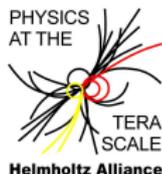


Introduction to Monte Carlos

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CTEQ Summer School 2013
Pittsburgh, 7–17 July 2013



- ▶ Part I — Basics
 - ▶ Introduction
 - ▶ Monte Carlo techniques
- ▶ Part II — Perturbative physics
 - ▶ Hard scattering
 - ▶ Parton showers
- ▶ Part III — Non-perturbative physics
 - ▶ Hadronization
 - ▶ Hadronic decays
 - ▶ Comparison to data

Thanks to my colleagues

Frank Krauss, Leif Lönnblad, Steve Mrenna, Peter Richardson,
Mike Seymour, Torbjörn Sjöstrand.

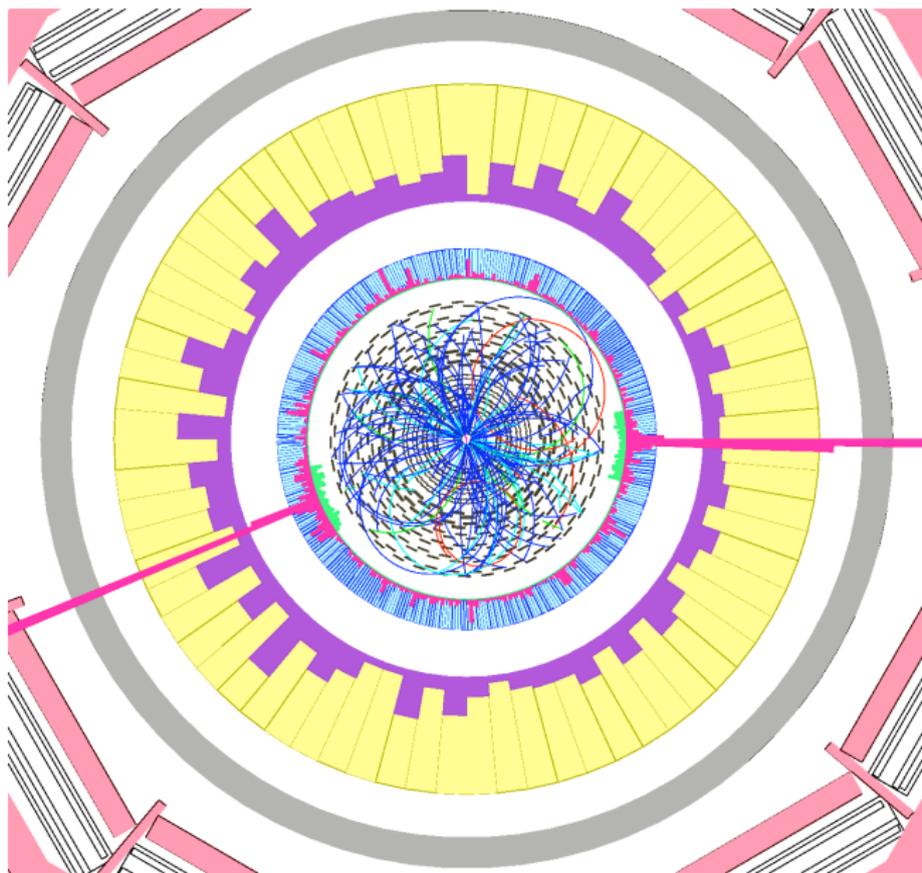
Introduction

Why Monte Carlos?

We want to understand

$$\mathcal{L}_{\text{int}} \longleftrightarrow \text{Final states} .$$

Can you spot the Higgs?



Why Monte Carlos?

LHC experiments require
sound understanding of signals and *backgrounds*.



Full detector simulation.



Fully exclusive hadronic final state.

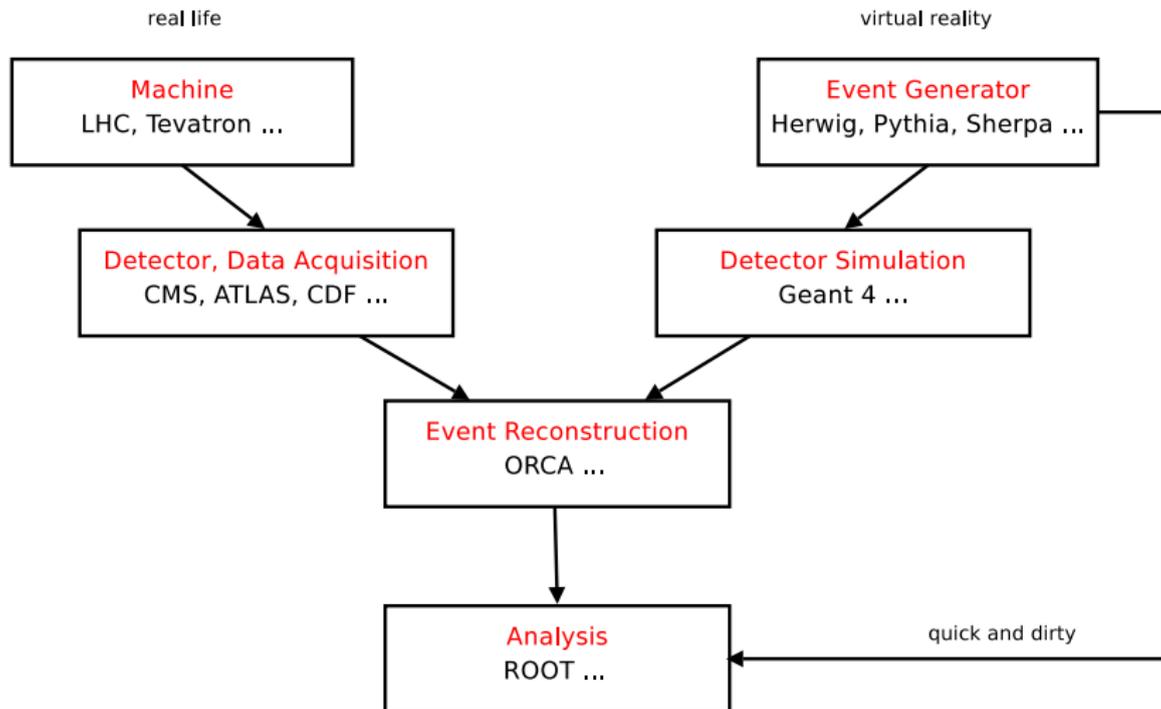


Monte Carlo event generator with
parton shower, hadronization model, decays of unstable
particles.



Parton level computations.

Experiment and Simulation

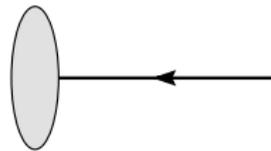
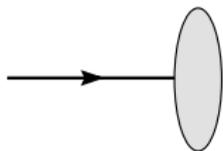


- ▶ Complex final states in full detail (jets).
- ▶ Arbitrary observables and cuts from final states.
- ▶ Studies of new physics models.

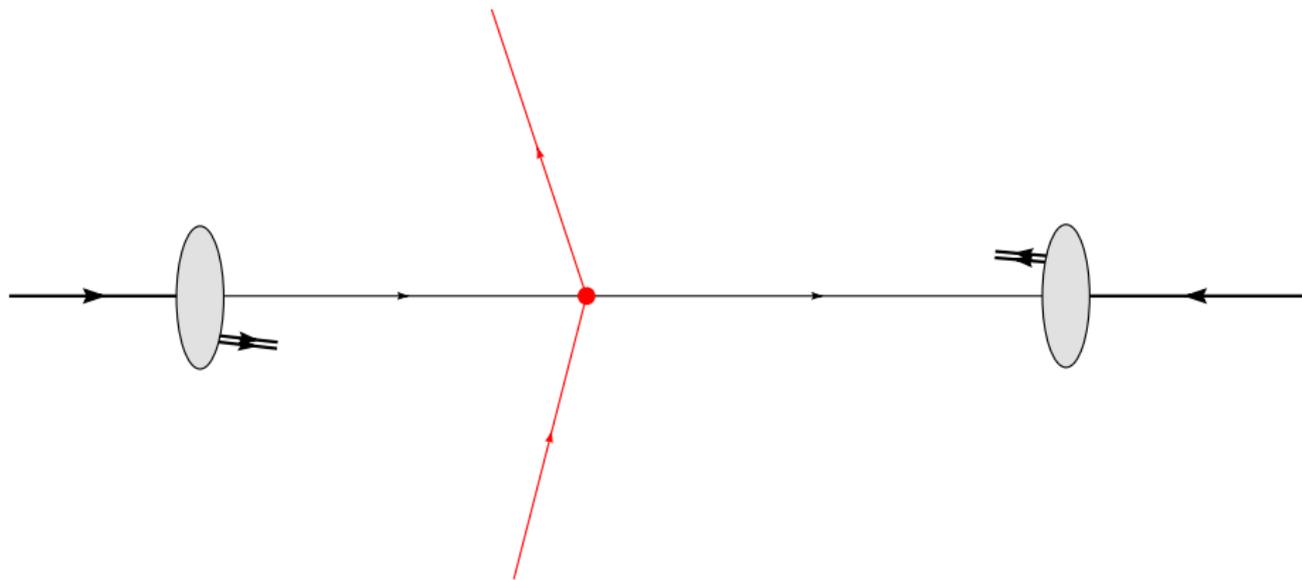
- ▶ Rates and topologies of final states.
- ▶ Background studies.
- ▶ Detector Design.
- ▶ Detector Performance Studies (Acceptance).

- ▶ *Obvious* for calculation of observables on the quantum level

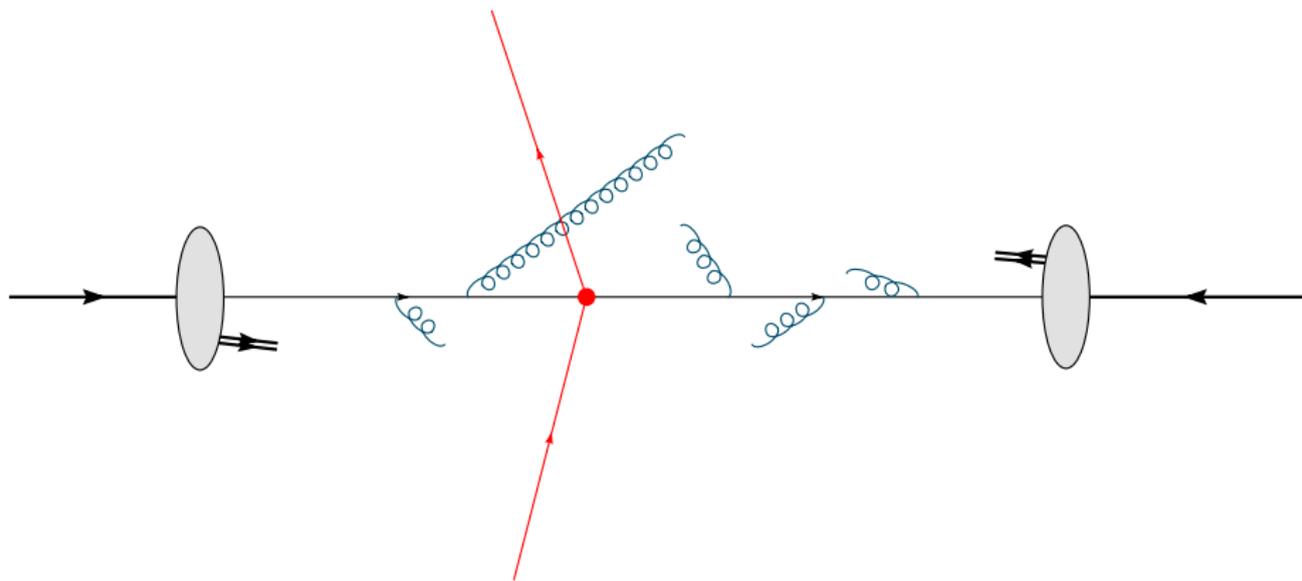
$$|A|^2 \longrightarrow \text{Probability.}$$



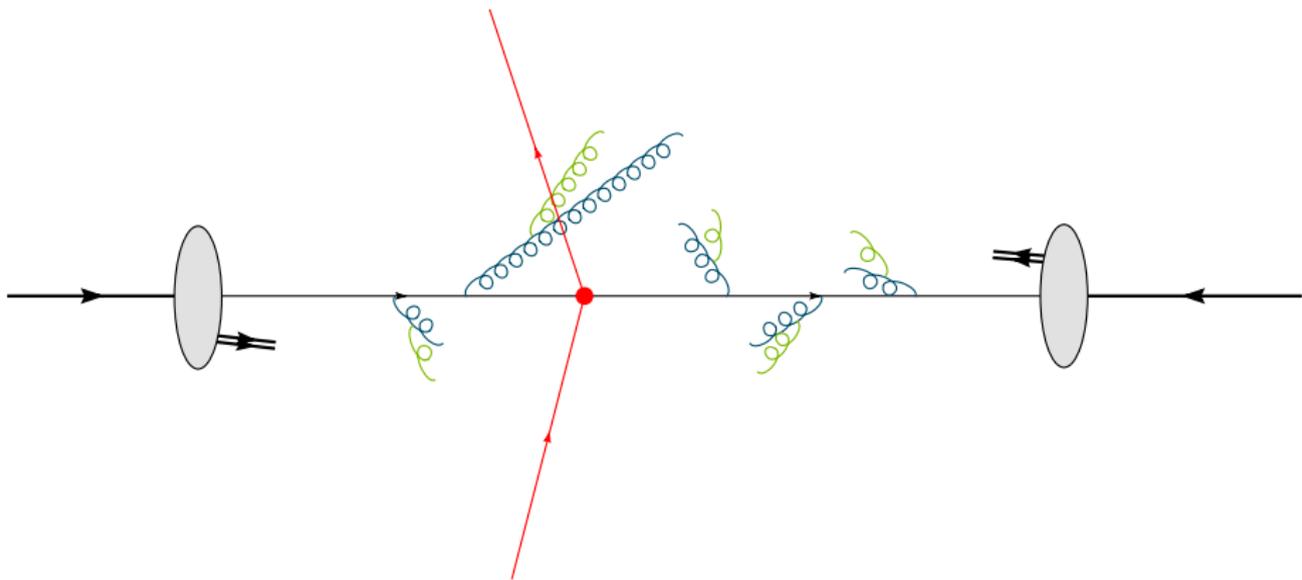
pp Event Generator



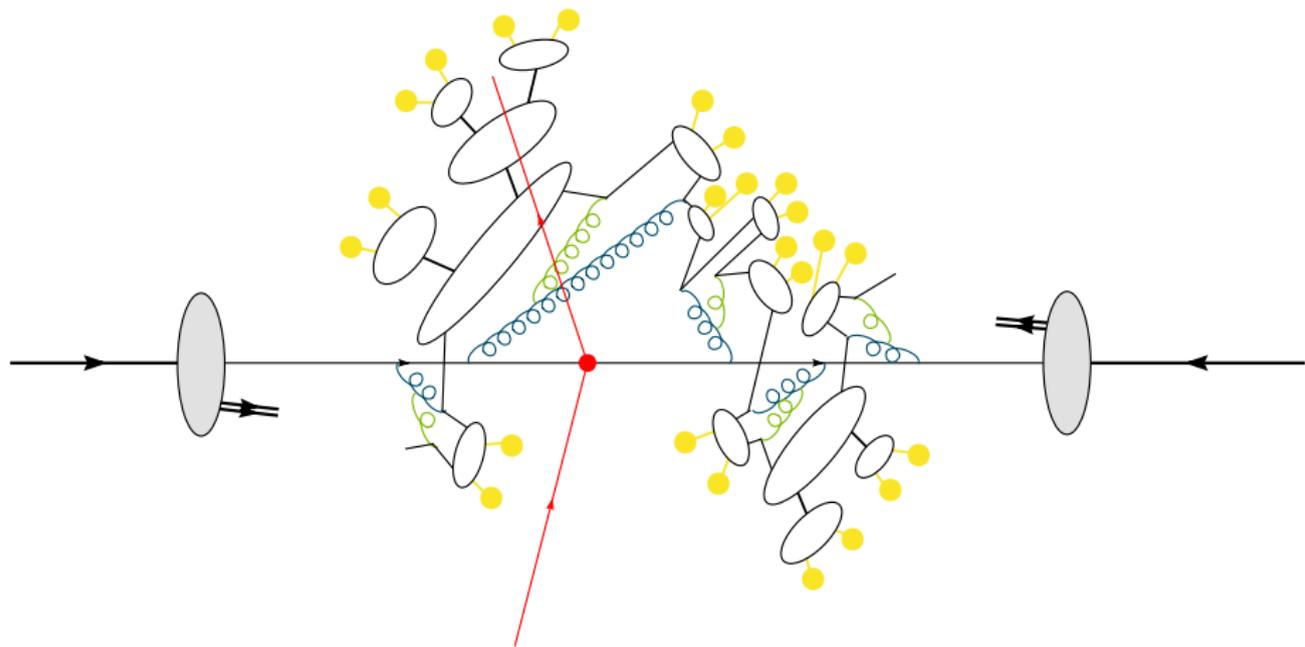
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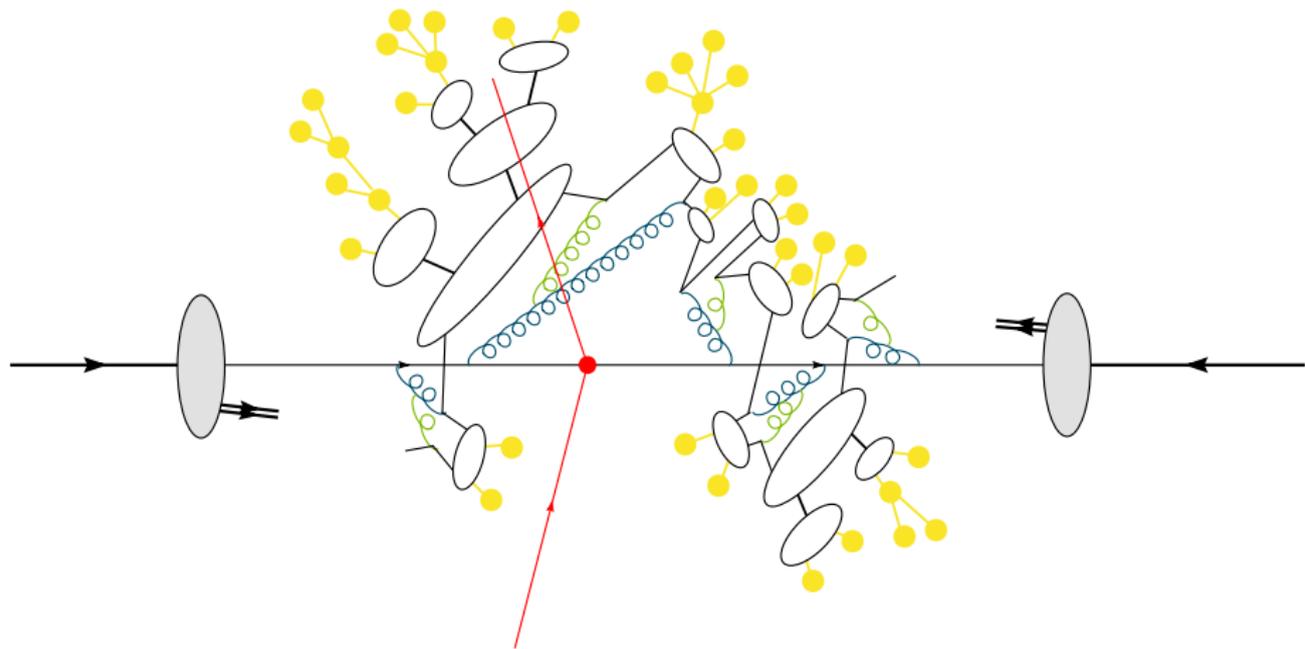
pp Event Generator



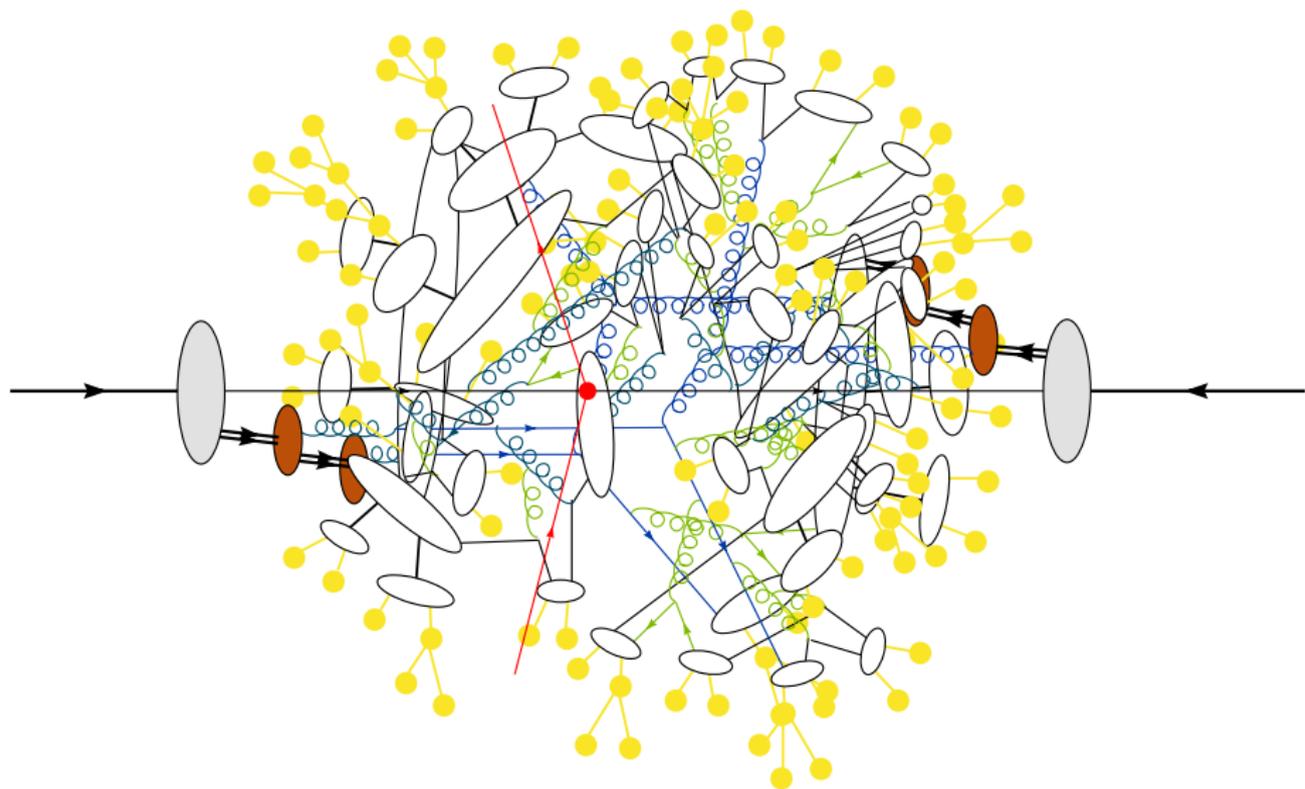
pp Event Generator



pp Event Generator



pp Event Generator



Partonic cross section from Feynman diagrams

$$d\sigma = d\sigma_{\text{hard}} dP(\text{partons} \rightarrow \text{hadrons})$$

Note, that

$$\int dP(\text{partons} \rightarrow \text{hadrons}) = 1 ,$$

- ▶ σ remains unchanged
- ▶ introduce realistic fluctuations into distributions.

Partonic cross section from Feynman diagrams

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Note, that

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- ▶ σ remains unchanged
- ▶ introduce realistic fluctuations into distributions.

Simulation steps governed by different scales

→ separation into ($Q_0 \approx 1 \text{ GeV} > \Lambda_{\text{QCD}}$)

$$\begin{aligned} dP(\text{partons} \rightarrow \text{hadrons}) = & dP(\text{resonance decays}) && [\Gamma > Q_0] \\ & \times dP(\text{parton shower}) && [\text{TeV} \rightarrow Q_0] \\ & \times dP(\text{hadronisation}) && [\sim Q_0] \\ & \times dP(\text{hadronic decays}) && [O(\text{MeV})] \end{aligned}$$

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Quite complicated integration.

$$\begin{aligned} dP(\text{partons} \rightarrow \text{hadrons}) &= dP(\text{resonance decays}) && [\Gamma > Q_0] \\ &\times dP(\text{parton shower}) && [\text{TeV} \rightarrow Q_0] \\ &\times dP(\text{hadronisation}) && [\sim Q_0] \\ &\times dP(\text{hadronic decays}) && [O(\text{MeV})] \end{aligned}$$

Quite complicated integration.

Monte Carlo is the only choice.

Monte Carlo Methods

Introduction to the most important MC sampling
(= integration) techniques.

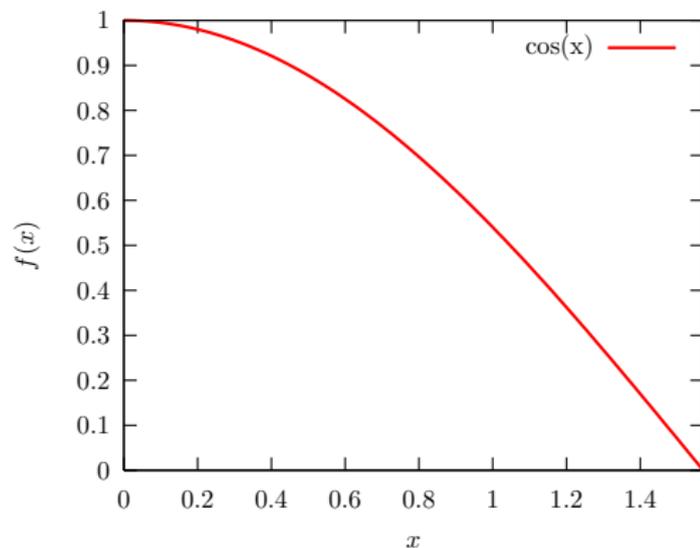
1. Hit and miss.
2. Simple MC integration.
3. (Some) methods of variance reduction.
4. Multichannel.

Probability density:

$$dP = f(x) dx$$

is probability to find value x .

Example: $f(x) = \cos(x)$.



Probability \sim Area

Probability density:

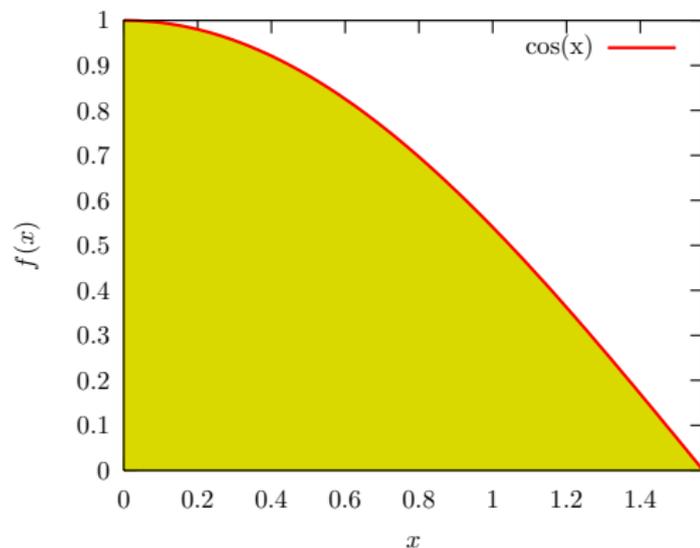
$$dP = f(x) dx$$

is probability to find value x .

$$F(x) = \int_{x_0}^x f(x) dx$$

is called *probability distribution*.

Example: $f(x) = \cos(x)$.



Probability \sim Area

Hit and Miss

Hit and miss method:

- ▶ throw N random points (x, y) into region.
- ▶ Count hits N_{hit} ,
i.e. whenever $y < f(x)$.

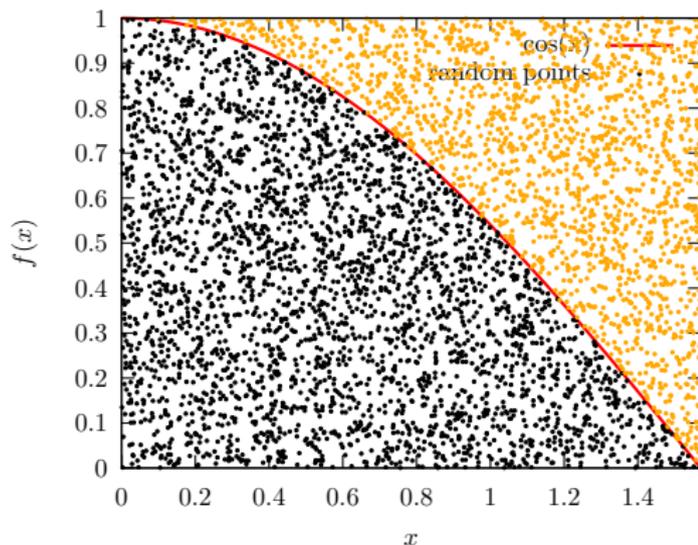
Then

$$I \approx V \frac{N_{\text{hit}}}{N}.$$

approaches 1 again in our example.

Every **accepted** value of x can be considered an **event** in this picture. As $f(x)$ is the 'histogram' of x , it seems obvious that the x values are distributed as $f(x)$ from this picture.

Example: $f(x) = \cos(x)$.



This method is used in many event generators. However, it is not sufficient as such.

- ▶ Can handle any density $f(x)$, however wild and unknown it is.
- ▶ $f(x)$ should be bounded from above.
- ▶ Sampling will be very *inefficient* whenever $\text{Var}(f)$ is large.

Improvements go under the name **variance reduction** as they improve the error of the crude MC at the same time.

Mean value theorem of integration:

$$\begin{aligned} I &= \int_{x_0}^{x_1} f(x) dx \\ &= (x_1 - x_0) \langle f(x) \rangle \\ &\approx (x_1 - x_0) \frac{1}{N} \sum_{i=1}^N f(x_i) \end{aligned}$$

(Riemann integral).

Sum doesn't depend on ordering

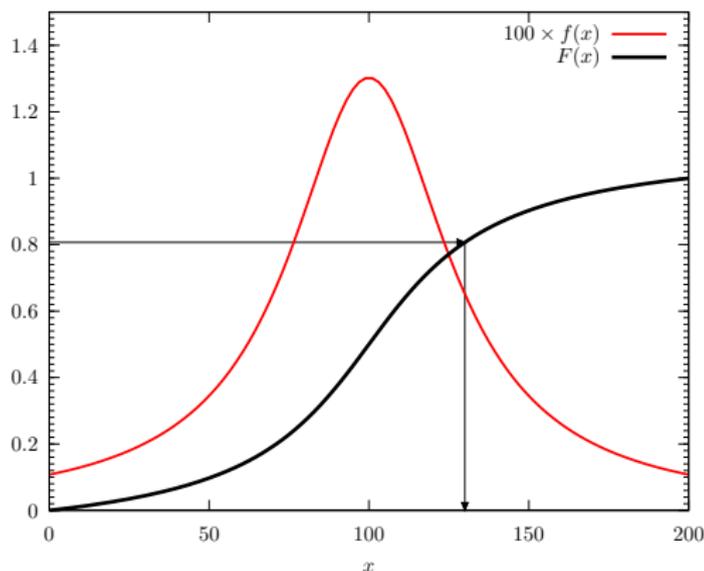
→ randomize x_i .

Yields a flat distribution of events x_i ,
but weighted with *weight* $f(x_i)$ (→ unweighting).

Inverting the Integral

- ▶ Probability density $f(x)$. Not necessarily normalized.
- ▶ Integral $F(x)$ known,
- ▶ $P(x < x_s) = F(x_s)$.
- ▶ Probability = 'area', distributed evenly,

$$\int_{x_0}^x dP = r \cdot \text{area}$$



Sample x according to $f(x)$ with

$$x = F^{-1} \left[F(x_0) + r(F(x_1) - F(x_0)) \right].$$

Sample x according to $f(x)$ with

$$x = F^{-1} \left[F(x_0) + r(F(x_1) - F(x_0)) \right].$$

Optimal method, but we need to know

- ▶ The integral $F(x) = \int f(x) dx$,
- ▶ It's inverse $F^{-1}(y)$.

That's rarely the case for real problems.

But very powerful in combination with other techniques.

Importance sampling

Error on Crude MC $\sigma_{MC} = \sigma/\sqrt{N}$.

\implies Reduce error by reducing variance of integrand.

Importance sampling

Error on Crude MC $\sigma_{MC} = \sigma/\sqrt{N}$.

⇒ Reduce error by reducing variance of integrand.

Idea: *Divide out the singular structure.*

$$I = \int f dV = \int \frac{f}{p} p dV \approx \left\langle \frac{f}{p} \right\rangle \pm \sqrt{\frac{\langle f^2/p^2 \rangle - \langle f/p \rangle^2}{N}}.$$

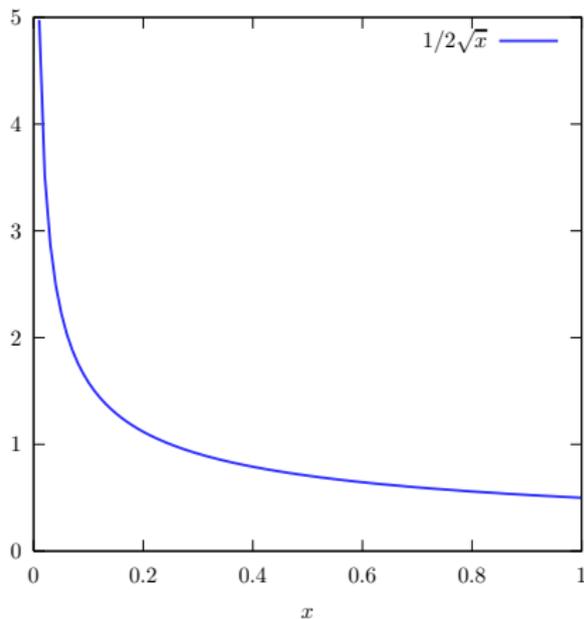
where we have chosen $\int p dV = 1$ for convenience.

Note: need to sample flat in $p dV$, so we better know $\int p dV$ and it's inverse.

Importance sampling — better example

More interesting for **divergent integrands**, eg

$$\frac{1}{2\sqrt{x}},$$



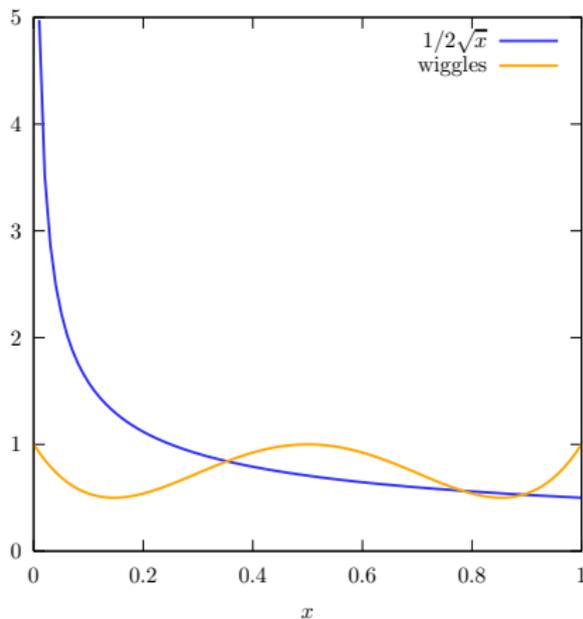
Importance sampling — better example

More interesting for **divergent integrands**, eg

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with some wiggles,

$$p(x) = 1 - 8x + 40x^2 - 64x^3 + 32x^4.$$



Importance sampling — better example

More interesting for **divergent integrands**, eg

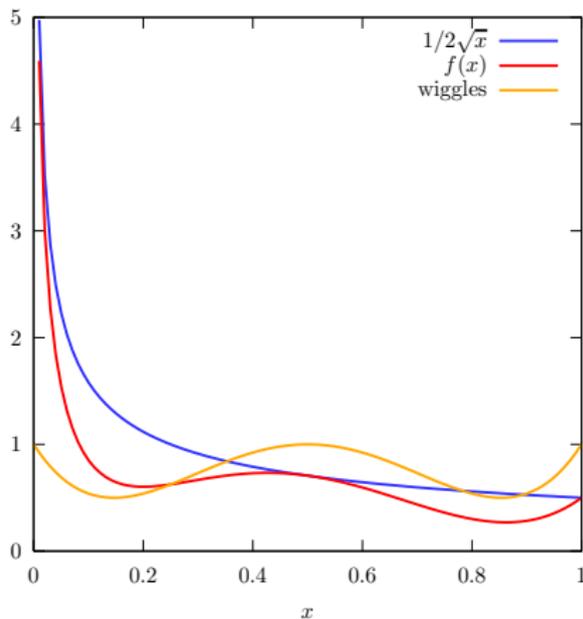
$$\frac{1}{2\sqrt{x}},$$

with some wiggles,

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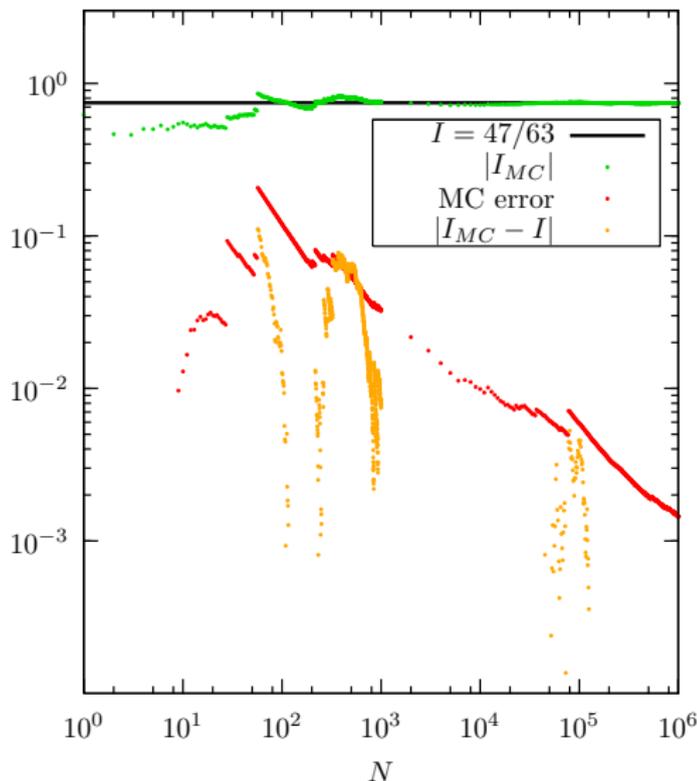
i.e. we want to integrate

$$f(x) = \frac{p(x)}{2\sqrt{x}}.$$



Importance sampling — better example

- ▶ Crude MC gives result in reasonable 'time'.
- ▶ Error a bit unstable.
- ▶ Event generation with maximum weight $w_{\max} = 20$. (that's arbitrary.)
- ▶ hit/miss/events with $(w > w_{\max}) = 36566/963434/617$ with 1M generated events.

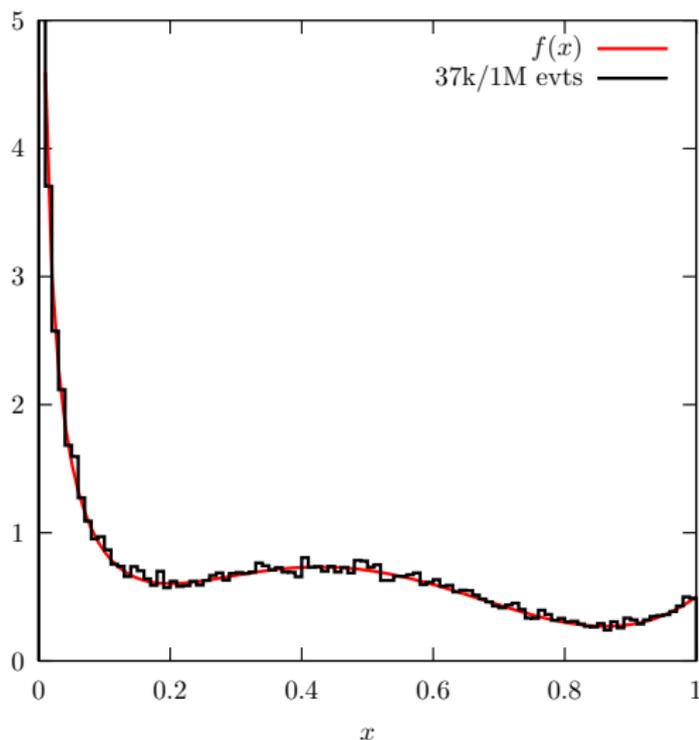


Importance sampling — example

Want events:

use hit+mass variant
here:

- ▶ Choose new random number r
- ▶ $w = f(x)$ in this case.
- ▶ if $r < w/w_{\max}$ then “hit”.
- ▶ MC efficiency = hit/ N .
- ▶ Efficiency for MC events only 3.7%.
- ▶ Note the wiggly histogram.



Importance sampling — example

Now importance sampling, i.e. divide out $1/2\sqrt{x}$.

$$\begin{aligned}\int_0^1 \frac{p(x)}{2\sqrt{x}} dx &= \int_0^1 \left(\frac{p(x)}{2\sqrt{x}} / \frac{1}{2\sqrt{x}} \right) \frac{dx}{2\sqrt{x}} \\ &= \int_0^1 p(x) d\sqrt{x} \\ &= \int_0^1 p(x(\rho)) d\rho \\ &= \int_0^1 1 - 8\rho^2 + 40\rho^4 - 64\rho^6 + 32\rho^8 d\rho\end{aligned}$$

so,

$$\rho = \sqrt{x}, \quad d\rho = \frac{dx}{2\sqrt{x}}$$

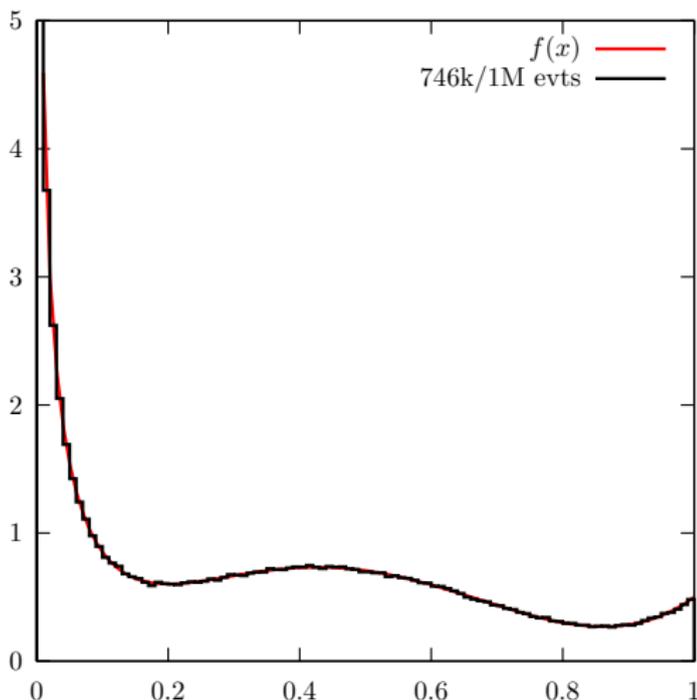
x sampled with *inverting the integral* from flat random numbers ρ , $x = \rho^2$.

Importance sampling — example

$$\int_0^1 \frac{p(x)}{2\sqrt{x}} dx = \int_0^1 p(x(\rho)) d\rho$$

with

$$\rho = \sqrt{x}, \quad d\rho = \frac{dx}{2\sqrt{x}}$$



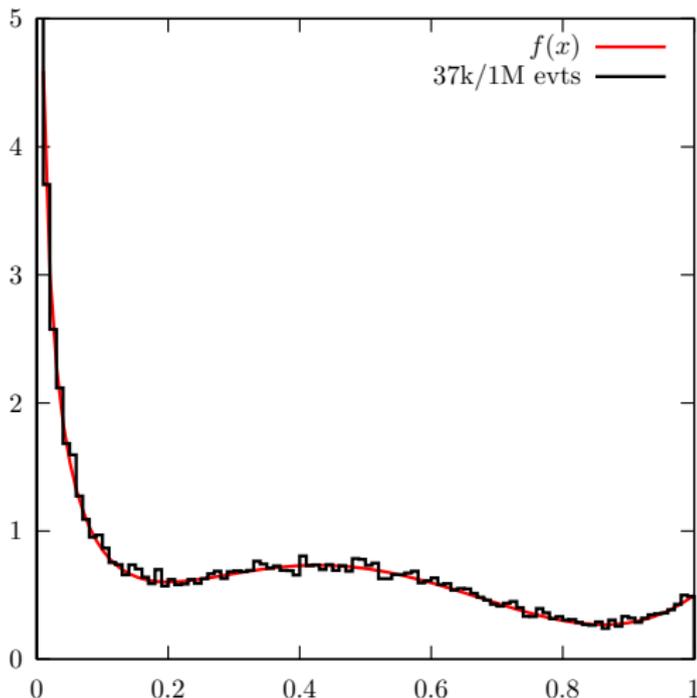
Events generated with $w_{\max} = 1$, as $p(x) \leq 1$, no guesswork needed here! Now, we get **74.6%** MC efficiency.

Importance sampling — example

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with

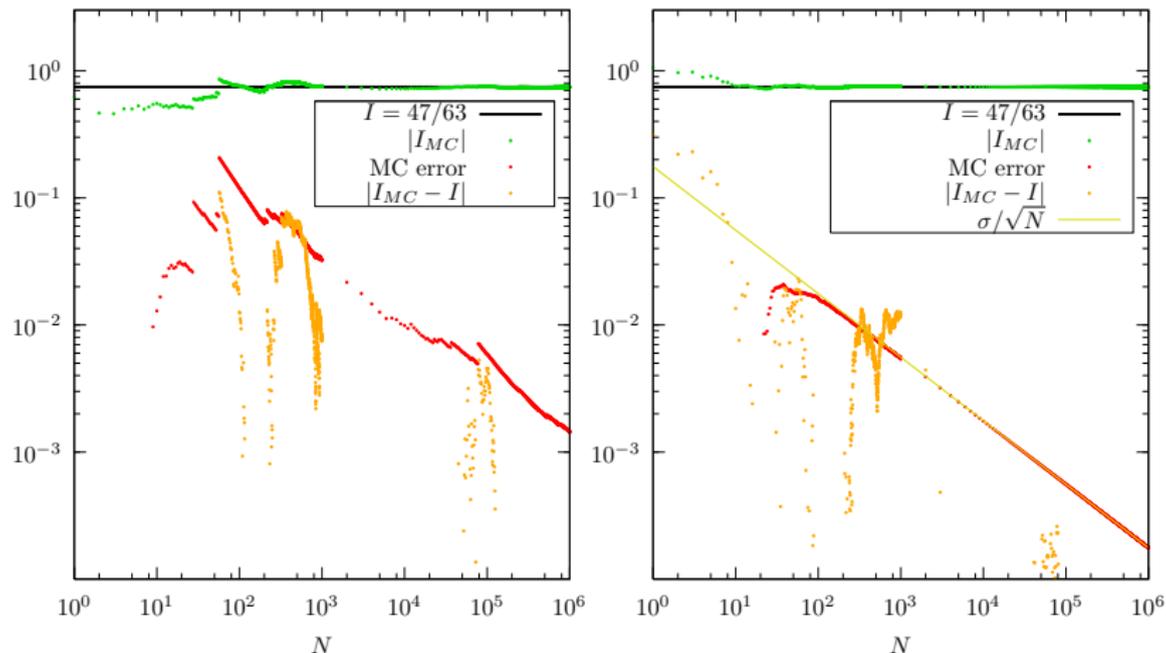
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Events generated with $w_{\max} = 1$, as $p(x) \leq 1$, no guesswork needed here! Now, we get 74.6% MC efficiency.

... as opposed to 3.7%.

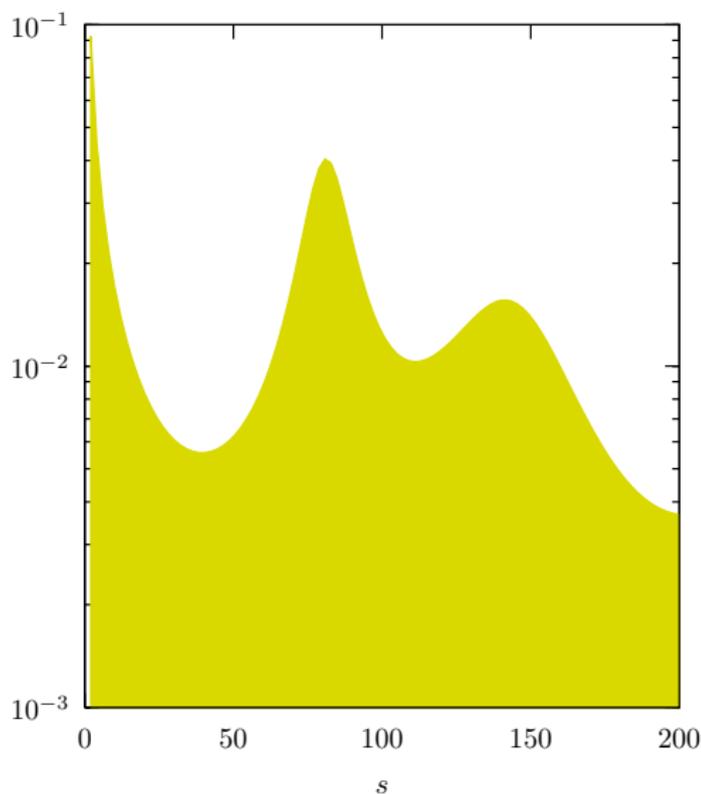
Crude MC vs Importance sampling.



100× more events needed to reach same accuracy.

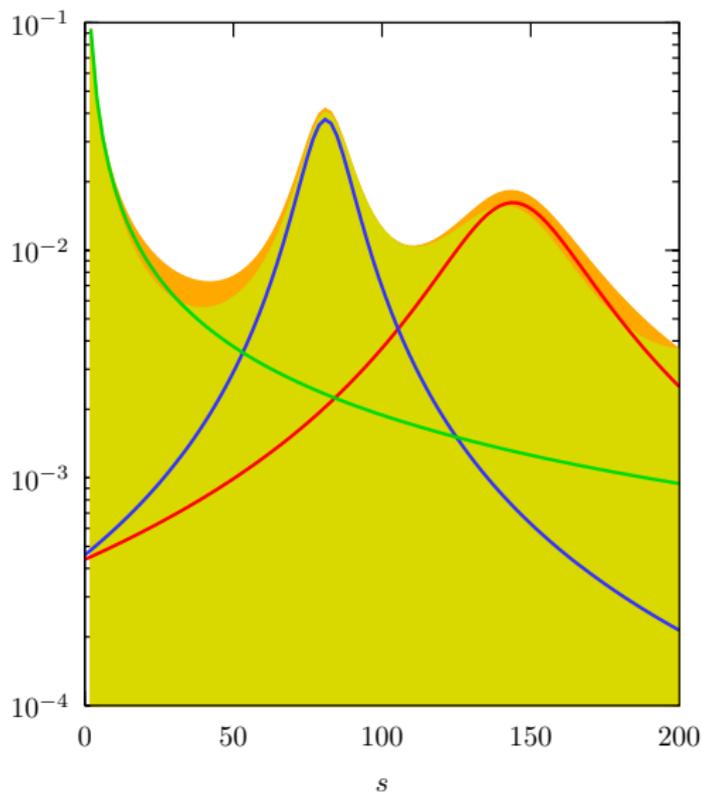
Typical problem:

- ▶ $f(s)$ has multiple peaks (\times wiggles from ME).



Typical problem:

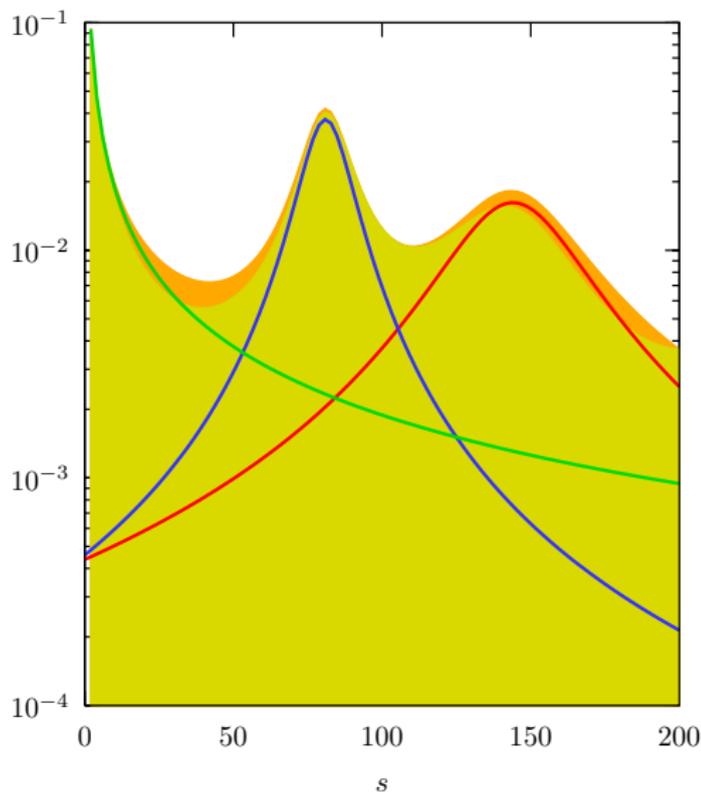
- ▶ $f(s)$ has multiple peaks (\times wiggles from ME).
- ▶ Usually have some idea of the peak structure.



Typical problem:

- ▶ $f(s)$ has multiple peaks (\times wiggles from ME).
- ▶ Usually have some idea of the peak structure.
- ▶ Encode this in sum of sample functions $g_i(s)$ with weights $\alpha_i, \sum_i \alpha_i = 1$.

$$g(s) = \sum_i \alpha_i g_i(s) .$$



Now rewrite

$$\begin{aligned}\int_{s_0}^{s_1} f(s) ds &= \int_{s_0}^{s_1} \frac{f(s)}{g(s)} g(s) ds \\ &= \int_{s_0}^{s_1} \frac{f(s)}{g(s)} \sum_i \alpha_i g_i(s) ds \\ &= \sum_i \alpha_i \int_{s_0}^{s_1} \frac{f(s)}{g(s)} g_i(s) ds\end{aligned}$$

Now $g_i(s) ds = d\rho_i$ (inverting the integral).

Now rewrite

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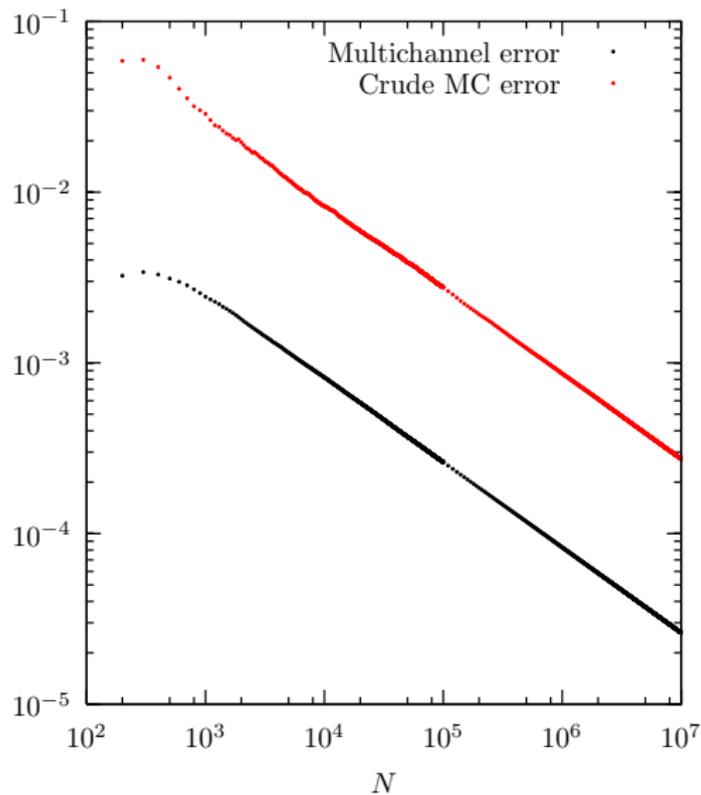
Now $g_i(s) ds = d\rho_i$ (inverting the integral).

Select the distribution $g_i(s)$ you'd like to sample next event from acc to weights α_i .

α_i can be optimized after a number of trials.

Multichannel MC

Works quite well:

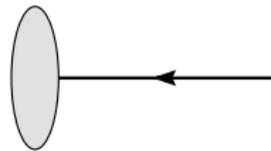
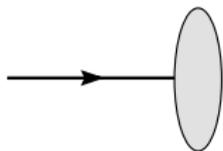


Some Remarks/Real Life MC

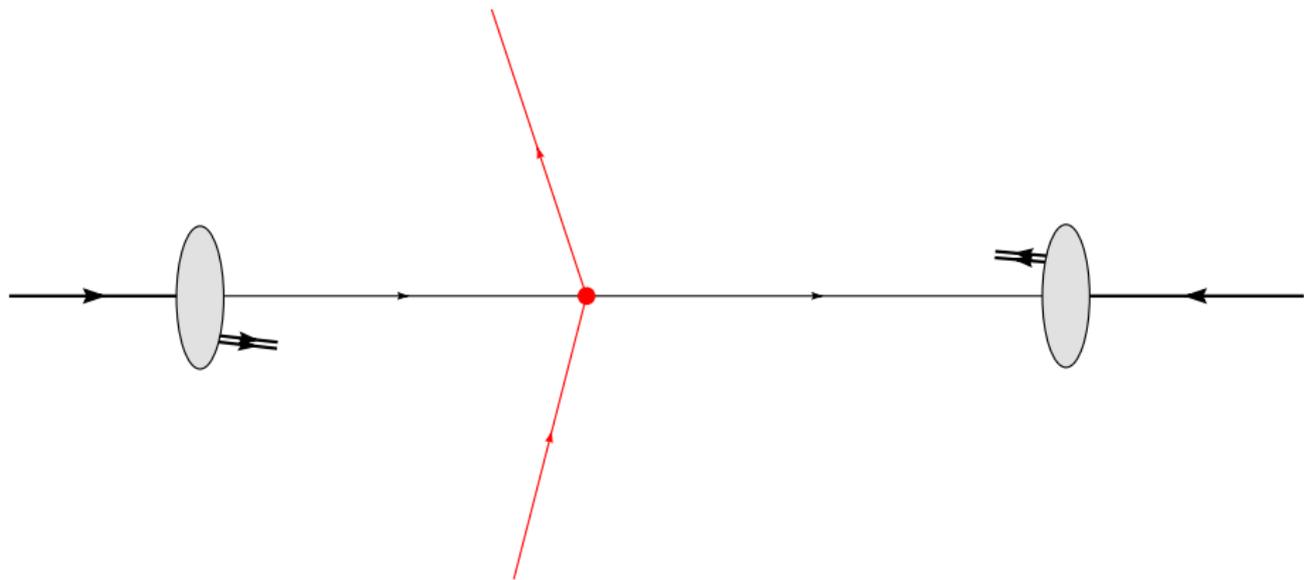
- ▶ Didn't discuss random number generators. Please make sure to use 'good' random numbers.
- ▶ Didn't discuss *stratified sampling* (VEGAS).
Sample where variance is biggest.
(not necessarily where PS is most populated).
- ▶ Only discussed one-dimensional case here. N -particle PS has $3N - 4$ dimensions...
- ▶ Didn't discuss tools geared towards this, like RAMBO (generates flat N particles PS).
- ▶ generalisation straightforward, particularly
 $\text{MCError} \sim \frac{1}{\sqrt{N}}$,
compare eg Trapezium rule $\text{Error} \sim \frac{1}{N^{2/D}}$.
- ▶ Many important techniques covered here in detail! Should be good starting point.

Hard Scattering

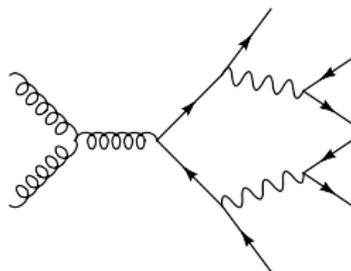
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Hard scattering

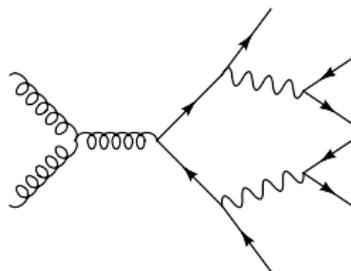


- ▶ Perturbation theory/Feynman diagrams give us (fairly accurate) final states for a few number of legs ($O(1)$).



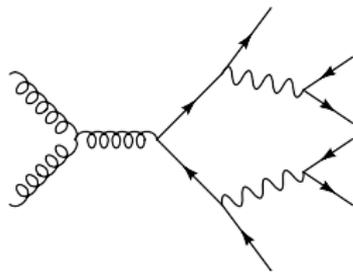
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- ▶ OK for very inclusive observables.
- ▶ Starting point for further simulation.
- ▶ Want exclusive final state at the LHC ($O(100)$).

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- ▶ OK for very inclusive observables.
- ▶ Starting point for further simulation.
- ▶ Want exclusive final state at the LHC ($O(100)$).
- ▶ Want arbitrary cuts.
- ▶ → use Monte Carlo methods.

Where do we get (LO) $|M|^2$ from?

- ▶ Most/important simple processes (SM) are 'built in'.
- ▶ Calculate yourself (≤ 3 particles in final state).
- ▶ Matrix element generators:
 - ▶ MadGraph/MadEvent.
 - ▶ Comix/AMEGIC (part of Sherpa).
 - ▶ HELAC/PHEGAS.
 - ▶ Whizard.
 - ▶ CalcHEP/CompHEP.

generate code or event files that can be further processed.

- ▶ → FeynRules interface to ME generators.

Cross section formula

From Matrix element, we calculate

$$\sigma = \int f_i(x_1, \mu^2) f_j(x_2, \mu^2) \frac{1}{F} \sum \overline{|M|^2} \, dx_1 dx_2 d\Phi_n ,$$

Cross section formula

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now,

$$\frac{1}{F} dx_1 dx_2 d\Phi_n = J(\vec{x}) \prod_{i=1}^{3n-2} dx_i \quad \left(d\Phi_n = (2\pi)^4 \delta^{(4)}(\dots) \prod_{i=1}^n \frac{d^3\vec{p}}{(2\pi)^3 2E_i} \right)$$

such that

$$\begin{aligned} \sigma &= \int g(\vec{x}) d^{3n-2}\vec{x} , & \left(g(\vec{x}) = J(\vec{x}) f_i f_j \overline{\sum} |M|^2 \Theta(\text{cuts}) \right) \\ &= \frac{1}{N} \sum_{i=1}^N \frac{g(\vec{x}_i)}{p(\vec{x}_i)} = \frac{1}{N} \sum_{i=1}^N w_i . \end{aligned}$$

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We generate **events** \vec{x}_i with **weights** w_i .

Mini event generator

- ▶ We generate pairs (\vec{x}_i, w_i) .

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Generate events with same frequency as in nature!

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- ▶ Use immediately to book weighted histogram of arbitrary observable (possibly with additional cuts!)
- ▶ Keep event \vec{x}_i with probability

$$P_i = \frac{w_i}{w_{\max}},$$

where w_{\max} has to be chosen sensibly.

→ reweighting, when $\max(w_i) = \bar{w}_{\max} > w_{\max}$, as

$$P_i = \frac{w_i}{\bar{w}_{\max}} = \frac{w_i}{w_{\max}} \cdot \frac{w_{\max}}{\bar{w}_{\max}},$$

i.e. reject events with probability $(w_{\max}/\bar{w}_{\max})$ afterwards.
(can be ignored when #(events with $w_i > \bar{w}_{\max}$) small.)

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Generate events with same frequency as in nature!

Some comments:

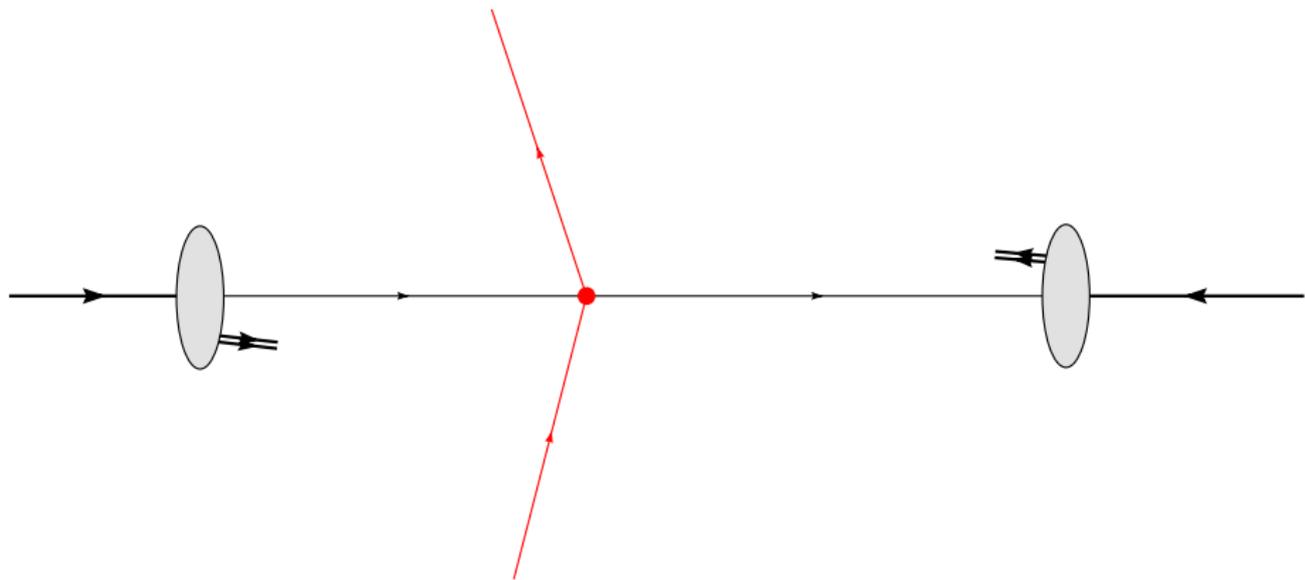
- ▶ Use techniques from above to generate events efficiently.
Goal: small variance in w_i distribution!

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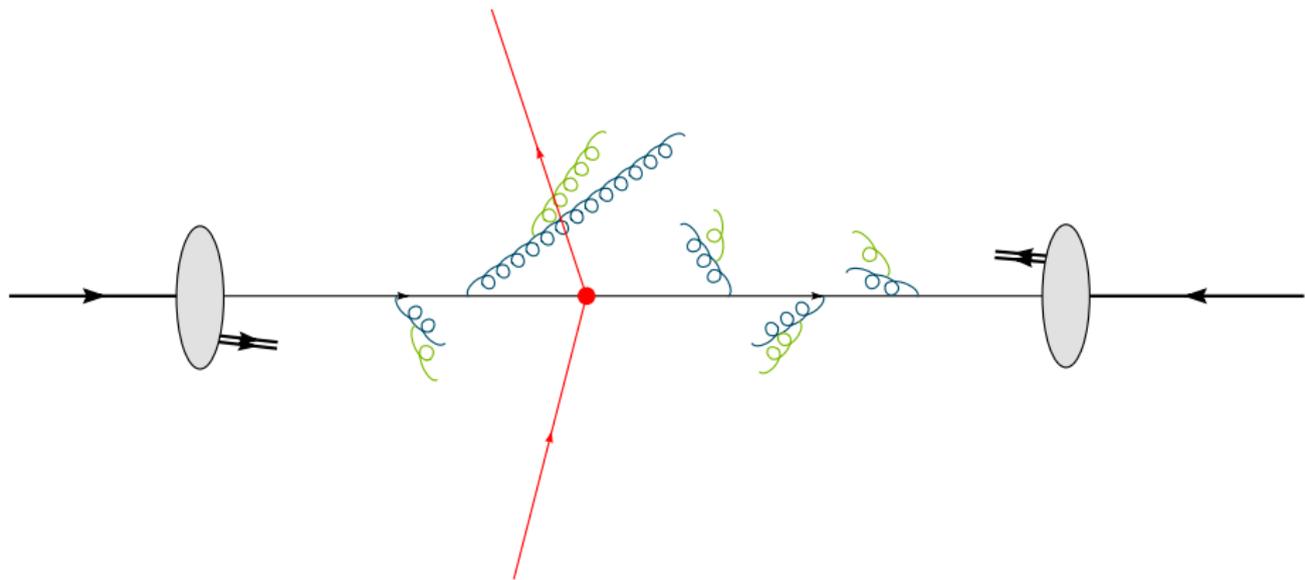
- ▶ Use techniques from above to generate events efficiently.
Goal: small variance in w_i distribution!
- ▶ Clear from above: efficient generation closely tied to knowledge of $f(\vec{x}_i)$, *i.e.* the matrix element's propagator structure.
→ build phase space generator already while generating ME's automatically.

Parton Showers

Hard matrix element



Hard matrix element \rightarrow parton showers



Quarks and gluons in final state, pointlike.

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- ▶ Know short distance (short time) fluctuations from matrix element/Feynman diagrams: $Q \sim \text{few GeV to } O(\text{TeV})$.

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Generated from emissions *ordered* in Q .
Soft and/or collinear emissions.

Good starting point: $e^+e^- \rightarrow q\bar{q}g$:

Final state momenta in one plane (orientation usually averaged).

Write momenta in terms of

$$x_i = \frac{2p_i \cdot q}{Q^2} \quad (i = 1, 2, 3),$$

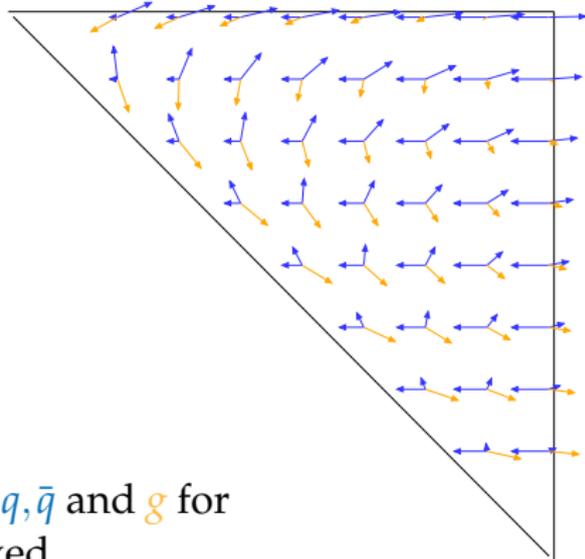
$$0 \leq x_i \leq 1, x_1 + x_2 + x_3 = 2,$$

$$q = (Q, 0, 0, 0),$$

$$Q \equiv E_{cm}.$$

Fig: momentum configuration of q, \bar{q} and g for given point (x_1, x_2) , \bar{q} direction fixed.

$(x_1, x_2) = (x_q, x_{\bar{q}})$ -plane:

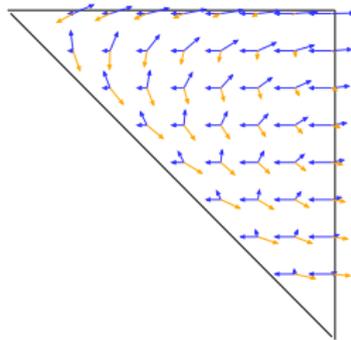
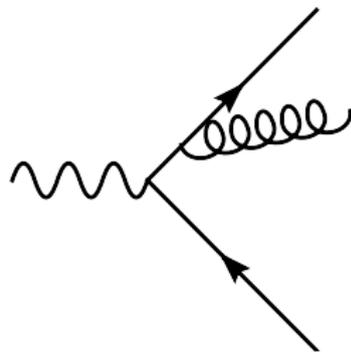


Differential cross section:

$$\frac{d\sigma}{dx_1 dx_2} = \sigma_0 \frac{C_F \alpha_S}{2\pi} \frac{x_1 + x_2}{(1-x_1)(1-x_2)}$$

Collinear singularities: $x_1 \rightarrow 1$ or $x_2 \rightarrow 1$.

Soft singularity: $x_1, x_2 \rightarrow 1$.



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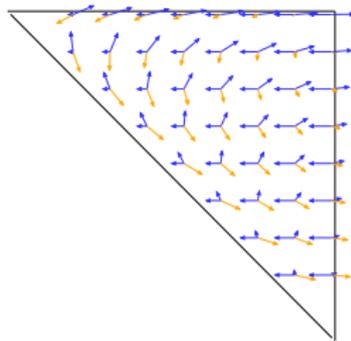
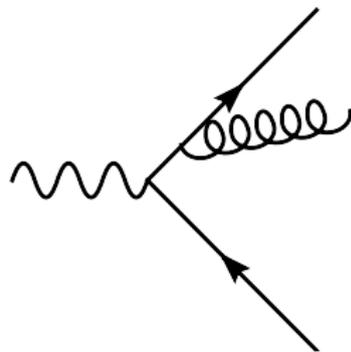
Collinear singularities: $x_1 \rightarrow 1$ or $x_2 \rightarrow 1$.

Soft singularity: $x_1, x_2 \rightarrow 1$.

Rewrite in terms of x_3 and $\theta = \angle(q, g)$:

$$\frac{d\sigma}{d\cos\theta dx_3} = \sigma_0 \frac{C_F \alpha_S}{2\pi} \left[\frac{2}{\sin^2\theta} \frac{1 + (1-x_3)^2}{x_3} - x_3 \right]$$

Singular as $\theta \rightarrow 0$ and $x_3 \rightarrow 0$.



Can separate into two jets as

$$\begin{aligned}\frac{2d\cos\theta}{\sin^2\theta} &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\theta}{1+\cos\theta} \\ &= \frac{d\cos\theta}{1-\cos\theta} + \frac{d\cos\bar{\theta}}{1-\cos\bar{\theta}} \\ &\approx \frac{d\theta^2}{\theta^2} + \frac{d\bar{\theta}^2}{\bar{\theta}^2}\end{aligned}$$

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So, we rewrite $d\sigma$ in collinear limit as

$$d\sigma = \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} C_F \frac{1+(1-z)^2}{z^2} dz$$

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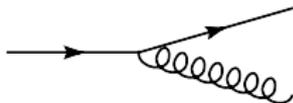
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with DGLAP splitting function $P(z)$.

Universal DGLAP splitting kernels for collinear limit:

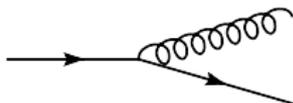
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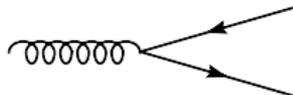
$$P_{q \to qg}(z) = C_F \frac{1+z^2}{1-z}$$



$$P_{g \to gg}(z) = C_A \frac{(1-z(1-z))^2}{z(1-z)}$$



$$P_{q \to gq}(z) = C_F \frac{1+(1-z)^2}{z}$$



$$P_{g \to qq}(z) = T_R(1-2z(1-z))$$

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$$d\sigma = \sigma_0 \sum_{\text{jets}} \frac{d\theta^2}{\theta^2} \frac{\alpha_S}{2\pi} P(z) dz$$

Note: Other variables may equally well characterize the collinear limit:

$$\frac{d\theta^2}{\theta^2} \sim \frac{dQ^2}{Q^2} \sim \frac{dp_{\perp}^2}{p_{\perp}^2} \sim \frac{d\tilde{q}^2}{\tilde{q}^2} \sim \frac{dt}{t}$$

whenever $Q^2, p_{\perp}^2, t \rightarrow 0$ means “collinear”.

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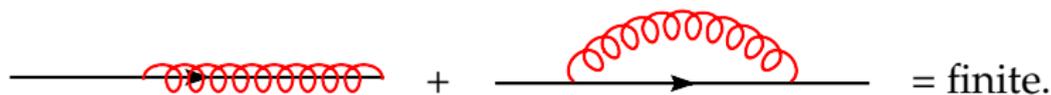
whenever $Q^2, p_{\perp}^2, t \rightarrow 0$ means “collinear”.

- ▶ θ : HERWIG
- ▶ Q^2 : PYTHIA ≤ 6.3 , old SHERPA.
- ▶ p_{\perp} : PYTHIA ≥ 6.4 , ARIADNE, Catani–Seymour showers in HERWIG++ and SHERPA.
- ▶ \tilde{q} : Herwig++.

Need to introduce **resolution** t_0 , e.g. a cutoff in p_{\perp} . Prevent us from the singularity at $\theta \rightarrow 0$.

Emissions below t_0 are **unresolvable**.

Finite result due to virtual corrections:



The diagram shows two Feynman diagrams separated by a plus sign, followed by an equals sign and the word "finite". The first diagram is a horizontal line with a red wavy loop below it. The second diagram is a horizontal line with a red wavy loop above it that is semi-circular and ends with an arrow pointing to the right.

unresolvable + virtual emissions are included in Sudakov form factor via unitarity (see below!).

Starting point: factorisation in collinear limit, single emission.

$$\sigma_{2+1}(t_0) = \sigma_2(t_0) \int_{t_0}^t \frac{dt'}{t'} \int_{z_-}^{z_+} dz \frac{\alpha_S}{2\pi} \hat{P}(z) = \sigma_2(t_0) \int_{t_0}^t dt W(t) .$$

Towards multiple emissions

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Simple example:

Multiple photon emissions, strongly ordered in t .

We want

$$W_{\text{sum}} = \sum_{n=1} W_{2+n} = \frac{\int \left| \text{diagram}_1 \right|^2 d\Phi_1 + \int \left| \text{diagram}_2 \right|^2 d\Phi_2 + \int \left| \text{diagram}_3 \right|^2 d\Phi_3 + \dots}{\left| \text{diagram}_0 \right|^2}$$

for any number of emissions.

Towards multiple emissions

$$(n=1) \text{ } \begin{array}{c} \nearrow \\ \bullet \\ \searrow \\ \text{---} \end{array}$$

$$W_{2+1} = \left(\int \left| \begin{array}{c} \nearrow \\ \text{---} \\ \searrow \end{array} \right|^2 + \left| \begin{array}{c} \nearrow \\ \text{---} \\ \searrow \end{array} \right|^2 d\Phi_1 \right) / \left| \begin{array}{c} \nearrow \\ \bullet \\ \searrow \end{array} \right|^2 = \frac{2}{1!} \int_{t_0}^t dt W(t) .$$

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$$(n = 2) \text{ } \bullet \begin{array}{l} \nearrow \\ \text{---} \\ \searrow \end{array}$$

$$W_{2+2} = \left(\int \left| \begin{array}{l} \nearrow \\ \text{---} \\ \searrow \end{array} \right|^2 + \left| \begin{array}{l} \nearrow \\ \text{---} \\ \searrow \end{array} \right|^2 + \left| \begin{array}{l} \nearrow \\ \text{---} \\ \searrow \end{array} \right|^2 + \left| \begin{array}{l} \nearrow \\ \text{---} \\ \searrow \end{array} \right|^2 d\Phi_2 \right) / \left| \begin{array}{l} \nearrow \\ \bullet \\ \searrow \end{array} \right|^2$$
$$= 2^2 \int_{t_0}^t dt' \int_{t_0}^{t'} dt'' W(t') W(t'') = \frac{2^2}{2!} \left(\int_{t_0}^t dt W(t) \right)^2.$$

We used

$$\int_{t_0}^t dt_1 \dots \int_{t_0}^{t_{n-1}} dt_n W(t_1) \dots W(t_n) = \frac{1}{n!} \left(\int_{t_0}^t dt W(t) \right)^n.$$

Towards multiple emissions

Easily generalized to n emissions  by induction. *i.e.*

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So, in total we get

$$\sigma_{>2}(t_0) = \sigma_2(t_0) \sum_{k=1}^{\infty} \frac{2^k}{k!} \left(\int_{t_0}^t dt W(t) \right)^k = \sigma_2(t_0) \left(e^{2 \int_{t_0}^t dt W(t)} - 1 \right)$$

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Sudakov Form Factor

$$\Delta(t_0, t) = \exp \left[- \int_{t_0}^t dt W(t) \right]$$

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Sudakov Form Factor in QCD

$$\Delta(t_0, t) = \exp \left[- \int_{t_0}^t dt W(t) \right] = \exp \left[- \int_{t_0}^t \frac{dt}{t} \int_{z_-}^{z_+} \frac{\alpha_S(z, t)}{2\pi} \hat{P}(z, t) dz \right]$$

Note that

$$\begin{aligned}\sigma_{\text{all}} &= \sigma_2 + \sigma_{>2} = \sigma_2 + \sigma_2 \left(\frac{1}{\Delta^2(t_0, t)} - 1 \right), \\ \Rightarrow \Delta^2(t_0, t) &= \frac{\sigma_2}{\sigma_{\text{all}}}.\end{aligned}$$

Two jet rate = $\Delta^2 = P^2$ (No emission in the range $t \rightarrow t_0$).

Sudakov form factor = No emission probability .

Often $\Delta(t_0, t) \equiv \Delta(t)$.

- ▶ Hard scale t , typically CM energy or p_{\perp} of hard process.
- ▶ Resolution t_0 , two partons are resolved as two entities if inv mass or relative p_{\perp} above t_0 .
- ▶ P^2 (not P), as we have two legs that evolve independently.