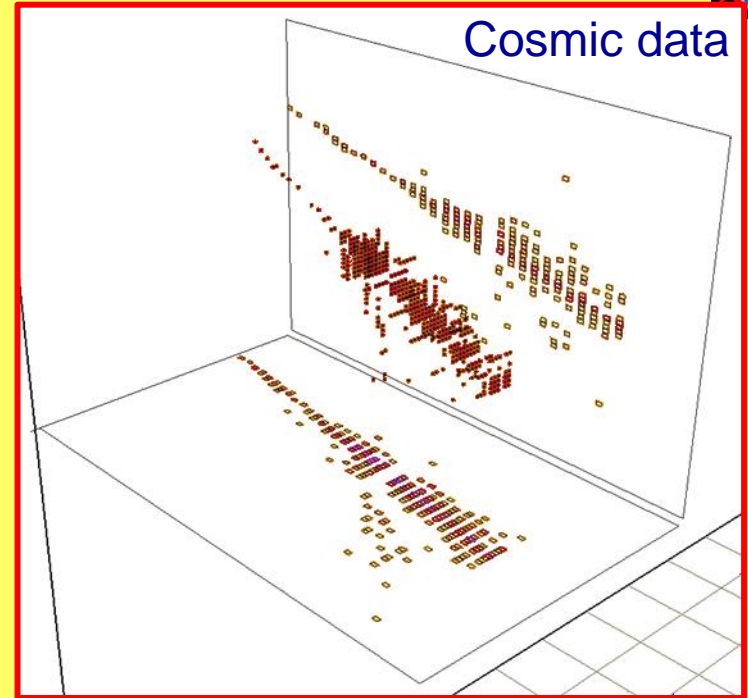




# Performance w/ Cosmics



- 9.0 kT/14 kT instrumented.
- Lots of cosmics observed.
- Reconstruction & calibration algorithms tested on cosmics

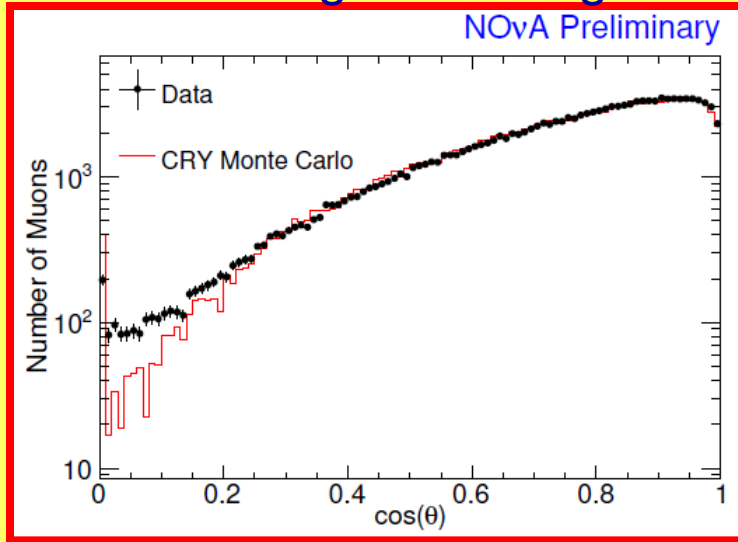




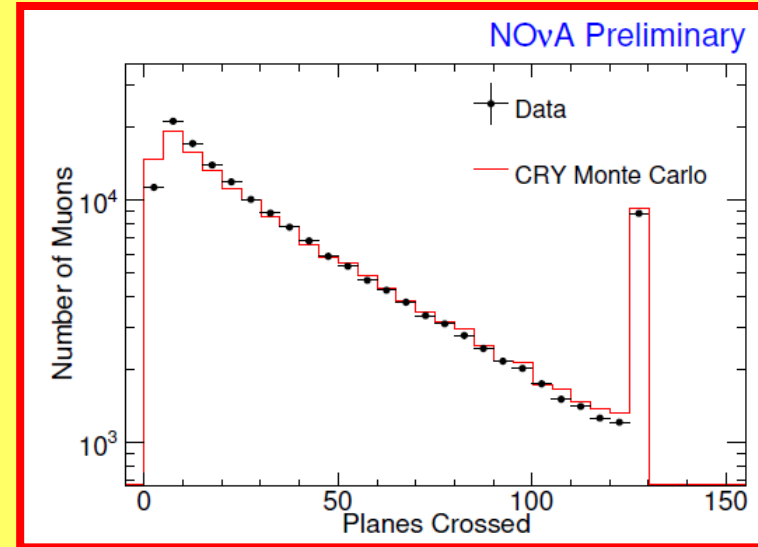
# Performance w/ Cosmics (2)



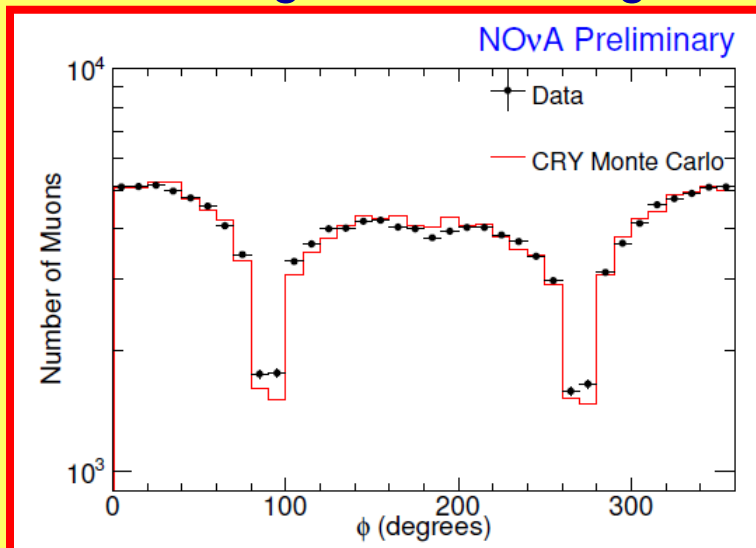
## Incoming zenith angle



## Planes crossed



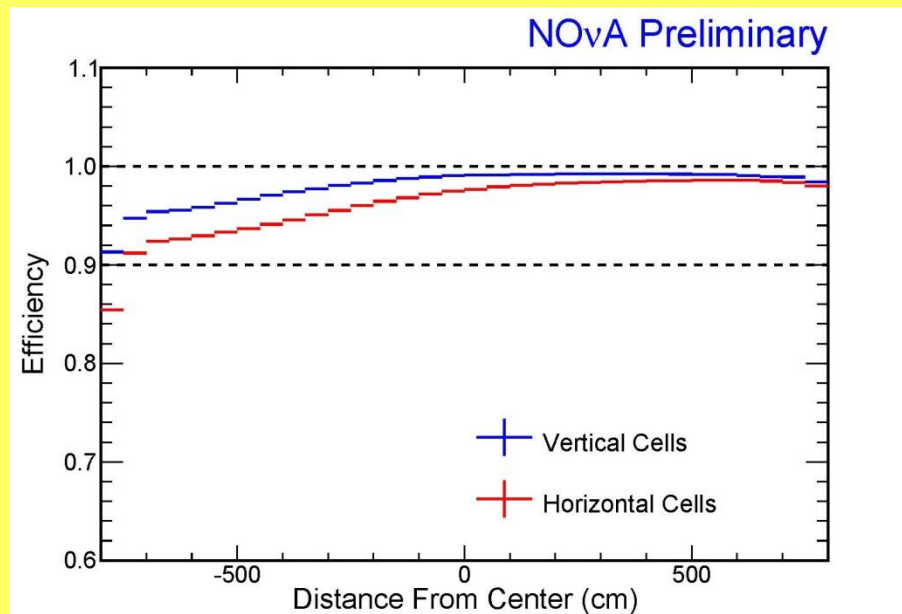
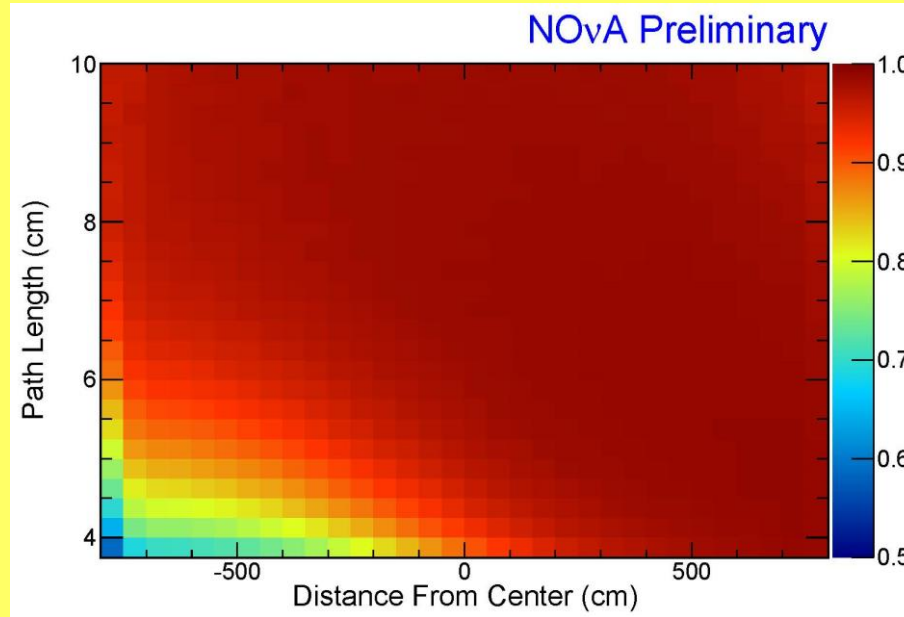
## Incoming azimuthal angle



- 1/7 detector only.
- 14k data triggers.
- Tracks must pass  $> 7$  planes.
- MC/data agreement good.



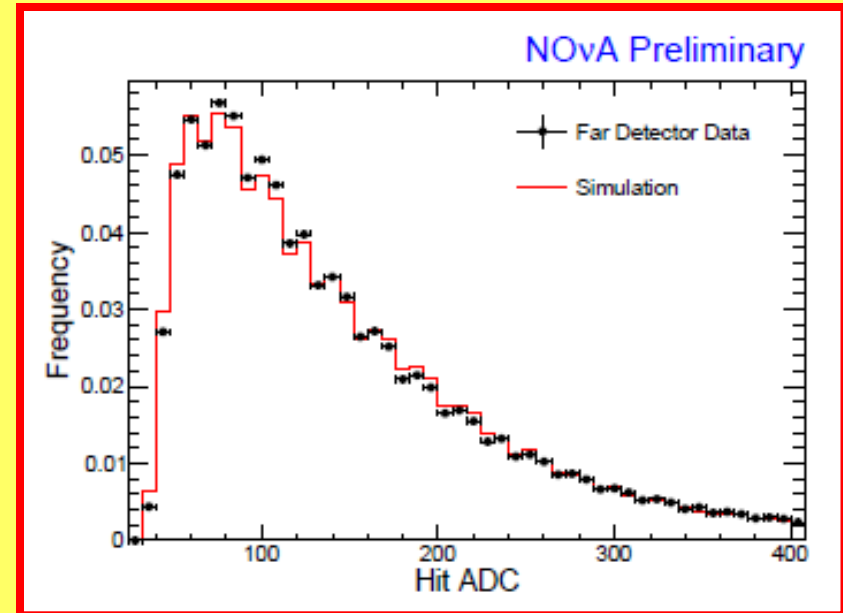
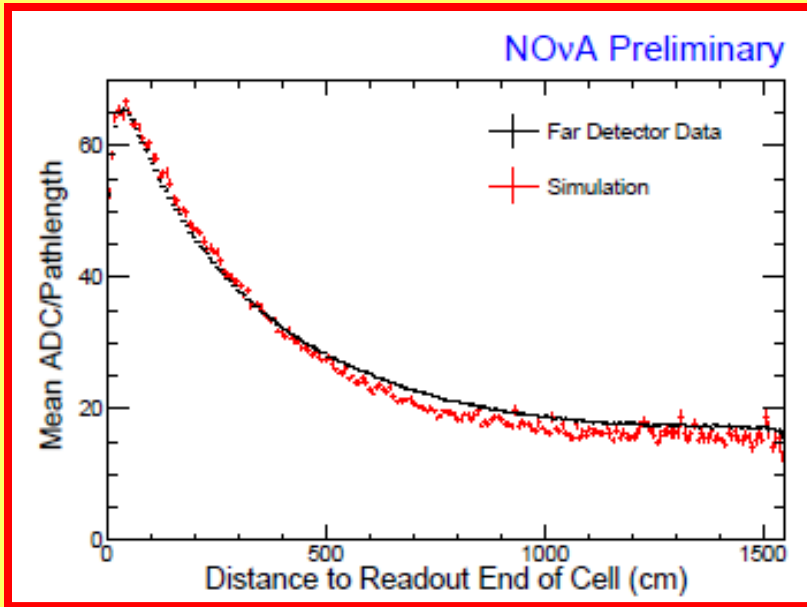
# Hit Efficiency w/ Cosmics







# Far Detector: MC v Data



- Decent MC/data agreement.
- Expected to improve.

# Physics Measurement Scheme

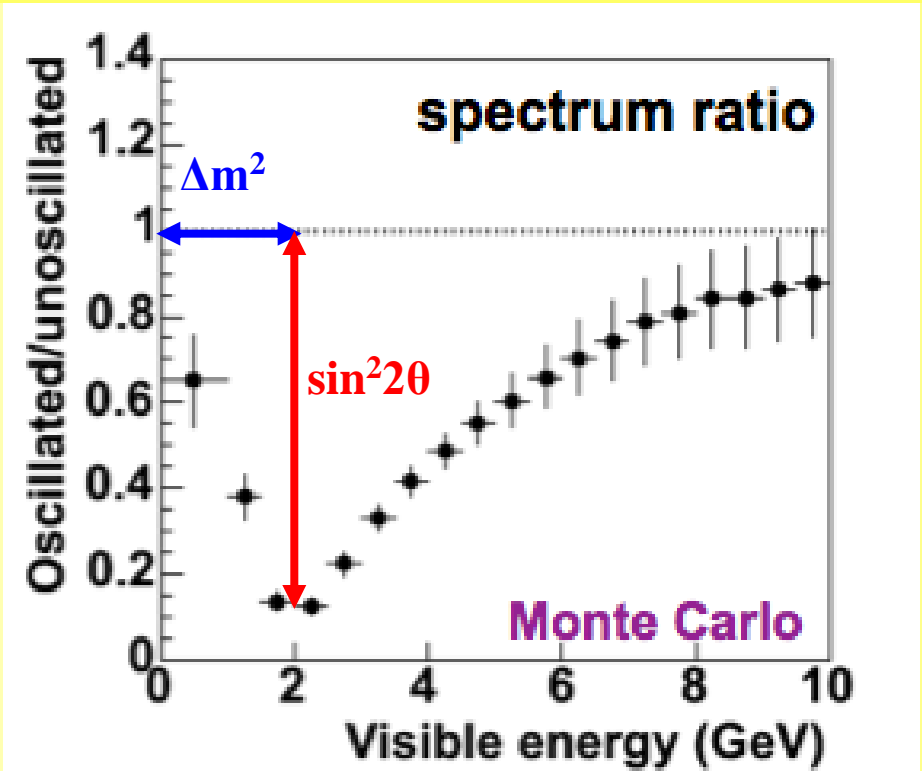
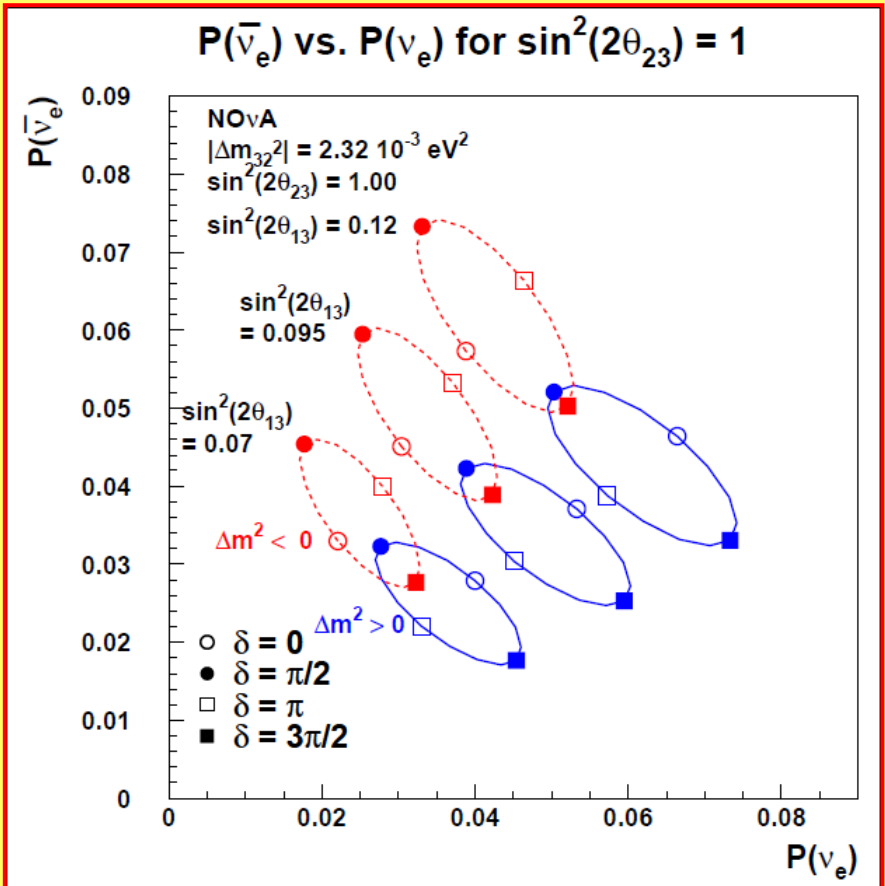


NOvA measures probability of 4 reactions:

$$\begin{array}{ll} \nu_{\mu} \rightarrow \nu_e & \bar{\nu}_{\mu} \rightarrow \bar{\nu}_e \\ \nu_{\mu} \rightarrow \nu_{\mu} & \bar{\nu}_{\mu} \rightarrow \bar{\nu}_{\mu} \end{array}$$

“appearance” measurements

“disappearance” measurements





# Anticipated Event Yields



- $6 \times 10^{20}$  POT/yr  $\nu$ 's &  $6 \times 10^{20}$  POT/yr anti- $\nu$ 's
- 3 yrs  $\nu$ 's + 3 yrs anti- $\nu$ 's
- $\sin^2 2\theta_{13} = 0.095$ ,  $\sin^2 2\theta_{23} = 0.95$  or  $1.0$
- $\Delta m^2_{32} = 2.35 \times 10^{-3} \text{ eV}^2$

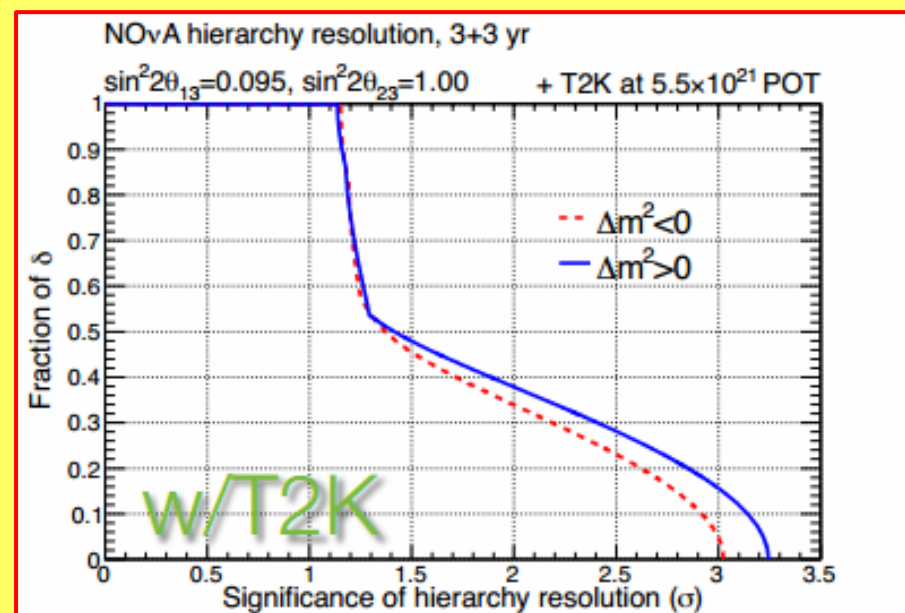
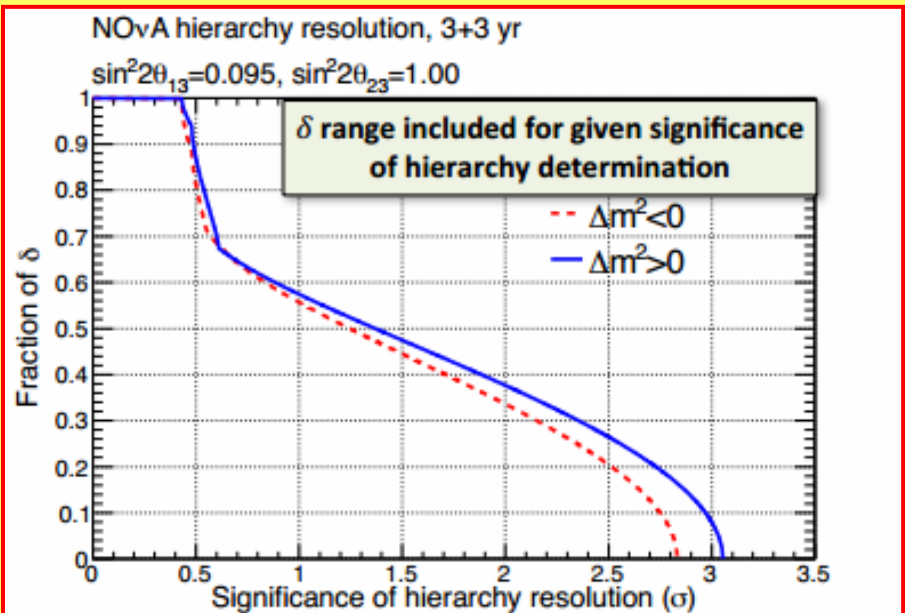
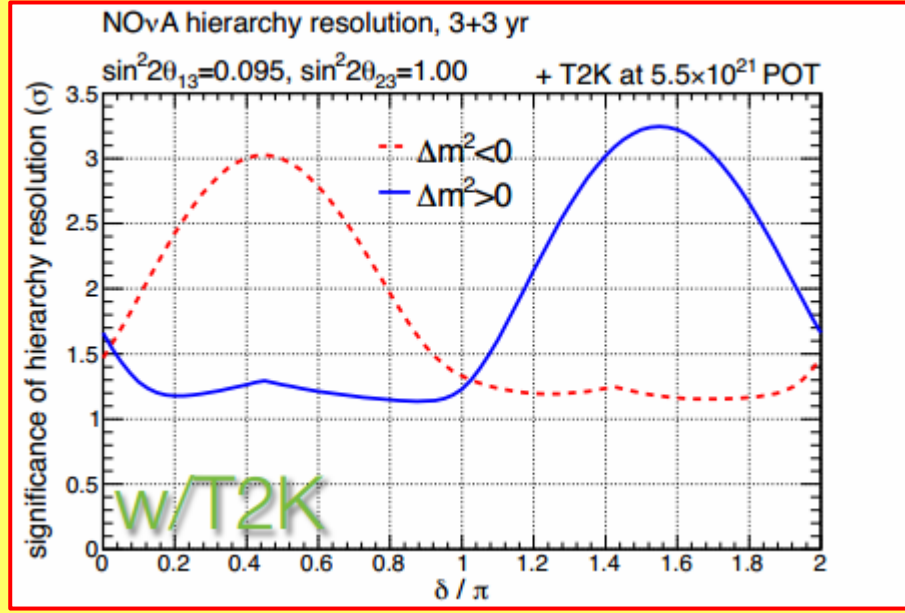
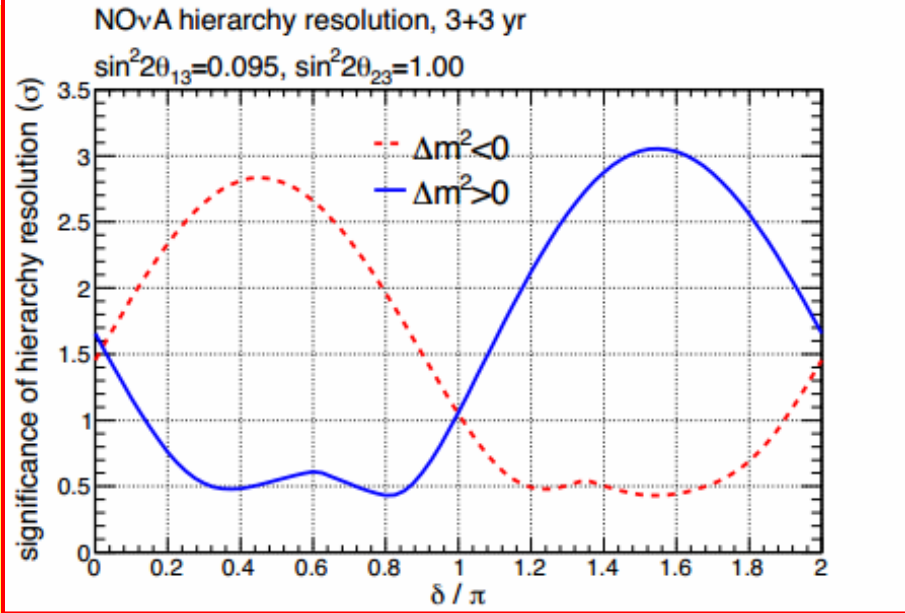
$\nu_e$ selected	$\nu$	anti- $\nu$
NC	19	10
$\nu_\mu$ CC	5	<1
Beam $\nu_e$	8	5
Total bkg	32	15
Signal	68	32

$\nu_\mu$ selected	$\nu$	anti- $\nu$
QE signal	82	49
NC bkg	< 1	< 1
Non-QE signal	168	78
NC bkg	14	6
Uncont. signal	233	134
NC bkg	6	3

Need to be patient !!



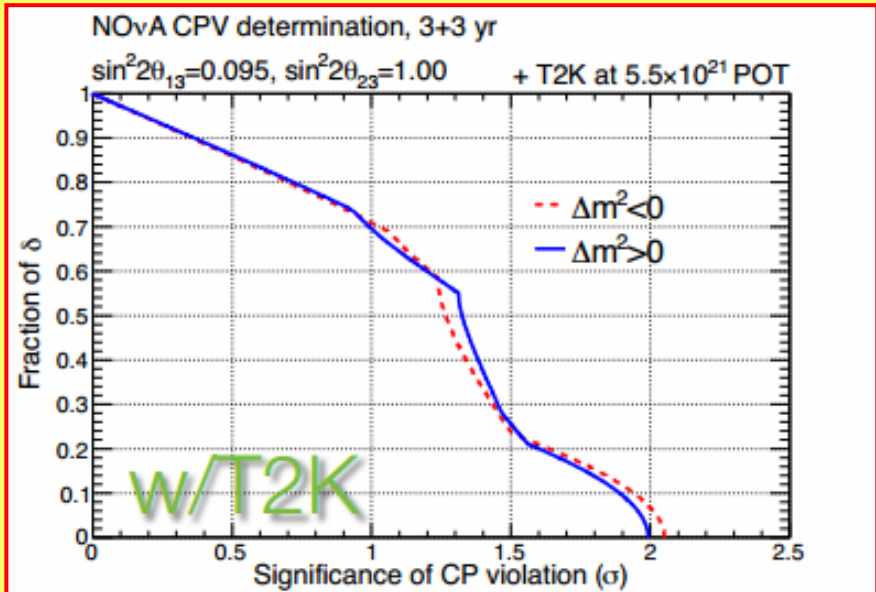
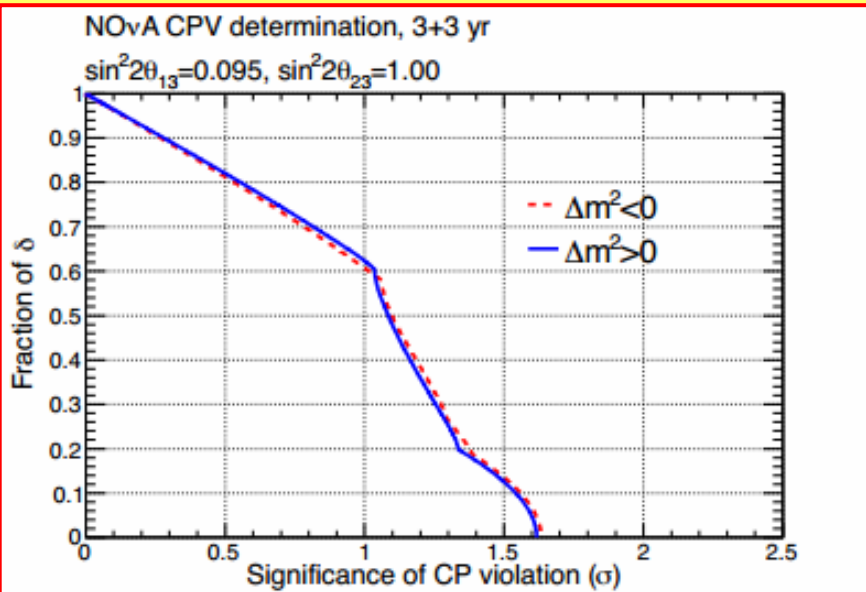
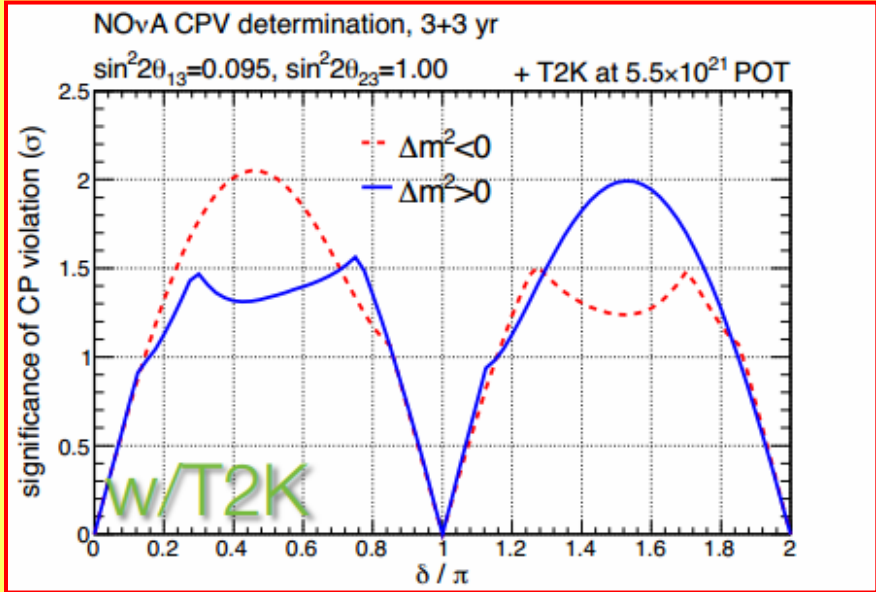
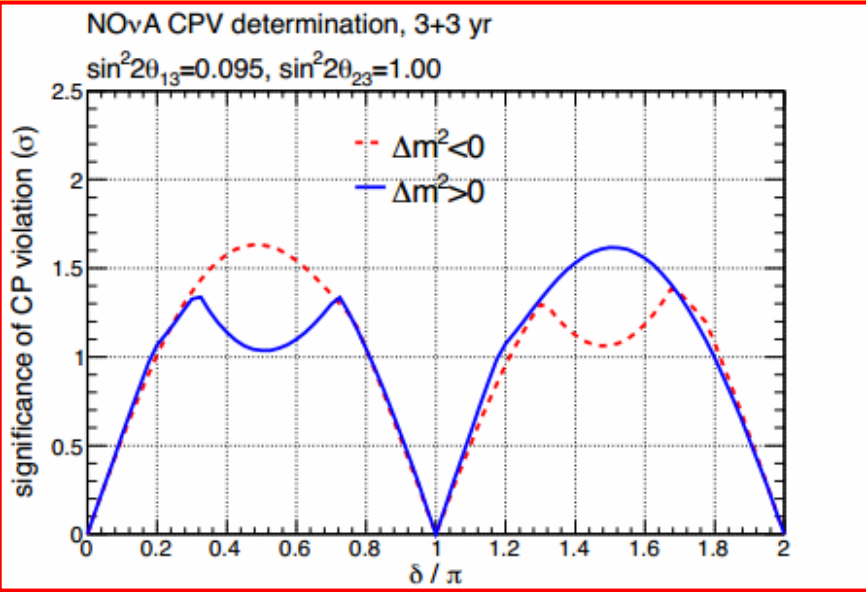
# Mass Hierarchy Sensitivity







# $\delta$ \_CP Sensitivity

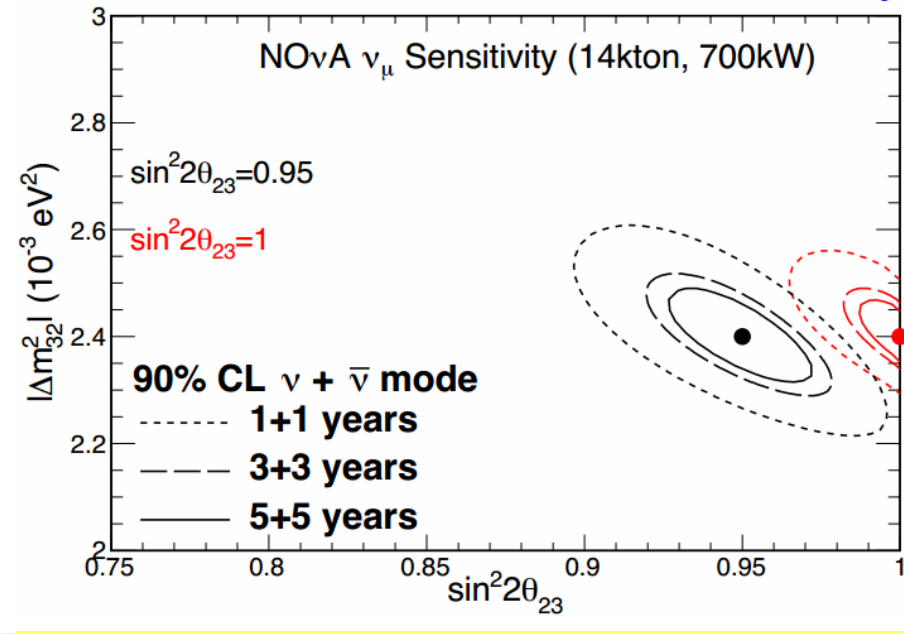
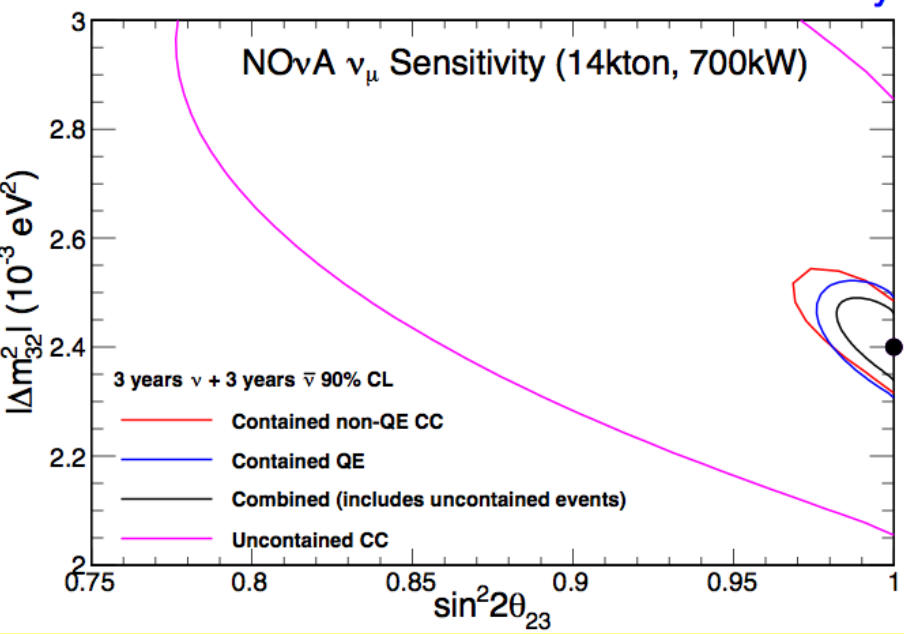


# $\nu_\mu$ Disappearance Measurements



NOvA Preliminary

NOvA Preliminary





# NOvA Summary

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- Both detectors rapidly being completed.
- Far detector performing well.
- Precise measurement of  $\sin^2(2\theta_{23})$  &  $\Delta m^2_{23}$ .
- Constrain  $\theta_{23}$  octant.
- Constrain CP violating angle  $\delta_{CP}$ .
- Information on neutrino mass hierarchy.
- Measure  $\theta_{13}$ .
- NOvA detector finished summer 2014
- Expect first publication late 2014.



# Backup





# $\nu$ Transition Probability



$$\begin{aligned}
P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) &\approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2} \\
&+ 2\alpha \sin\theta_{13} \sin\delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \sin\Delta \\
&+ 2\alpha \sin\theta_{13} \cos\delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \cos\Delta
\end{aligned}$$

$\alpha = \Delta m_{21}^2 / \Delta m_{31}^2$        $\Delta = \Delta m_{31}^2 L / (4E)$        $A = \frac{(-)}{+} G_{f\nu_e} L / (\sqrt{2}\Delta)$