The Color Glass Condensate:

An effective theory of QCD at high energies

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Recent Reviews on the CGC:

L. McLerran, hep-ph/0311028

E. Iancu & R. Venugopalan, hep-ph/0303204

• A. H. Mueller, hep-ph/9911289

Road map of the strong interactions



Momentum Resolution Q^2

QCD evolution equations at small x

a) The DGLAP equation (Dokshitzer-Gribov-Lipatov-Altarelli-Parisi) If $x_B < 1$, gluon bremsstrahlung is dominant in QCD evolution: $\mathcal{P}_{qq} > \mathcal{P}_{qq} > \mathcal{P}_{qq}$ Q^2 Large Logs from Bremsstrahlung $^{\mathbf{k}_{\perp}\mathbf{r}_{2}} \stackrel{\frown}{\cap} \alpha_{S} \int \frac{dx}{x} \frac{d^{2}k_{\perp}}{k_{\perp}^{2}} \to \alpha_{S}^{p} \ln^{m}(1/x) \ln^{n}(Q^{2})$

For $Q^2>>\Lambda_{\rm QCD}^2$ and x~1, resum $\alpha_S\ln(Q^2)$ At small x, sum double logs

of gluons grows rapídly…



But... the phase space density decreases -the proton becomes more dilute Thus far, our discussion has focused on the Bjorken limit in QCD:

$$Q^2 \to \infty; s \to \infty; x_{\rm Bj} \approx \frac{Q^2}{s} = \text{fixed}$$

Asymptotic freedom, Factorization Theorems, machinery of precision physics in QCD...

Other interesting limit-is the Regge limit of QCD:

 $x_{\rm Bj} \to 0; s \to \infty; Q^2(>>\Lambda_{\rm QCD}^2) = \text{fixed}$

Physics of strong fields in QCD, multi-particle productionpossibly discover novel universal properties of theory in this limit





Phase space density grows rapidly-BFKL evolution breaks down when phase space density $f \sim 1$

Gluon density saturates at $f = \frac{1}{\alpha_S}$



Recombination effects compete with DGLAP Bremsstrahlung effects when

 $\alpha_S x G(x, Q^2) \sim R^2 Q^2$

Saturation of the gluon density for $Q \equiv Q_s(x)$

A hadron at high energies



Unlike QED, the QCD light cone vacuum is very complicated: -various topological objects, Instantons, Monopoles, Skyrmions, ... Hadrons may be bags or flux tubes or solitons: Complex phenomena - Chiral symmetry breaking, Confinement,...



Given this, how does one describe the structure of hadrons in high energy scattering?

How does one construct a Lorentz invariant wave fn for a hadron?

Partíal answer: formulate the theory on the líght cone



Quantize theory on light like surface: $x^+=0$

Quantum field theories quantized on light like surfaces have remarkable properties

Weinberg, 1966 Susskind, 1968





TWO DIMENSIONAL QUANTUM MECHANICS



Light cone pert. theory = Rayleigh-Schrodinger pert. theory

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- The MV-model
- Quantum evolution: a Wilsonian RG
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THE MV MODEL

McLerran, RV; Kovchegov Jalilian-Marian,Kovner,McLerran,Weigert

Consider large nucleus in the IMF frame: $P^+ \rightarrow \infty$





One large component of the current-others suppressed by $\frac{1}{P^+}$ Wee partons see a large density of valence color charges at small transverse resolutions.



<u>Born-Oppenheimer</u>: separation of large x and small x modes

Limiting fragmentation



Suggestive that valence partons are recoil-less sources-unaffected by Bremsstrahlung of wee partons



Gaussian random sources

$$\begin{aligned} & \textbf{THE EFFECTIVE ACTION} \\ & \textbf{Scale separating} \\ & \textbf{sources and fields} \\ & \mathcal{Z}[j] = \int [d\rho] \, W_{\Lambda^+}[\rho] \left\{ \frac{\int^{\Lambda^+} [dA] \, \delta(A^+) \, e^{iS[A,\rho] - \int j \cdot A}}{\int^{\Lambda^+} [dA] \, \delta(A^+) \, e^{iS[A,\rho]}} \right\} \end{aligned}$$

Gauge invariant weight functional describing distribution of the sources

$$\begin{split} S[A,\rho] &= \frac{-1}{4} \int d^4x \, F_{\mu\nu}^2 + \frac{i}{N_c} \int d^2x_{\perp} dx^- \delta(x^-) \text{Tr} \left(\rho(x_{\perp}) \, U_{-\infty,\infty}[A^-] \right) \\ \text{where } U_{-\infty,\infty}[A^-] &= \mathcal{P} \exp\left(ig \int dx^+ A^{-,a} T^a \right) \\ \text{To lowest order,} &= -J^+ A^- \text{ with } J^+ = g \, \rho(x_{\perp}) \, \delta(x^-) \end{split}$$

For a large nucleus,

$$W[\rho] = \exp\left(-\int d^2x_\perp \frac{\rho^a \rho^a}{2\,\mu_A^2}\right)$$
 where, for valence quark sources, one has $\mu_A^2 = \frac{g^2 A}{2\pi R_A^2} \propto A^{1/3}$ fm

-2

For A >>1,
$$\mu_A^2 >> \Lambda_{
m QCD}^2$$
 and $\alpha_S(\mu_A^2) << 1$

Effective action describes a weakly coupled albeit non-perturbative system

THE CLASSICAL FIELD OF THE NUCLEUS AT HIGH ENERGIES

Saddle point of effective action-> Yang-Mills equations

 $D_{\mu}F^{\mu\nu} = \delta^{\nu+}\delta(x^{-})\rho^{a}(x_{\perp})$

Solutions are <mark>non-Abelian</mark> Weizsäcker-Williams fields

 $\begin{aligned} A^+ &= A^- = 0 ;\\ F^{ij} &= 0 \Longrightarrow A^i = \theta(x^-)\alpha^i ,\\ \text{where } \alpha^i &= \frac{-1}{ig} U \nabla^i U^\dagger \\ \text{and } \nabla \cdot \alpha &= g\rho \end{aligned}$

Careful solution requires smearing in



Random Electric & Magnetic fields in the plane of the fast moving nucleus







✓ Gluons are colored

 Random sources evolving on time scales much larger than natural time scales-very similar to spin glasses
 Hadron/nucleus at high energies is a Color Glass Condensate The Color Glass Condensate: An effective field theory of QCD at high energies

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Constructive EFT: Wilson RG at small x



Color charge grows due to inclusion of fields into hard source with decreasing x: $\rho' = \rho + \delta \rho => W_x[\rho] \to W_{x'}[\rho']$



Because of strong fields $A\sim 1/g$ All insertions are O(1)

 $W_x[
ho]$ obeys a non-línear Wílson renormalízatíon group equation-the JIMWLK equation

(Jalilian-Marian, Iancu, McLerran, Weigert, Leonidov, Kovner)

At each step in the evolution, compute 1-point and 2-point functions in the background field

 $\sigma^a(x)[\rho] = \langle \delta \rho_Y^a(x) \rangle_{\rho} \; ; \; \chi^{ab}(x,y)[\rho] = \langle \delta \rho_Y^a(x) \delta \rho_Y^b(y) \rangle_{\rho}$



The JIMWLK (functional RG) equation:

$$\frac{\partial W_x[\rho]}{\partial \ln(1/x)} = \frac{1}{2} \int_{x_\perp, y_\perp} \frac{\delta}{\delta \rho_x^a(x_\perp)} \chi^{ab}(x_\perp, y_\perp)[\rho] \frac{\delta}{\delta \rho_x^b(y_\perp)} W_x[\rho]$$

⇒ An infinite hierarchy of ordinary differential equations for the correlators $\langle A_1 A_2 \cdots A_n \rangle_y$

Correlation Functions

Change of variables: $\rho^a \to \alpha^a\,;\, \nabla^2 \alpha = \rho$

$$< O[\alpha] >_{Y} = \int [d\alpha] O[\alpha] W_{Y}[\alpha]$$

Iancu, McLerran; Weigert

Brownian motion in functional space: Fokker-Planck equation!

"diffusion coefficient"

1

$$=>\frac{\partial}{\partial Y} < O[\alpha]>_Y = <\frac{1}{2}\int_{x,y}\frac{\delta}{\delta\alpha_Y^a(x)}\chi_{x,y}^{ab}\frac{\delta}{\delta\alpha_Y^b(y)}O[\alpha]>_Y$$
 "time"

Consider the 2-point function: $<lpha(x_{\perp})lpha(y_{\perp})>_{Y}$

Can solve JIMWLK in the weak field limit: $g\,lpha << 1$

Recover the BFKL equation in this low density limit





NOVEL REGIME OF QCD EVOLUTION AT HIGH ENERGIES



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THE BALITSKY-KOVCHEGOV EQUATION



Path ordered exponential $V^{\dagger}(x) = \mathcal{P} \exp\left(ig \int dx^{-} \alpha_{a}(x^{-}, x)T^{a}\right)$

> Weak field limit: $V^{\dagger}(x) \approx 1 + ig\alpha(x)$; $g\alpha << 1$

 $=> \mathcal{N}_{Y}(r) \sim \alpha_{s} r^{2} \frac{xG(x, 1/r^{2})}{\pi R^{2}} \quad \text{violates unitarity bound if } \mathcal{N} > 1$ $\blacktriangleright \text{ For } r > 1/Q_{s}(Y) \text{ dipole probes strong fields } (g\alpha \sim 1)$

Iancu-McLerran RPA => $< V^{\dagger}(x)V(y) >_{Y} << 1$ for $|x - y| >> 1/Q_{s}(Y)$ => N ~ 1 - dipole unitarizes



BK: Evolution eqn. for the dipole cross-section

• The 2-point correlator $\langle V^{\dagger}(x)V(y) \rangle$ in JIMWLK has

a closed form expression for $N_c \to \infty$ and A >> 1

$$\frac{\partial \mathcal{N}_{Y}(x,y)}{\partial Y} = \bar{\alpha}_{s} \int_{z} \frac{(x-y)^{2}}{(x-z)^{2}(y-z)^{2}} \left\{ \frac{\mathcal{N}_{Y}(x,z) + \mathcal{N}_{Y}(z,y) - \mathcal{N}_{Y}(x,y) - \mathcal{N}_{Y}(x,z)\mathcal{N}_{Y}(z,y)}{\mathsf{BFKL}} - \frac{\mathcal{N}_{Y}(x,z)\mathcal{N}_{Y}(z,y)}{\mathsf{Non-linear}} \right\}$$

- For small dipole, $(r << 1/Q_s(Y)) => BFKL$ eqn. $\mathcal{N}_Y(r) \approx (r^2 Q_0^2)^{1/2} e^{\omega \bar{\alpha}_s Y} \exp\left(-\frac{\ln^2(1/r^2 Q_0^2)}{2\beta \bar{\alpha}_s Y}\right)$
- From saturation condition,

 $\mathcal{N} = 1/2 \text{ when } r \sim 1/Q_s(Y) = > \ \ Q_s^2(Y) \approx Q_0^2 \, e^{\lambda Y} \text{ with } \lambda \sim 4.8 \, \alpha_s$

• For large dipole, $(r >> 1/Q_s(Y))$

$$\begin{split} \mathcal{N}_Y(r) &\approx 1 - \kappa \exp\left(-\frac{1}{4c}\ln^2(r^2Q_s^2(Y))\right) & \text{Levin,Tuchin;} \\ c &\approx 4.8 \end{split}$$

Numerical solutions of the BK-Eqn.



Figure 1: The functions $k\phi(k, Y)$ constructed from solutions to the BFKL and the Balitsky-Kovchegov equations for different values of the evolution parameter $Y = \ln(1/x)$ ranging from 1 to 10. The coupling constant $\alpha_s = 0.2$.

From K. Golec-Biernat, L. Motyka, A. M. Stasto, Phys Rev D65 (2002) 074037; hep-ph/0110325

No infrared diffusion a la BFKL in BK Exact analogy to travelling waves => Munier, Peschanski

Synopsis of CGC numerics

Numerical simulations of BK-eqn display Geometrical Scaling (Armesto, Braun; Golec-Biernat, Stasto, Motyka)

Infrared diffusion pathology of BFKL is cured.

State of the art: numerical simulations of JIMWLK n-point correlators by Rummukainen & Weigert

Running coupling effects important & still to be understood...

Iancu,Itakura, McLerran; Mueller,Triantafyllopolous

Can write the solution of BFKL as:

$$\begin{split} \mathcal{N}_{Y}(r_{\perp}) &\approx \exp\left(\omega\bar{\alpha}_{s}Y - \frac{\rho}{2} - \frac{\rho^{2}}{2\beta\bar{\alpha}_{s}Y}\right) \text{ with } \rho = \ln\frac{1}{r^{2}Q_{0}^{2}}\\ \rho_{S} \quad \text{ soln. where argument vanishes} \end{split}$$

$$= Q_s^2 = Q_0^2 e^{c\bar{\alpha}_s Y}$$
, with $c = 4.84$

For $r_\perp < 1/Q_s$ (but close!), can write $ho =
ho_S(Y) + \ln rac{1}{r_\perp^2 Q_s^2} \equiv
ho_S + \delta
ho$

Plugging into N_Y, can show simply

$$\mathcal{N}_Y \approx \left(r_\perp^2 Q_s^2(Y)\right)^\gamma \text{ for } Q_s^2 << Q^2 << \frac{Q_s^4}{Q_0^2}$$

 $\gamma \sim 0.64$ is large than BFKL anomalous dimension ~0.5

Geometrical scaling at HERA



How does Q_s behave as function of Y?

Fixed coupling LO BFKL: $Q_s^2 = Q_0^2 e^{c \bar{\alpha}_s Y}$ LO BFKL+ running coupling: $Q_s^2 = \Lambda_{QCD}^2 e^{\sqrt{2b_0 c(Y+Y_0)}}$ Re-summed NLO BFKL + CGC:



A-DEPENDENCE OF SATURATION SCALE



Such interesting systematics may be tested at LHC & eRHIC

Hadron & Nuclear Scattering at high energies

I: Universality: collinear versus k_t factorization



Hadronic collisions in the CGC framework



Solve Yang-Mills equations for two light cone sources: ho_{p1} & ho_{p2}

For observables $O(A_{\mu}(\rho_{p1}, \rho_{p2}))$ average over $W_{x1}[\rho_{p1}] \& W[\rho_{p2}]$

Gluon production in high energy pA collisions:

$$N_g = \frac{4g^2 N_c}{\pi^2 (N_c^2 - 1)q_{\perp}^2} \int \frac{d^3 q}{(2\pi)^3 2E_q} \frac{d^2 k_{\perp}}{(2\pi)^2} \int d^2 x_{\perp} \frac{d\phi_p(k_{\perp}, x_{\perp})}{d^2 X_{\perp}} \frac{d\phi_A(q_{\perp} - k_{\perp}, x_{\perp} - b)}{d^2 X_{\perp}}$$

$$\phi_A(k_\perp, x_\perp) \propto < U_{ab}^{\dagger} U_{bc} >_{\rho_A}$$

is non-linear in the gluon density to all orders-recover unintegrated gluon dist at large k_t

Quark production:

$$\begin{aligned} \frac{d\sigma^{pA \to q\bar{q}X}}{dy_p dy_A d^2 p_\perp q_\perp} &\propto \phi_p \times \left[A\phi_{g,g} + \left(B\phi_{g;q\bar{q}} + c.c\right) + C\phi_{q\bar{q};q\bar{q}}\right] \\ &< U_A(x_\perp) U_A^{\dagger}(y_\perp) > &< U_F(x_\perp) \tau^a U_F^{\dagger}(y_\perp) U_F(y'_\perp) \tau^b U_F(x'_\perp) > \\ &< U_F(x_\perp) \tau^a U_F^{\dagger}(y_\perp) \tau^{b'} (U_A^{a'b'})^{\dagger}(y'_\perp) > \end{aligned}$$

Not k_t factorizable!

Gluon & Quark production in the dense/AA region



Breaks k_t factorization? Contribution is same order in coupling

- Wave-fn evolution effects (beyond MV) difficult to include-work of Rummukainen & Weigert is promising.
- Classical evolution shows re-scattering-hence energy loss at eta=0 must be especially strong in AA!

COLLIDING SHEETS OF COLORED GLASS AT HIGH ENERGIES

$$\tau >> \frac{1}{Q_s}$$
 but $\tau << R$

Space-time evolution in heavy ion collisions



Initial conditions determined by saturation scale Q_s >> T_i

$$\varepsilon \propto \frac{Q_s^4}{\alpha_S}$$
 at $\tau \sim \frac{1}{Q_s}$

Q_s=1.4-2 GeV from HERA data extrapolations and numerical simulations

Do initial state (wave fn.) effects or final state (parton re-scattering) dominate in the space-time evolution?